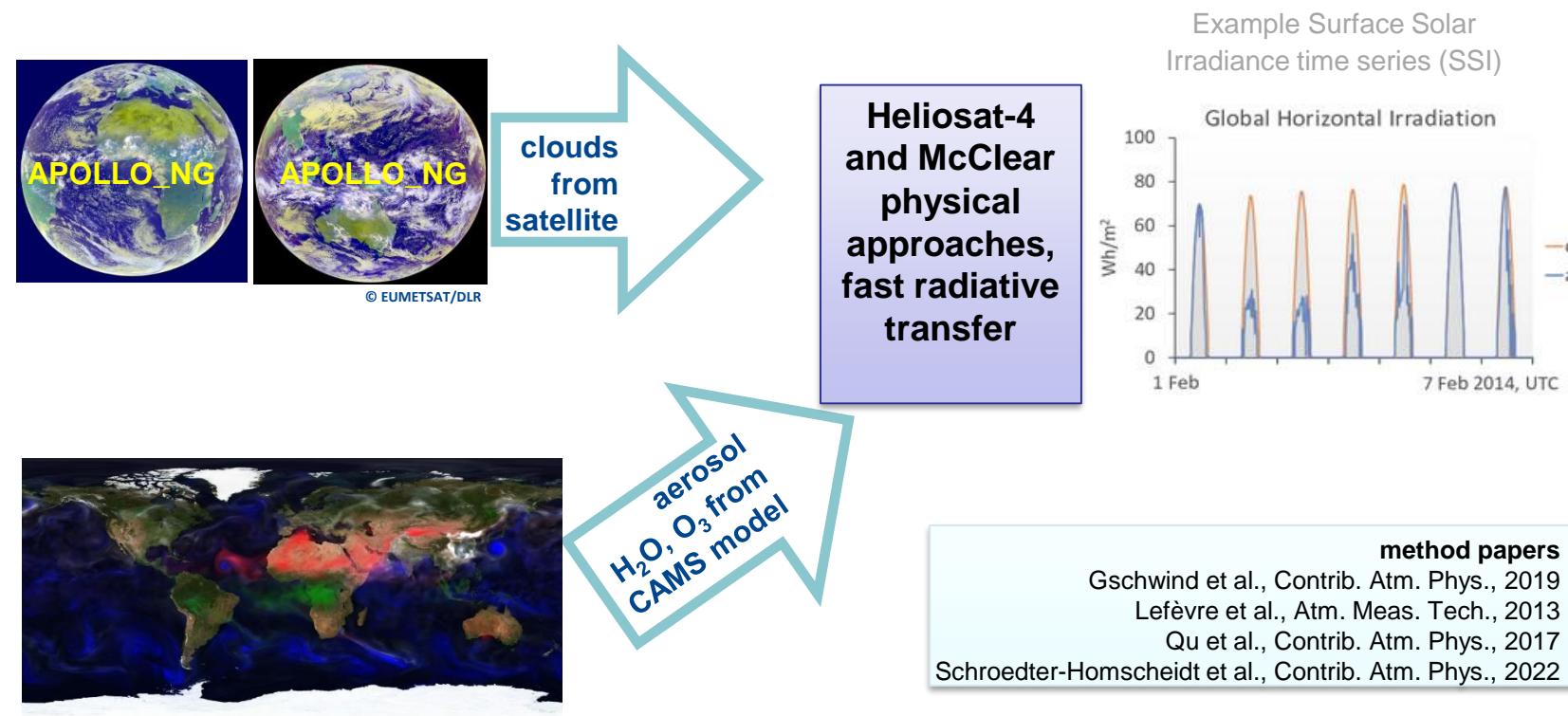


CAMS radiation service for solar energy: Exploring the error space with data-driven and spatially resolved methods and service evolution

