



## RESEARCH LETTER

10.1002/2014GL059480

## Key Points:

- Representative data set of acetone in the UT/LMS on a global scale from CARIBIC
- HO<sub>x</sub> production due to acetone degradation and from O<sub>3</sub> photolysis are contrasted
- Year-around acetone is a significant source of HO<sub>x</sub> in the upper troposphere

## Correspondence to:

M. Neumaier,  
marco.neumaier@kit.edu

## Citation:

Neumaier, M., R. Ruhnke, O. Kirner, H. Ziereis, G. Stratmann, C. A. M. Brenninkmeijer, and A. Zahn (2014), Impact of acetone (photo)oxidation on HO<sub>x</sub> production in the UT/LMS based on CARIBIC passenger aircraft observations and EMAC simulations, *Geophys. Res. Lett.*, 41, 3289–3297, doi:10.1002/2014GL059480.

Received 3 FEB 2014

Accepted 28 MAR 2014

Accepted article online 1 APR 2014

Published online 1 MAY 2014

## Impact of acetone (photo)oxidation on HO<sub>x</sub> production in the UT/LMS based on CARIBIC passenger aircraft observations and EMAC simulations

M. Neumaier<sup>1</sup>, R. Ruhnke<sup>1</sup>, O. Kirner<sup>2</sup>, H. Ziereis<sup>3</sup>, G. Stratmann<sup>3</sup>, C. A. M. Brenninkmeijer<sup>4</sup>, and A. Zahn<sup>1</sup>

<sup>1</sup>Karlsruhe Institute of Technology (KIT), Institute of Meteorology and Climate Research (IMK-ASF), Karlsruhe, Germany, <sup>2</sup>Karlsruhe Institute of Technology (KIT), Steinbuch Centre for Computing (SCC), Karlsruhe, Germany, <sup>3</sup>Deutsches Zentrum für Luft- und Raumfahrt (DLR), Institute of Atmospheric Physics, Wessling, Germany, <sup>4</sup>Max Planck Institute for Chemistry (MPIC), Air Chemistry Division, Mainz, Germany

**Abstract** Until a decade ago, acetone was assumed to be a dominant HO<sub>x</sub> source in the dry extra-tropical upper troposphere (ex-UT). New photodissociation quantum yields of acetone and the lack of representative data from the ex-UT challenged that assumption. Regular mass spectrometric observations onboard the Civil Aircraft for the Regular Investigation of the atmosphere Based on an Instrument Container (CARIBIC) passenger aircraft deliver the first representative distribution of acetone in the UT/LMS (UT/lowermost stratosphere). Based on diverse CARIBIC trace gas data and non-observed parameters taken from the model ECHAM5/MESy for Atmospheric Chemistry, we quantify the HO<sub>x</sub> source in the UT/LMS from (photo)oxidation of acetone. The findings are contrasted to HO<sub>x</sub> production from ozone photolysis, overall the dominant tropospheric HO<sub>x</sub> source. It is shown that HO<sub>x</sub> production from acetone (photo)oxidation reaches up to 95% of the HO<sub>x</sub> source from ozone photolysis in autumn in the UT and on average ~61% in summer. That is, acetone is a significant source of HO<sub>x</sub> in the UT/LMS.

### 1. Introduction

HO<sub>x</sub> radicals (HO<sub>x</sub> = OH + HO<sub>2</sub>) govern the chemical lifetime of basically all organic species and many other crucial trace gases such as CH<sub>4</sub> and CO and thus largely control the oxidative capacity of the troposphere [e.g., Prinn, 2003]. The main primary source of OH near ground and in the mid-troposphere is reaction of O<sup>1</sup>D (formed *via* photolysis of ozone (O<sub>3</sub>)) with water (H<sub>2</sub>O), in the following denoted as O<sub>3</sub>-HO<sub>x</sub> source [Folkins and Chatfield, 2000; Jaeglé *et al.*, 2000; Monks, 2005]. This source becomes less dominant in the UT/lowermost stratosphere (LMS) due to the rapid decrease of H<sub>2</sub>O with altitude, whereas other HO<sub>x</sub> precursors like formaldehyde (CH<sub>2</sub>O), methyl hydroperoxide (CH<sub>3</sub>OOH), or acetone (CH<sub>3</sub>C(O)CH<sub>3</sub>) are gaining importance [Folkins and Chatfield, 2000].

Acetone photolysis was first suggested in a pioneering study by Singh *et al.* [1995] to be an important HO<sub>x</sub> source in the mid-latitude UT. This hypothesis was confirmed by subsequent studies [Arnold *et al.*, 1997; Jaeglé *et al.*, 1997; Wennberg *et al.*, 1998; Müller and Brasseur, 1999; Folkins and Chatfield, 2000; Jaeglé *et al.*, 2001]. However, the picture began to alter with the laboratory measurements of new temperature-dependent photolysis quantum yields (QYs) of acetone by Blitz *et al.* [2004]. As shown by Arnold *et al.* [2004, 2005], the reduced QYs result in acetone photolysis rates which are a factor of 3–10 smaller in the upper troposphere, compared to previously recommended photolysis rates based on QYs from Gierczak *et al.* [1998].

The reduced QYs obtained by Blitz *et al.* [2004] at long wavelengths and low temperatures, presently also recommended by expert panels [Atkinson *et al.*, 2006; Sander *et al.*, 2011], have recently been substantiated by Nádasdi *et al.* [2007] and Khamaganov and Crowley [2013], confirming that the importance of acetone as HO<sub>x</sub> precursor was indeed overestimated in the past.

A further crucial shortcoming in quantifying the role of acetone was the lack of representative data from the UT/LMS. This deficit can now largely be ameliorated by the CARIBIC (Civil Aircraft for the Regular Investigation of the atmosphere Based on an Instrument Container) project which delivers the largest in situ data set of acetone from the UT/LMS on a nearly world-wide scale [Sprung and Zahn, 2010].

The objective of this paper is to quantify the HO<sub>x</sub> source from (photo)oxidation (i.e., photolysis and reaction with OH) of acetone in the UT/LMS and compare it to the O<sub>3</sub>-HO<sub>x</sub> source. As a basis, we used representative tropopause-referenced distributions of acetone, O<sub>3</sub>, water vapor, and NO (which controls the yield of HO<sub>x</sub> from acetone degradation) as measured onboard CARIBIC. Parameters not accessible via the CARIBIC project, i.e., photolysis rates, the OH and HO<sub>2</sub> field, and the temperature field, were taken from the global chemistry model EMAC (ECHAM5/MESSy for Atmospheric Chemistry) [Jöckel *et al.*, 2006]. In a sensitivity study, we investigated the uncertainty of our findings, and it will be demonstrated that besides the photolysis rates of acetone and ozone, the acetone distribution is the most sensitive input parameter.