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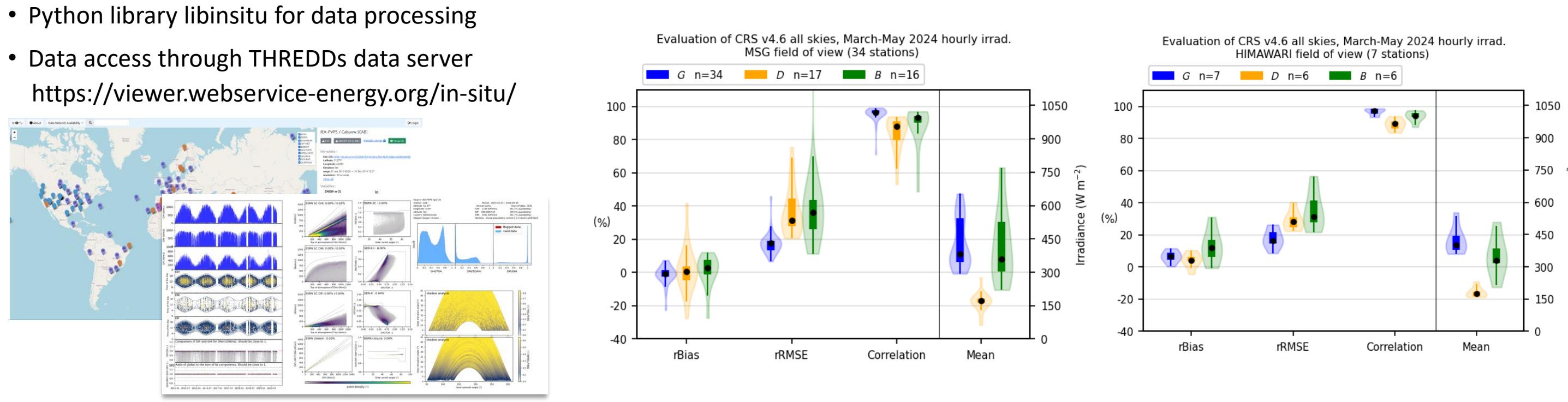
Copernicus Atmospheric Monitoring service (CAMS) solar radiation services (CRS)



- Primary product: SSI Time Series**
 - global, diffuse, direct and direct normal irradiance
 - 2004 onwards in MSG & 2016 onwards for HIMAWARI field of view (FOV)
 - 1 min, 15 min, 1 hour, 1 day, 1 month resolution
 - interactive access on CAMS ADS and SODA portal
 - OGC script access possible or via open source library Pvlib
 - access to all input data in expert mode (1 min)
- Pre-calculated SSI gridded data**
 - global, diffuse, direct and direct normal irradiance
 - 15 min temporal resolution selected
 - 2005-2023 in MSG & 2016-2023 in HIMAWARI FOV
 - interactive access on CAMS ADS
 - 0.1° spatial grid selected

CRS routine evaluations and Quality control (EQC)

- Provision of ground based SSI measurements collected on global scale to support CRS evaluations
- Consolidated metadata; harmonized netCDF structure following Climate and Forecast (CF) conventions and the FAIR (Findable, Accessible, Interoperable, Reproducible) principle
- Automated visual Quality Control (QC)
- Python library libinsitu for data processing
- Data access through THREDDS data server <https://viewer.webservice-energy.org/in-situ/>
- Publicly available <https://atmosphere.copernicus.eu/supplementary-services>



CAMEO: the detailed assessment of the error space of CRS

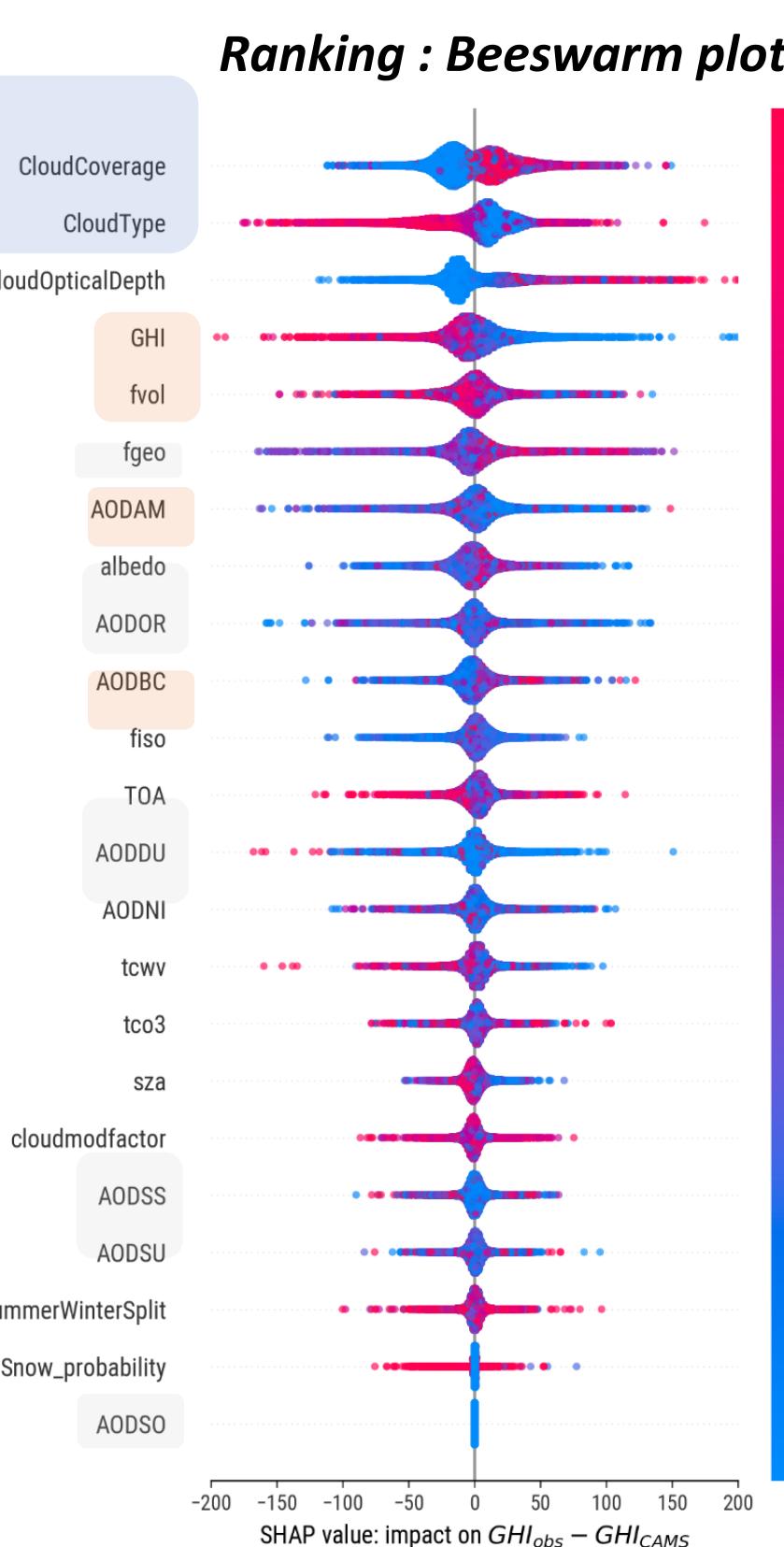