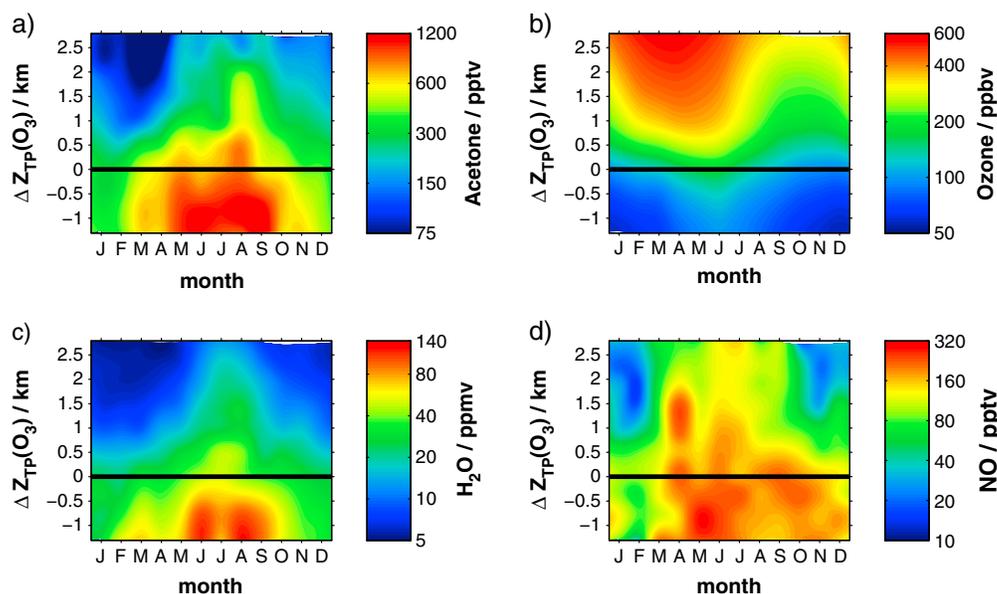
## 2. Experimental

In CARIBIC, a modified airfreight container equipped with 15 instruments for measurement of altogether ~100 trace gases and aerosol parameters is deployed onboard a passenger aircraft (Lufthansa, Airbus 340-600) during a sequence of two to six flights per month [Brenninkmeijer *et al.*, 2007]. The data analyzed here were collected at 9–12 km altitude (300–200 hPa), at 35°N–60°N and 124°W–142°E from 2005 to 2014. Among the most frequent flight destinations were Guangzhou (China), Caracas (Venezuela), Manila (Philippines), Chennai (India), and Vancouver (Canada) (overview available at [http://www.caribic.de/2005/Flight\\_Scheduling.html](http://www.caribic.de/2005/Flight_Scheduling.html)). The systematic and long-term nature of the CARIBIC data enables, e.g., the inference of seasonal variations. Moreover, the data cover a large part of the globe and altogether provide fairly representative distributions of trace species in the UT/LMS. For acetone, data from 185 flights were considered here, for O<sub>3</sub> and H<sub>2</sub>O ~318 and for NO 177 flights. The smaller number for acetone and NO is because of occasional technical problems. Onboard CARIBIC, acetone is measured by a proton-transfer-reaction mass spectrometer which is a considerably modified version of a commercial instrument from Ionicon (Innsbruck, Austria); see Brenninkmeijer *et al.* [2007] and Sprung and Zahn [2010]. Protonated acetone is detected at 59 amu. Isobaric propanal has a much smaller lifetime compared to acetone (~9 h vs. ~5 weeks), has a weaker source, and can thus be neglected in the UT/LMS [Warneke *et al.*, 2003]. The computed proton affinity of likewise isobaric glyoxal (C<sub>2</sub>H<sub>2</sub>O<sub>2</sub>) is too low to be effectively ionized by proton transfer from H<sub>3</sub>O<sup>+</sup> [Wróblewski *et al.*, 2007].

Ozone was measured by two instruments, an accurate UV-photometer and a fast chemiluminescence instrument [Zahn *et al.*, 2012]. Water vapor was measured by a photo-acoustic laser spectrometer and a commercial frost point hygrometer. The measurement of NO is based on the chemiluminescence of excited NO<sub>2</sub> molecules formed by reaction of NO with O<sub>3</sub> [Ziereis *et al.*, 2000; Stratmann, 2013]. NO was measured only during daylight.

Figure 1 shows the seasonal variation of acetone, O<sub>3</sub>, H<sub>2</sub>O vapor, and NO relative to the tropopause (TP) at mid-latitudes (35°N–60°N). The distance relative to the TP,  $\Delta Z_{TP}(O_3)$ , was derived from measured ozone and the well-documented gradient of O<sub>3</sub> in the LMS from O<sub>3</sub> sondes as described in Sprung and Zahn [2010].  $\Delta Z_{TP}(O_3)$  is an in situ measured and mixing-based altitude relative to the thermal TP and is far more accurate than the often used model-based tropopause height retrieved by, e.g., the ECMWF model. All four species (acetone, O<sub>3</sub>, H<sub>2</sub>O, NO) show strong vertical gradients in the UT/LMS and considerable seasonal variations. The most prominent features are the well-known descent of ozone-rich stratospheric air in spring within the downward branch of the Brewer-Dobson circulation and the considerable buildup of acetone and water vapor in the summertime UT and its propagation into the LMS. The CARIBIC acetone data set was recently contrasted to results from the LMDz-INCA global chemistry climate model in respect to spatial distribution, temporal variability, and representativeness [Elias *et al.*, 2011]. The summer maximum of acetone in the UT can qualitatively be explained by the (i) then peaking emission of precursor species and chemical turnover rates which together lead to maximum photochemical production of acetone, and (ii) by the enhanced lofting of polluted (acetone- and H<sub>2</sub>O-rich) near-surface air due to convection.

Mixing ratios of NO at the TP are higher between spring and autumn (~180 pptv) compared to winter (~100 pptv) because of (i) stronger photolysis of NO precursors (e.g., N<sub>2</sub>O and HNO<sub>3</sub>), (ii) stronger convective transport of polluted air into the UT, (iii) more frequent lightning events in the northern hemisphere (NH), and (iv) the seasonal cycle of NO<sub>y</sub> (NO<sub>x</sub> + HNO<sub>3</sub> + PAN + HONO + N<sub>2</sub>O<sub>5</sub> + NO<sub>3</sub> + ...) with enhanced levels in summer. The NO<sub>x</sub> winter minimum is not only due to the reduced solar actinic flux but is also associated with the heterogeneous conversion of reactive nitrogen to HNO<sub>3</sub> [Gao *et al.*, 1997; Brunner *et al.*, 2001]. NO above 1000 pptv (~0.6% of the data), mainly caused by fresh contrails and lightning, was not taken into account.



**Figure 1.** Seasonal variation of (a) acetone, (b) ozone, (c) H<sub>2</sub>O, and (d) NO relative to the thermal TP ( $\Delta Z_{TP}(O_3)=0$ , black line) for mid-latitudes (35°N–60°N). The data were collected with CARIBIC between May 2005 and January 2014.

### 3. EMAC Model

Not measured but necessary parameters to infer HO<sub>x</sub> production rates were taken from the Chemistry Climate Model EMAC [Jöckel *et al.*, 2006]. This included (a) photolysis rates of acetone, formaldehyde, methyl hydroperoxide, peracetic acid (CH<sub>3</sub>C(O)OOH), NO<sub>2</sub>, and O<sub>3</sub>, (b) the OH and HO<sub>2</sub> field, and (c) the temperature field.

EMAC is a combination of the general circulation model ECHAM5 [Roeckner *et al.*, 2006] with a selectable set of submodels such as the chemistry module MECCA1 [Sander *et al.*, 2005] combined through the Modular Earth Submodel System MESSy [Jöckel *et al.*, 2005]. For this study, we performed an EMAC (Version 1.10) simulation from 2008 to 2013 with the horizontal resolution T42 (2.8° × 2.8°) at 39 layers, covering the atmosphere from the surface up to ~80 km (0.01 hPa) and showing a vertical resolution of ~1.3 km at the TP. We applied the identical submodels as described by Kirner *et al.* [2011] and nudged the prognostic variables temperature, vorticity, divergence, and the surface pressure below 1 hPa towards ERA-Interim reanalysis [Dee *et al.*, 2011] to simulate realistic synoptic conditions.

The simulation included a comprehensive atmospheric chemistry setup with reaction rate coefficients following the latest JPL recommendations [Sander *et al.*, 2011]. For this study, photolysis rates  $J_i^\ddagger$ , averaged monthly and zonally between 35°N and 60°N were used. As mean daylight (symbolized by  $^\circ$ ) photolysis rates  $J_i^\circ$ , we applied as a first approximation

$$J_i^\circ = J_i^\ddagger \cdot f_s = J_i^\ddagger \cdot \frac{24}{(\text{hours of sunlight})_i} \quad (1)$$

where  $f_s$  accounts for the number of hours of sunlight in the respective month ranging from ~3.0 in December to ~1.5 in June and July. As HO<sub>x</sub> production from acetone (photo)oxidation takes place mainly during daylight, all other parameters (i.e., OH and HO<sub>2</sub> from EMAC) were also taken as daylight averaged monthly and zonally means.

The EMAC fields and CARIBIC data were interpolated and binned to the same grid with a vertical resolution of 0.45 km. Diurnally averaged OH and HO<sub>2</sub> distributions around the TP from EMAC (not shown) exhibit a strong seasonal variation with the highest concentrations in summer ([OH] ~ 1.1 · 10<sup>6</sup> cm<sup>-3</sup>, [HO<sub>2</sub>] ~ 2.5 · 10<sup>7</sup> cm<sup>-3</sup>) and significantly smaller concentration in winter ([OH] ~ 2 · 10<sup>5</sup> cm<sup>-3</sup>, [HO<sub>2</sub>] ~ 1 · 10<sup>7</sup> cm<sup>-3</sup>), in good agreement with Spivakovsky *et al.* [2000]. The corresponding daylight averaged values are higher by a factor of ~1.5 (June) to ~2.5 (December).

#### 4. Calculation of HO<sub>x</sub> Production Rates PHO<sub>x</sub>(O<sup>1</sup>D) and PHO<sub>x</sub>(Ac)

This section outlines the calculation of the HO<sub>x</sub> production rates from (a) O<sub>3</sub> photolysis and subsequent reaction of O<sup>1</sup>D with H<sub>2</sub>O denoted by PHO<sub>x</sub>(O<sup>1</sup>D) and from (b) photolysis of acetone and its reaction with OH denoted by PHO<sub>x</sub>(Ac). The most important input parameters—the distributions of the precursor gases O<sub>3</sub>, H<sub>2</sub>O, and acetone were taken from CARIBIC (Figures 1a–1c), in the following symbolized by an asterisk. Also, NO, necessary for the calculation of PHO<sub>x</sub>(Ac) (see section 4.2), was taken from CARIBIC (Figure 1d).

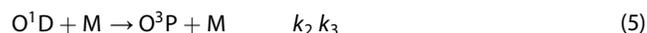
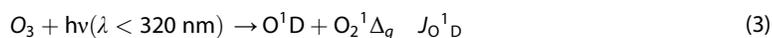
As NO<sub>2</sub> is measured by CARIBIC only during night, daylight NO<sub>2</sub> was calculated from the photostationary-state equation [Westberg *et al.*, 1971]

$$[\text{NO}_2] = \frac{k_{\text{NO}_2} \cdot [\text{NO}]^* \cdot [\text{O}_3]^*}{J_{\text{NO}_2}^\circ} \quad (2)$$

where  $k_{\text{NO}_2}$  is the rate constant for the  $\text{NO} + \text{O}_3 \rightarrow \text{NO}_2 + \text{O}_2$  reaction and  $J_{\text{NO}_2}^\circ$  is the mean daylight photolysis rate of NO<sub>2</sub>.

##### 4.1. HO<sub>x</sub> Production Rate From Ozone Photolysis

OH formation via photolysis of O<sub>3</sub> and subsequent reaction of O<sup>1</sup>D with H<sub>2</sub>O is given by reaction (3) – (5) [e.g. Levy, 1971; Logan *et al.*, 1981; Monks, 2005]



$J(\text{O}^1\text{D})$  is the photolysis rate of O<sub>3</sub> and  $k_1$  the rate constant for reaction (4) taken from Sander *et al.* [2011]. Reaction (5) is the electronic quenching of O<sup>1</sup>D due to collisions with N<sub>2</sub> ( $k_2$ ) or O<sub>2</sub> ( $k_3$ ) [Sander *et al.*, 2011]. The mean daylight HO<sub>x</sub> production rate PHO<sub>x</sub>(O<sup>1</sup>D) then is given by

$$\text{PHO}_x(\text{O}^1\text{D}) = 2 \cdot J(\text{O}^1\text{D})^\circ \cdot [\text{O}_3]^* \cdot \frac{k_1[\text{H}_2\text{O}]^*}{k_1[\text{H}_2\text{O}]^* + k_2[\text{N}_2] + k_3[\text{O}_2]} \quad (6)$$