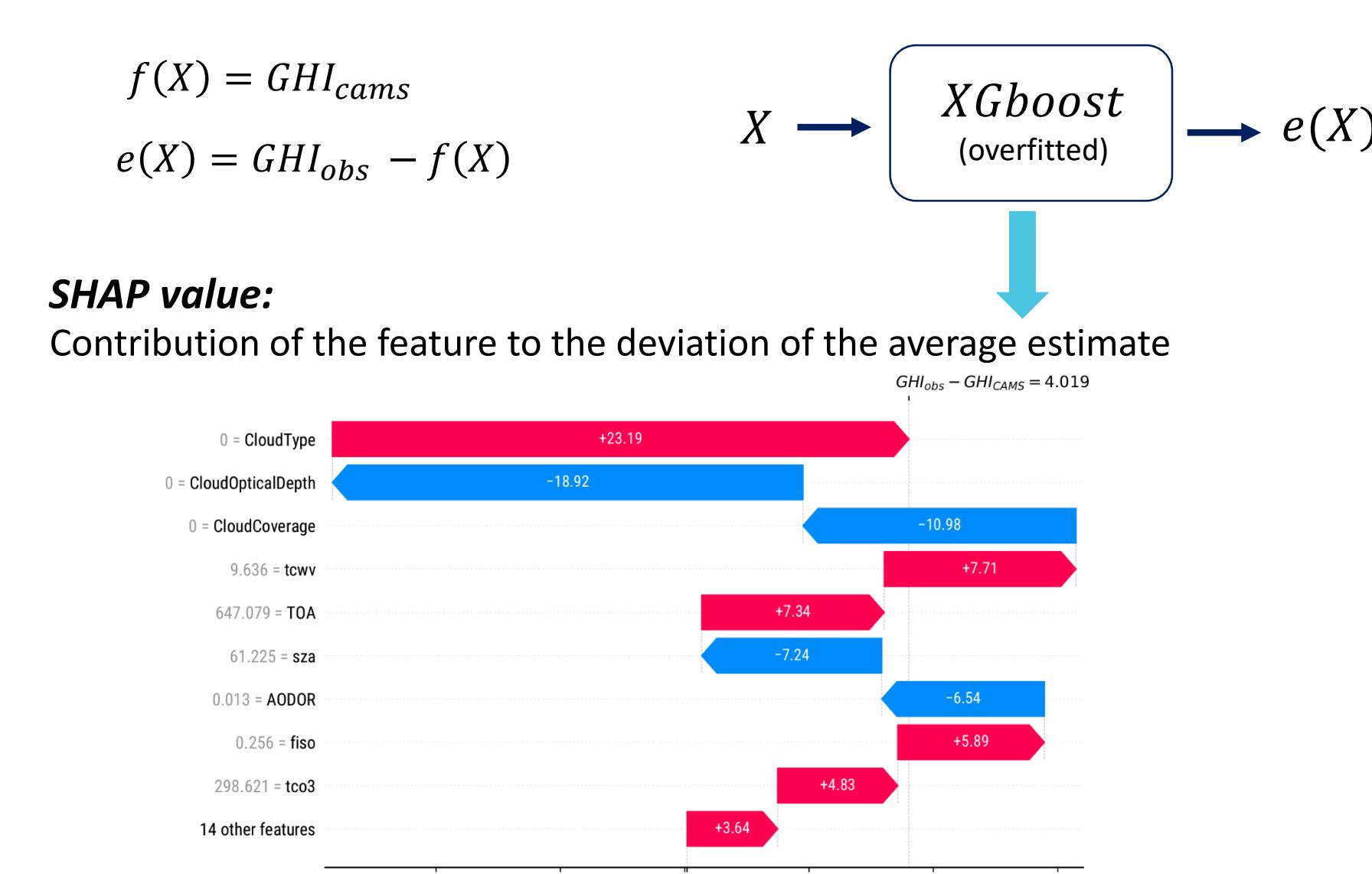
Questions to be addressed in CAMEO