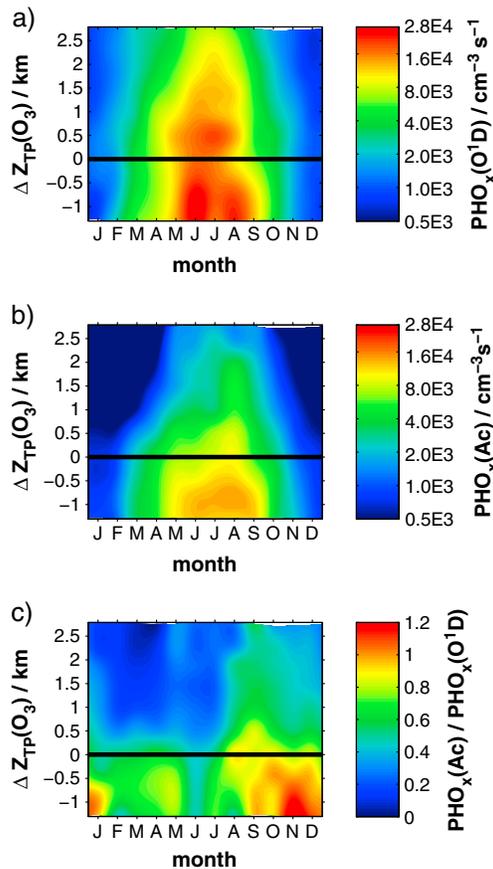
$J(\text{O}^1\text{D})^\circ$  maximizes with  $\sim 2.5 \cdot 10^{-5} \text{ s}^{-1}$  in summer, while the lowest values of  $\sim 5 \cdot 10^{-6} \text{ s}^{-1}$  are found in winter (not shown). The seasonal variation of PHO<sub>x</sub>(O<sup>1</sup>D) relative to the extra-tropical TP (Figure 2a) also maximizes in summer  $\sim 1 \text{ km}$  below the TP with  $\sim 2.7 \cdot 10^4 \text{ cm}^{-3} \text{ s}^{-1}$  because of maximizing water VMRs of  $\sim 120 \text{ ppmv}$  and the then peaking O<sub>3</sub> photolysis rate. Conversely, HO<sub>x</sub> production rates at the TP bottom out in winter with  $\sim 1 \cdot 10^3 \text{ cm}^{-3} \text{ s}^{-1}$ . The decrease of PHO<sub>x</sub>(O<sup>1</sup>D) with altitude is strongly correlated with the decrease of water. The summer maximum of PHO<sub>x</sub>(O<sup>1</sup>D) of  $\sim 2.7 \cdot 10^4 \text{ cm}^{-3} \text{ s}^{-1}$  is in excellent agreement with PHO<sub>x</sub>(O<sup>1</sup>D) values given by Colomb *et al.* [2006] for background air in the UT ( $\sim 10 \text{ km}$  altitude) over Europe in July of  $3 \cdot 10^{-3} \text{ ppt s}^{-1}$  ( $\sim 2.5 \cdot 10^4 \text{ cm}^{-3} \text{ s}^{-1}$ ).

##### 4.2. HO<sub>x</sub> Production Rate From Acetone Photolysis and Reaction With OH

To calculate HO<sub>x</sub> production rates from acetone photo(oxidation), we followed the analytical approach by Folkins and Chatfield [2000]. As the chemical sinks for acetone are (i) photolysis and (ii) oxidation by OH, the HO<sub>x</sub> production rate PHO<sub>x</sub>(Ac) is

$$\text{PHO}_x(\text{Ac}) = \text{HY}_{\text{hv}} \cdot J(\text{Ac})^\circ \cdot [\text{Ac}]^* + \text{HY}_{\text{OH}} \cdot k_{\text{OH}} \cdot [\text{OH}] \cdot [\text{Ac}]^* \quad (7)$$

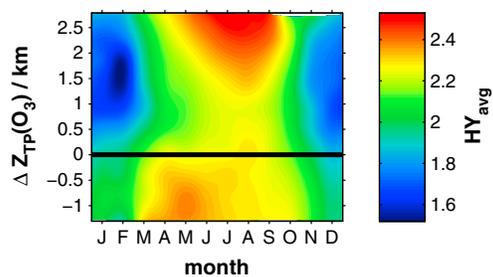
with  $k_{\text{OH}}$  the rate constant for the reaction of acetone with OH [Sander *et al.*, 2011], acetone [Ac]<sup>\*</sup> from CARIBIC (Figure 1a) and with  $\text{HY}_{\text{hv}}$  and  $\text{HY}_{\text{OH}}$  the net HO<sub>x</sub> yields per photolysed or oxidized acetone molecule, respectively. For the photolysis rate  $J(\text{Ac})^\circ$ , we used the latest JPL recommended values based on QYs by Blitz *et al.* [2004] and Sander *et al.* [2011] as daylight averaged monthly and zonally means.



**Figure 2.** (a) Seasonal variation of the daylight mean HO<sub>x</sub> production rate PHO<sub>x</sub>(O<sup>1</sup>D) from ozone photolysis and subsequent reaction of O<sup>1</sup>D with water vapor relative to the thermal TP (35°N–60°N). (b) Daylight mean seasonal variation of PHO<sub>x</sub>(Ac) calculated by using acetone data from CARIBIC. (c) Ratio PHO<sub>x</sub>(Ac)/PHO<sub>x</sub>(O<sup>1</sup>D).

~0.95 in September/October and ~1.2 in November. At higher altitudes, the impact of acetone becomes smaller because of the quickly decreasing acetone VMR, but it is still a noticeable HO<sub>x</sub> source with R<sub>(Ac/O<sup>1</sup>D)</sub> ranging from ~0.16 (MAM) to ~0.44 (SON) 2.5 km above the TP.

As shown by *Folkens and Chatfield* [2000], HY<sub>avg</sub> increases rapidly from 0.5 at NO<sub>x</sub> = 5 pptv to 2.6 at NO<sub>x</sub> = 100 pptv but increases only slightly above these levels to 2.7 at NO<sub>x</sub> = 1000 pptv.



**Figure 3.** Average net HO<sub>x</sub> yield (35°N–60°N) for acetone photolysis and reaction with OH based on the and HO<sub>x</sub> fields from EMAC and on NO from CARIBIC.

In accordance with *Folkens and Chatfield* [2000], by introducing a combined, average yield HY<sub>avg</sub>, which is the net number of formed HO<sub>x</sub> molecules per each photolysed and oxidized acetone molecule, equation (7) can be expressed as

$$PHO_x(Ac) = HY_{avg} \cdot [Ac]^* \cdot (J(Ac)^{\circ} + k_{OH} \cdot [OH]) \quad (8)$$

The yields HY<sub>avg</sub>, HY<sub>hv</sub>, and HY<sub>OH</sub> are determined by several reactions being part of the (photo)oxidation mechanism of acetone, i.e., reactions of acetone and its oxidation products with NO<sub>x</sub>, HO<sub>x</sub>, and several photolysis channels [*Folkens and Chatfield*, 2000]. HY<sub>avg</sub> depends on NO (taken from CARIBIC), NO<sub>2</sub> (from equation (2)), OH, and HO<sub>2</sub> (from EMAC) and on the reaction rate constants and photolysis rates of the particular reactions.

### 4.3. Results

As shown in Figure 2b, PHO<sub>x</sub>(Ac) shows a pronounced maximum up to ~14,000 cm<sup>-3</sup> s<sup>-1</sup> in summer ~1 km below the TP that can be traced back to (i) high acetone VMRs (Figure 1), (ii) a high photolysis rate, and (iii) a higher yield HY<sub>avg</sub> compared to winter (Figure 3). The average HO<sub>x</sub> production rate during summer (JJA) amounts to ~8500 cm<sup>-3</sup> s<sup>-1</sup> at the TP. *Colomb et al.* [2006] estimated HO<sub>x</sub> production rates from photolysis of acetone of 0.7 · 10<sup>-3</sup> ppt s<sup>-1</sup> (~5900 cm<sup>-3</sup> s<sup>-1</sup>) for background air in the UT over Europe in July which is in reasonable agreement with our findings.

Figure 2c shows the ratio R<sub>(Ac/O<sup>1</sup>D)</sub> = PHO<sub>x</sub>(Ac)/PHO<sub>x</sub>(O<sup>1</sup>D) that amounts at the TP to ~0.61 on average in summer and to ~0.83 in autumn with values up to

Figure 3 shows the seasonal variation of HY<sub>avg</sub> based on the HO<sub>x</sub> fields from EMAC and on the measured NO (and calculated NO<sub>2</sub>) from CARIBIC. HY<sub>avg</sub> maximizes in spring, the lowest values are found in winter above the TP. Photolysis of acetone and oxidation by OH both lead to net production of HO<sub>x</sub>. The contribution of the acetone + OH reaction to PHO<sub>x</sub>(Ac) amounts to ~25% at most at the TP in winter. Ratios of R<sub>(Ac/O<sup>1</sup>D)</sub> and R<sub>(Ac/O<sup>1</sup>D)</sub><sup>hv</sup> (neglecting OH reaction) together with values of HY<sub>avg</sub>, HY<sub>hv</sub>, and HY<sub>OH</sub> are summarized in Table 1. As shown by *Winkler et al.* [2002], the interaction of acetone with cirrus ice particles is too weak to result in a

**Table 1.** Ratios  $R_{(Ac/O1D)}$ ,  $R_{(Ac/O1D)}^{hv}$  (Photolysis only), Average  $HO_x$  Yield,  $HY_{avg}$  and  $HY_{hv}$  From Acetone Photolysis, and  $HY_{OH}$  From Oxidation With OH for Winter (DJF), Spring (MAM), Summer (JJA), and Autumn (SON) at the TP ( $\Delta Z_{TP}(O_3) = 0$  km) and 2.5 km Above the TP

$\Delta Z_{TP}(O_3)$		$R_{(Ac/O1D)}$	$R_{(Ac/O1D)}^{hv}$	$HY_{avg}$	$HY_{hv}$	$HY_{OH}$
0 km	DJF	0.61	0.48	1.93	2.84	0.90
	MAM	0.58	0.48	2.20	2.93	0.98
	JJA	0.61	0.51	2.27	2.95	1.00
	SON	0.83	0.67	2.12	2.95	1.00
2.5 km	DJF	0.22	0.19	1.81	2.56	0.66
	MAM	0.16	0.14	2.19	2.76	0.82
	JJA	0.27	0.24	2.45	2.92	0.96
	SON	0.44	0.38	2.18	2.77	0.87

significant partition to the ice phase so that the role of acetone as an  $HO_x$  precursor is not perturbed by the presence of cirrus clouds.

Figure 4 shows  $R_{(Ac/O1D)}$  as a function of  $H_2O$  from winter to autumn. Using photolysis rates based on the latest QYs acetone (photo)oxidation can still be considered a significant ( $R_{(Ac/O1D)} > 0.5$ )  $HO_x$  source at water VMRs below  $\sim 100$  ppmv (in summer/autumn) and reaches values up to  $\sim 1.2$  between 30 and 50 ppmv in autumn. Our analysis is in good agreement with Jaeglé *et al.* [2001] using Gierczak QYs-based photolysis rates, i.e., acetone degradation becomes a dominant ( $R_{(Ac/O1D)} > 1$ )  $HO_x$  source below  $\sim 120$  ppmv water vapor (not shown).

### 5. Sensitivity Study

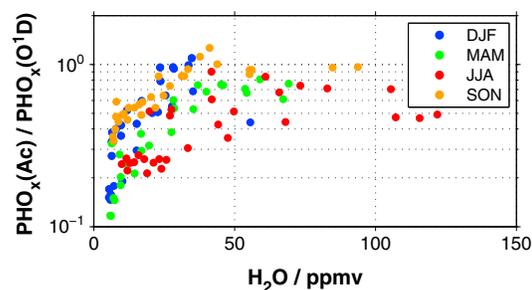
In a sensitivity study, we verified the uncertainty of our results by varying different input parameters.

#### 5.1. NO From CARIBIC

NO from CARIBIC ranges from  $111 \pm 100$  pptv (DJF) to  $188 \pm 106$  pptv (SON) at the TP and between  $30 \pm 42$  (DJF) and  $127 \pm 75$  pptv (JJA) 2.5 km above the TP. In the troposphere, NO is highly variable, and mixing ratios differ quite strongly dependent on the sampled region, e.g.,  $257 \pm 252$  pptv in spring over North America or  $276 \pm 414$  pptv in autumn over North Asia [Stratmann, 2013]. The lowest tropospheric mixing ratios during summer are observed over North America ( $99 \pm 104$  pptv) and during winter over the North Atlantic ( $54 \pm 49$  pptv) [Stratmann, 2013].

A comparison of CARIBIC data with data observed during the NOXAR (Nitrogen Oxides along Air Routes) project shows good agreement concerning the seasonal NO cycle over Europe/Asia and over the North Atlantic with maxima during summer. However, mixing ratios deviate partially strongly, which was to large part traced back to differences in flight routes [Brunner *et al.*, 2001; Stratmann, 2013].