- Are the aggregated error metrics provided to the users sufficient for expert applications?
- Can we assess the main error sources of the CRS irradiance estimates?
- Could we give more pixel-wise (non aggregated) uncertainty indicators for irradiance estimations to the users?

SHAP analysis as a rational for the selection of uncertainty model input parameters

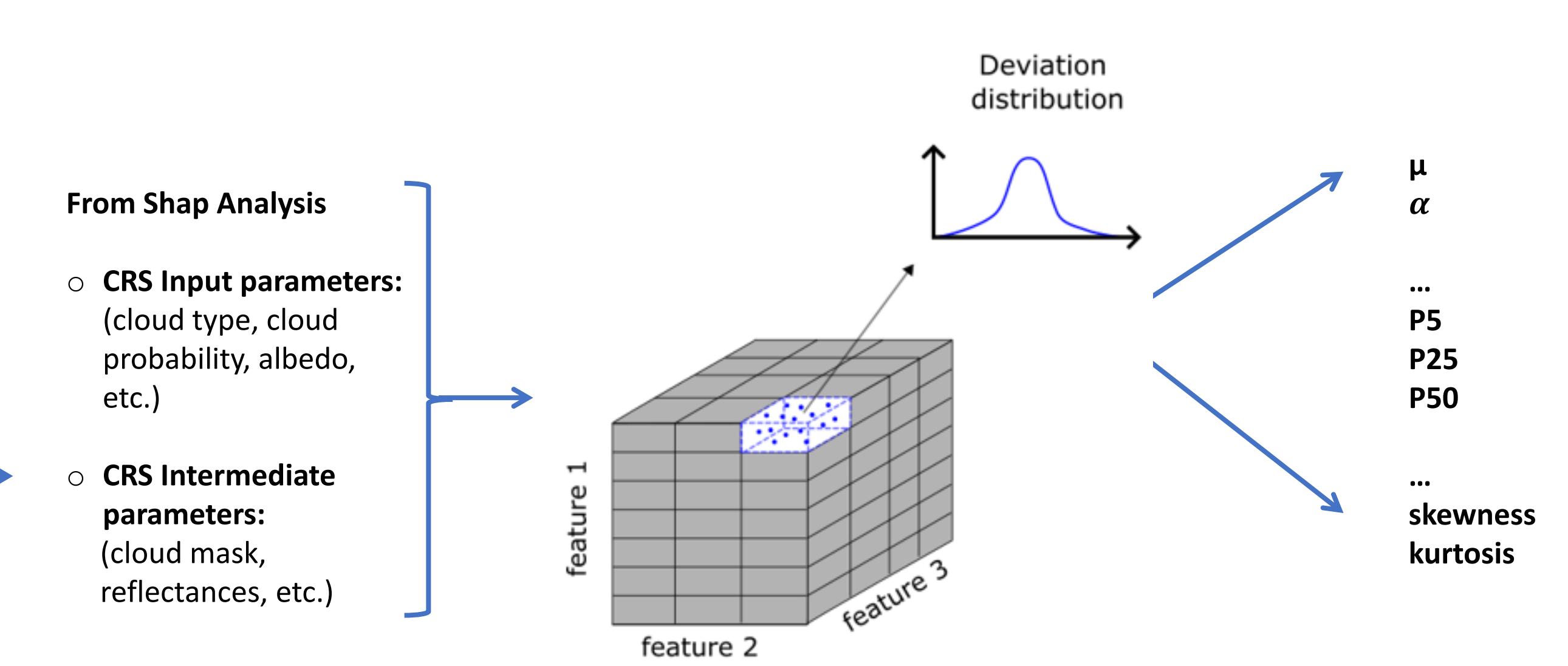
SHapley Additive exPlanations (SHAP) analysis:

Method (game theory) to understand the contributions of single model parameters/features on the total error of the CRS irradiance estimates



Pixel-wise error model development

Principle of the look up table (LUT) and neural network (NN) approaches:

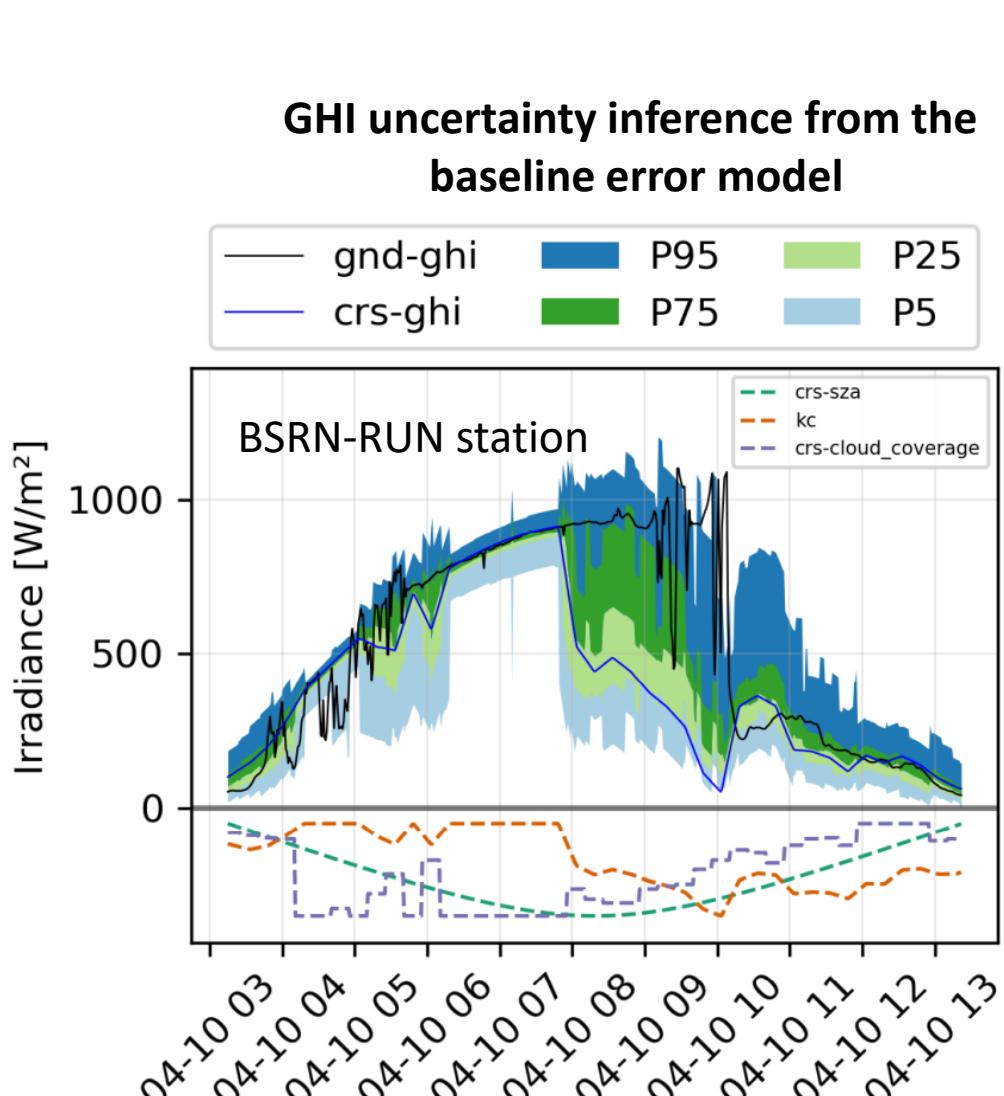


Methodology :

- Training :** find a probability distribution for every bin in our multidimensional feature space (56 internal cams parameters)
- Inference :** for the specific location, estimate the uncertainty from the probability distribution that corresponds to its specific feature combination (time dependent)
- Approaches :** LUT approach / NN approach optimizing moments of a parametric distribution (MBE (μ), STDE (α), kurtosis, skewness)

Localized error model 1: uncertainty inference based on LUT

- LUTs created using three parameters:
 - cloud probability: (from 0% to 100% every 5%) -> (20 ranges)
 - $K_c = GHI/GHI_c$: (from 0 to 1.4 every 0.1) -> (14 ranges)
 - CRS-solar zenith angle (crs-sza): (from 0° to 80° every 2°) -> (40 ranges)
- Percentiles used in order to achieve a probabilistic-based inference of the CRS errors (from P0 to P100 every 1 percentile) + (MBE, MAE, STDE, RMSE) -> (105 metrics)
- Multiple years are used for training