When considering longitudinal-averaged mixing ratios, deviations between CARIBIC and NOXAR largely cancel out. For instance typical upper-tropospheric  $NO_x$  mixing ratios during NOXAR were  $100 \pm 50$  pptv (winter) and  $150 \pm 100$  pptv (summer) with up to  $\sim 230$  pptv at the TP in summer, while the highest levels were typically located 2–3 km below the TP [Grewe *et al.*, 2001].



**Figure 4.** Ratio  $PHO_x(Ac)/PHO_x(O^1D)$  versus water vapor from CARIBIC for winter (DJF), spring (MAM), summer (JJA), and autumn (SON).

As a test, we scaled the NO field from CARIBIC (and the derived  $NO_2$  field) by a factor of 0.5 (scenario 1) and 2 (scenario 2), respectively, while leaving the EMAC  $HO_x$  fields unchanged. The results (see Table 2) stay basically the same:  $R_{(Ac/O1D)}$  changes at most in winter 2.5 km above the TP by  $\sim -8\%$  (scenario 1) and by  $\sim 4\%$  (scenario 2). This is because  $NO_x$  in the NH UT is high enough so that  $HY_{avg}$  changes only slightly with varying  $NO_x$ . Therefore,  $NO_x$  is not a critical parameter.

**Table 2.** Ratio  $R_{(Ac/O1D)}$  and Changes of  $R_{(Ac/O1D)}$  (in %) Upon Scaling the Input Parameters by the Factors Given in Parentheses<sup>a</sup>

$\Delta Z_{TP}(O_3)$		$R_{(Ac/O1D)}$	1 NO (0.5)	2 NO (2)	3 OH, HO <sub>2</sub> (1.3,1.5)	4 $J(O^1D)$ (0.7,1.3)	5 $J(Ac)$ (0.8,1.2)
0 km	DJF	0.61	-4.5	2.3	2.2	43/-23	-16/16
	MAM	0.58	-4.1	2.1	0.6	43/-23	-17/17
	JJA	0.61	-4.1	2.1	-0.3	43/-23	-17,17
	SON	0.83	-3.4	1.7	1.5	43/-23	-16/16
2.5 km	DJF	0.22	-8.3	4.4	-2.5	43/-23	-17/17
	MAM	0.16	-3.8	1.9	-0.8	43/-23	-18/18
	JJA	0.27	-3.5	1.8	-1.0	43/-23	-18/18
	SON	0.44	-7.7	4.0	-3.0	43/-23	-18/18

<sup>a</sup>Scenarios 1–5 (see text).

## 5.2. EMAC HO<sub>x</sub> Field

Recently, *Regelin et al.* [2013] have shown that calculated HO<sub>x</sub> levels from EMAC in the upper-troposphere (7–10 km) over Europe significantly underestimate observed levels by ~24% (OH) and ~41% (HO<sub>2</sub>). Upon scaling the OH/HO<sub>2</sub> field by factors of 1.3/1.5 (scenario **3**), while leaving all other parameters unchanged (Table 2),  $R_{(Ac/O1D)}$  decreases at most by ~-3% 2.5 km above the TP in autumn.

## 5.3. Photolysis Rates

For the photolysis cross sections and quantum yields of ozone, a combined uncertainty of 1.3 is given by *Sander et al.* [2011] which will be used here as the uncertainty of the O<sub>3</sub> photolysis rate. Scaling  $J(O^1D)$  by factors of 0.7 and 1.3, respectively, leads to changes of ~43% and ~-23% of  $R_{(Ac/O1D)}$  at the TP (scenario **4**). For the photolysis rate of acetone, we assume an uncertainty factor of 1.2, based on the overall error of 10–15% for the parameterization of the QYs given by *Blitz et al.* [2004] which directly converts into a  $R_{(Ac/O1D)}$  change of ~±17% at the TP (scenario **5**).

All other used photolysis rates affect the result only indirectly by influencing the yield  $HY_{avg}$ . For instance, scaling  $J(NO_2)$  by its combined uncertainty of 1.2 leads only to small changes of  $R_{(Ac/O1D)}$  by 1–3%. The impact on  $R_{(Ac/O1D)}$  upon scaling  $J(CH_2O)$  by its uncertainty factor of 1.4 [*Sander et al.*, 2011] is even less pronounced with small increases up to ~2%.

In summary, we estimate the total uncertainty of the acetone HO<sub>x</sub> production rate by using Gaussian error propagation to ~35% (30% estimated uncertainty being the sum of instrumental uncertainty and the limited representativeness of the data and 20% uncertainty of the acetone photolysis rates).

## 6. Summary

The HO<sub>x</sub> source from acetone degradation  $PHO_x(Ac)$  due to photolysis and reaction with OH was quantified and compared with the HO<sub>x</sub> source from ozone photolysis  $PHO_x(O^1D)$  in the mid-latitude UT/LMS. Representative distributions of acetone, ozone, water vapor, and NO collected onboard the CARIBIC passenger aircraft between 2005 and 2014 at 9–12 km at mid-latitudes were considered. Not detected parameters like photolysis rates and HO<sub>x</sub> distributions were taken from the global chemistry model EMAC.

Year-around, acetone was found to be an important precursor of HO<sub>x</sub> around the mid-latitude TP. Typical summer values of  $PHO_x(Ac)$  amount to ~8500 cm<sup>-3</sup>·s<sup>-1</sup> with a maximum of 14,000 cm<sup>-3</sup>·s<sup>-1</sup> ~1 km below the TP which is on average ~60% of the HO<sub>x</sub> production rate from ozone photolysis. The strongest impact by acetone is observed in autumn when  $PHO_x(Ac)$  is up to ~95% of  $PHO_x(O^1D)$ , i.e., when the UT is already quite dry with H<sub>2</sub>O ≈ 30 – 50 ppmv but still rich in acetone with on average ~500 pptv. Above the TP, the impact from acetone decreases due to the quickly decreasing acetone levels and reaches at 2.5 km above the TP ~44% in autumn.

In a sensitivity study, the uncertainty of our findings was quantified, and the parameters that impact them most were identified. It was demonstrated that besides the photolysis rates of acetone the acetone

distribution is the most sensitive input parameter. This is a crucial point, as in the past no representative distributions of acetone in the UT/LMS have been available. We studied the influence of (i) the CARIBIC NO distribution which affects the yield of formed HO<sub>x</sub> species during the degradation of acetone, (ii) of the OH and HO<sub>2</sub> fields from EMAC, and (iii) of several photolysis rates. Scaling the CARIBIC NO field by factors of 0.5 and 2, which results in unrealistically low (high) NO levels, changes R<sub>(Ac/O1D)</sub> by −5–2% at the TP. The dependence on the OH/HO<sub>2</sub> is comparable with R<sub>(Ac/O1D)</sub> changing by ~2% in winter upon scaling the OH and HO<sub>2</sub> field by a factor of 1.3 and 1.5. Altogether, the sensitivity study in section 5 indicates that the uncertainty of the inferred acetone-derived HO<sub>x</sub> production rates in the UT/LMS is ~35%.

#### Acknowledgments

We are deeply indebted to Andreas Waibel, Lufthansa for successfully guiding the scientific project CARIBIC. Likewise, we thank the operational flight planners, Mrs. Sieglinde Bacher and Mr. Peter Hartleib. CARIBIC obtains as part of IAGOS ([www.iagos.org](http://www.iagos.org)) funding from the Germany Ministry of Science and Education. We also acknowledge financial help from Frankfurt Airport, Fraport AG, and the Max Planck Society. Further, the authors would like to thank John Crowley (MPIC, Mainz) for discussions.

The Editor thanks two anonymous reviewers for their assistance in evaluating this paper.

#### References

- Arnold, S. R., M. P. Chipperfield, M. A. Blitz, D. E. Heard, and M. J. Pilling (2004), Photodissociation of acetone: Atmospheric implications of temperature-dependent quantum yields, *Geophys. Res. Lett.*, *31*, L07110, doi:10.1029/2003GL019099.
- Arnold, F., V. Bürger, B. Droste-Fanke, F. Grimm, A. Krieger, J. Schneider, and T. Stilp (1997), Acetone in the upper troposphere and lower stratosphere: Impact on trace gases and aerosols, *Geophys. Res. Lett.*, *24*(23), 3017–3020, doi:10.1029/97GL02974.
- Arnold, S. R., M. P. Chipperfield, and M. A. Blitz (2005), A three-dimensional model study of the effect of new temperature-dependent quantum yields for acetone photolysis, *J. Geophys. Res.*, *110*, 22305, doi:10.1029/2005JD005998.
- Atkinson, R., D. L. Baulch, R. A. Cox, J. N. Crowley, R. F. Hampson, R. G. Hynes, M. E. Jenkin, M. J. Rossi, and J. Troe (2006), Evaluated kinetic and photochemical data for atmospheric chemistry: Volume II – gas phase reactions of organic species, *Atmos. Chem. Phys.*, *6*(11), 3625–4055, doi:10.5194/acp-6-3625-2006.
- Blitz, M. A., D. E. Heard, M. J. Pilling, S. R. Arnold, and M. P. Chipperfield (2004), Pressure and temperature-dependent quantum yields for the photodissociation of acetone between 279 and 327.5 nm, *Geophys. Res. Lett.*, *31*, L06111, doi:10.1029/2003GL018793.
- Brenninkmeijer, C. A. M., et al. (2007), Civil Aircraft for the regular investigation of the atmosphere based on an instrumented container: The new CARIBIC system, *Atmos. Chem. Phys.*, *7*(18), 4953–4976, doi:10.5194/acp-7-4953-2007.
- Brunner, D., J. Staehelin, D. Jeker, H. Wernli, and U. Schumann (2001), Nitrogen oxides and ozone in the tropopause region of the northern hemisphere: Measurements from commercial aircraft in 1995/1996 and 1997, *J. Geophys. Res.*, *106*(D21), 27,673–27,699, doi:10.1029/2001JD900239.
- Colomb, A., et al. (2006), Airborne measurements of trace organic species in the upper troposphere over Europe: The impact of deep convection, *Environ. Chem.*, *3*(4), 244–259, doi:10.1071/EN06020.
- Dee, D. P., et al. (2011), The ERA-Interim reanalysis: Configuration and performance of the data assimilation system, *Q. J. R. Meteorol. Soc.*, *137*(656), 553–597, doi:10.1002/qj.828.
- Elias, T., S. Szopa, A. Zahn, T. Schuck, C. Brenninkmeijer, D. Sprung, and F. Slemr (2011), Acetone variability in the upper troposphere: Analysis of CARIBIC observations and LMDz-INCA chemistry-climate model simulations, *Atmos. Chem. Phys.*, *11*(15), 8053–8074, doi:10.5194/acp-11-8053-2011.
- Folkens, I., and R. Chatfield (2000), Impact of acetone on ozone production and OH in the upper troposphere at high NO<sub>x</sub>, *J. Geophys. Res.*, *105*(D9), 11,585–11,599, doi:10.1029/2000JD900067.
- Gao, R. S., et al. (1997), Partitioning of the reactive nitrogen reservoir in the lower stratosphere of the southern hemisphere: Observations and modeling, *J. Geophys. Res.*, *102*(D3), 3935–3949, doi:10.1029/96JD01967.
- Gierczak, T., J. B. Burkholder, S. Bauerle, and A. R. Ravishankara (1998), Photochemistry of acetone under tropospheric conditions, *Chem. Phys.*, *231*(2–3), 229–244, doi:10.1016/S0301-0104(98)00006-8.
- Grewe, V., D. Brunner, M. Dameris, J. L. Grenfell, R. Hein, D. Shindell, and J. Staehelin (2001), Origin and variability of upper tropospheric nitrogen oxides and ozone at northern mid-latitudes, *Atmos. Environ.*, *35*(20), 3421–3433, doi:10.1016/S1352-2310(01)00134-0.
- Jaeglé, L., D. J. Jacob, W. H. Brune, and P. O. Wennberg (2001), Chemistry of HO<sub>x</sub> radicals in the upper troposphere, *Atmos. Environ.*, *35*(3), 469–489, doi:10.1016/S1352-2310(00)00376-9.
- Jaeglé, L., et al. (1997), Observed OH and HO<sub>2</sub> in the upper troposphere suggest a major source from convective injection of peroxides, *Geophys. Res. Lett.*, *24*(24), 3181–3184, doi:10.1029/97GL03004.
- Jaeglé, L., et al. (2000), Photochemistry of HO<sub>x</sub> in the upper troposphere at northern midlatitudes, *J. Geophys. Res.*, *105*(D3), 3877–3892, doi:10.1029/1999JD901016.
- Jöckel, P., R. Sander, A. Kerkweg, H. Tost, and J. Lelieveld (2005), Technical Note: The Modular Earth Submodel System (MESSy) - a new approach towards Earth System Modeling, *Atmos. Chem. Phys.*, *5*(2), 433–444, doi:10.5194/acp-5-433-2005.
- Jöckel, P., et al. (2006), The atmospheric chemistry general circulation model ECHAM5/MESSy1: Consistent simulation of ozone from the surface to the mesosphere, *Atmos. Chem. Phys. Discuss.*, *6*(4), 6957–7050, doi:10.5194/acpd-6-6957-2006.
- Khamaganov, V. G., and J. N. Crowley (2013), Pressure dependent photolysis quantum yields for CH<sub>3</sub>C(O)CH<sub>3</sub> at 300 and 308 nm and at 298 and 228 K, *Phys. Chem. Chem. Phys.*, *15*(25), 10,500–10,509, doi:10.1039/c3cp50291k.
- Kirner, O., R. Ruhnke, J. Buchholz-Dietsch, P. Jöckel, C. Brühl, and B. Steil (2011), Simulation of polar stratospheric clouds in the chemistry-climate-model EMAC via the submodel PSC, *Geosci. Model Dev.*, *4*(1), 169–182, doi:10.5194/gmd-4-169-2011.
- Levy, H., II (1971), Normal atmosphere: Large radical and formaldehyde concentrations predicted, *Science*, *173*(3992), 141–143, doi:10.1126/science.173.3992.141.
- Logan, J. A., M. J. Prather, S. C. Wofsy, and M. B. McElroy (1981), Tropospheric chemistry: A global perspective, *J. Geophys. Res.*, *86*(C8), 7210–7254, doi:10.1029/JC086iC08p07210.
- Monks, P. S. (2005), Gas-phase radical chemistry in the troposphere, *Chem. Soc. Rev.*, *34*(5), 376–395, doi:10.1039/b307982c.
- Müller, J.-F., and G. Brasseur (1999), Sources of upper tropospheric HO<sub>x</sub>: A three-dimensional study, *J. Geophys. Res.*, *104*(D1), 1705–1715, doi:10.1029/1998JD100005.
- Nádasdi, R., G. Kovács, I. Szilágyi, A. Demeter, S. Dóbbé, T. Bérces, and F. Márta (2007), Exciplex laser photolysis study of acetone with relevance to tropospheric chemistry, *Chem. Phys. Lett.*, *440*(1–3), 31–35, doi:10.1016/j.cplett.2007.04.014.
- Prinn, R. G. (2003), The cleansing capacity of the atmosphere, *Annu. Rev. Environ. Resour.*, *28*(1), 29–57, doi:10.1146/annurev.energy.28.011503.163425.
- Regelin, E., et al. (2013), HO<sub>x</sub> measurements in the summertime upper troposphere over Europe: A comparison of observations to a box model and a 3-D model, *Atmos. Chem. Phys.*, *13*(21), 10,703–10,720, doi:10.5194/acp-13-10703-2013.