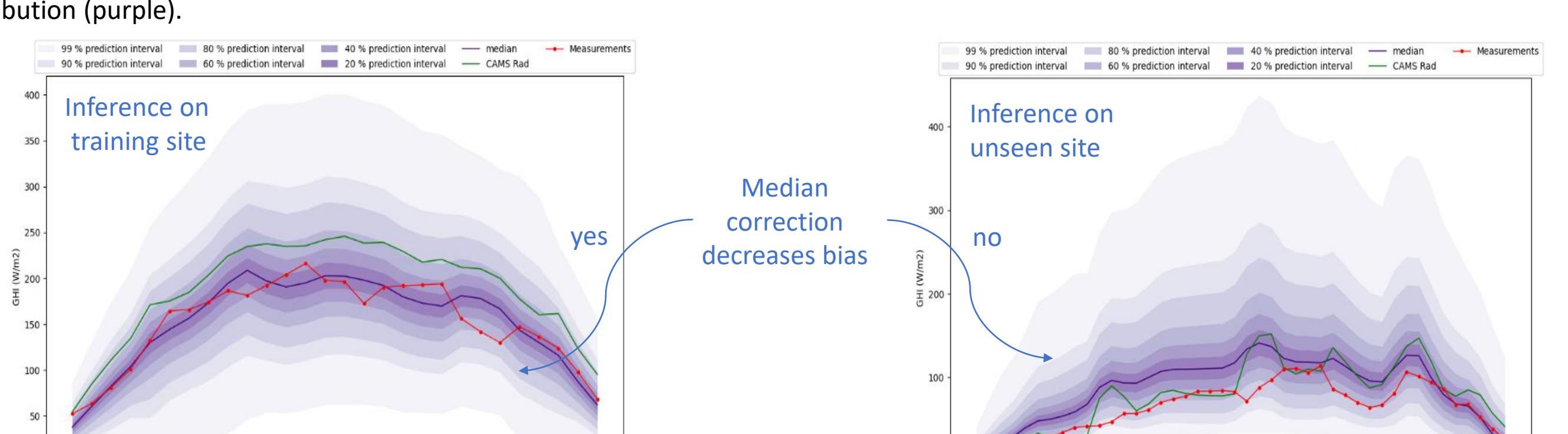
- Ground observation values well contained inside the 90% confidence interval (P5 - P95)
- The spread of the confidence intervals correlates well with the local variability situation, i.e., narrow intervals on clear and overcast situations and wider intervals on variable situations.
- Potential as a bias correction methodology

Localized error model # 2: uncertainty inference based on Deep learning

- Uncertainty modelled by a Johnson SU parametric distribution with four parameters controlling the first four moments of the distribution (mean, variance, skewness, and kurtosis)
- Six key variables for the calculation of SSI in CRS are selected as predictors:
 - clear sky index calculated with CRS estimates (K_c)
 - cloud type
 - cloud cover
 - cloud optical depth
 - solar zenith angle (SZA)
 - solar azimuth angle (SAA)

- 4 stations and 2 years are used for training
- The four parameters are predicted as a function of six predictors using a neural network (implementation in tensorflow probability)

Uncertainty inference for 2 days: CRS-GHI (green) and ground measurements (red), CRS-GHI-corrected with the median of the uncertainty distribution (purple).



- A conditional evaluation indicates that the model is over-dispersive at low values of the clear sky index and under-dispersive at high values of the clear sky index for cases of unseen sites. This lack of spatial generalisation can be attributed to an overtraining issue.
- The model is well calibrated when testing and training data (not shown here) are same but a significant decrease in accuracy is observed when the model is applied to stations that were not used for training.

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- User's Guide and validation reports at <https://atmosphere.copernicus.eu/supplementary-services>

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