- Roeckner, E., R. Brokopf, M. Esch, M. Giorgetta, S. Hagemann, L. Kornblueh, E. Manzini, U. Schlese, and U. Schulzweida (2006), Sensitivity of simulated climate to horizontal and vertical resolution in the ECHAM5 atmosphere model, *J. Clim.*, *19*(16), 3771–3791, doi:10.1175/JCLI3824.1.
- Sander, R., A. Kerkweg, P. Jöckel, and J. Lelieveld (2005), Technical note: The new comprehensive atmospheric chemistry module MECCA, *Atmos. Chem. Phys.*, *5*(2), 445–450, doi:10.5194/acp-5-445-2005.
- Sander, S. P., et al. (2011), Chemical kinetics and photochemical data for use in atmospheric studies, Evaluation Number 17, JPL Publication 10-6(17), Jet Propulsion Laboratory, Pasadena, Calif. [Available at <http://jpldataeval.jpl.nasa.gov>.]
- Singh, H. B., M. Kanakidou, P. J. Crutzen, and D. J. Jacob (1995), High concentrations and photochemical fate of oxygenated hydrocarbons in the global troposphere, *Nature*, *378*(6552), 50–54, doi:10.1038/378050a0.
- Spivakovsky, C. M., et al. (2000), Three-dimensional climatological distribution of tropospheric OH: Update and evaluation, *J. Geophys. Res.*, *105*(D7), 8931–8980, doi:10.1029/1999JD901006.
- Sprung, D., and A. Zahn (2010), Acetone in the upper troposphere/lowermost stratosphere measured by the CARIBIC passenger aircraft: Distribution, seasonal cycle, and variability, *J. Geophys. Res.*, *115*, D16301, doi:10.1029/2009JD012099.
- Stratmann, G. (2013), Stickoxidmessungen in der Tropopausenregion an Bord eines Linienflugzeugs: Großräumige Verteilung und Einfluss des Luftverkehrs, Technische Univ. München, Munich, Germany.
- Warneke, C., J. A. de Gouw, W. C. Kuster, P. D. Goldan, and R. Fall (2003), Validation of atmospheric VOC measurements by proton-transfer-reaction mass spectrometry using a gas-chromatographic pre-separation method, *Environ. Sci. Technol.*, *37*(11), 2494–2501, doi:10.1021/es026266i.
- Wennberg, P. O., et al. (1998), Hydrogen radicals, nitrogen radicals, and the production of O<sub>3</sub> in the upper troposphere, *Science*, *279*(5347), 49–53, doi:10.1126/science.279.5347.49.
- Westberg, K., N. Cohen, and K. W. Wilson (1971), Carbon monoxide: Its role in photochemical smog formation, *Science*, *171*(3975), 1013–1015, doi:10.1126/science.171.3975.1013.
- Winkler, A. K., N. S. Holmes, and J. N. Crowley (2002), Interaction of methanol, acetone and formaldehyde with ice surfaces between 198 and 223 K, *Phys. Chem. Chem. Phys.*, *4*(21), 5270–5275, doi:10.1039/b206258e.
- Wróblewski, T., L. Ziemczonek, A. M. Alhasan, and G. P. Karwasz (2007), Ab initio and density functional theory calculations of proton affinities for volatile organic compounds, *Eur. Phys. J. Spec. Top.*, *144*(1), 191–195, doi:10.1140/epjst/e2007-00126-7.
- Zahn, A., J. Weppner, H. Widmann, K. Schlote-Holubek, B. Burger, T. Kühner, and H. Franke (2012), A fast and precise chemiluminescence ozone detector for eddy flux and airborne application, *Atmos. Meas. Tech.*, *5*(2), 363–375, doi:10.5194/amt-5-363-2012.
- Ziereis, H., H. Schlager, P. Schulte, P. F. J. van Velthoven, and F. Slemr (2000), Distributions of NO, NO<sub>x</sub>, and NO<sub>y</sub> in the upper troposphere and lower stratosphere between 28° and 61°N during POLINAT 2, *J. Geophys. Res.*, *105*(D3), 3653–3664, doi:10.1029/1999JD900870.