

Age of air from in situ trace gas measurements: insights from a new technique

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Abstract. The age of air is an important transport diagnostic that can be derived from trace gas measurements and compared to global chemistry climate model output. We describe a new technique to calculate the age of air, measuring transport times from the Earth's surface to any location in the atmosphere based on simultaneous in situ measurements of multiple key long-lived trace gases. The primary benefits of this new technique include (1) optimized ages of air consistent with simultaneously measured $SF₆$ and $CO₂$; (2) age of air from the upper troposphere through the stratosphere; (3) estimates of the second moment of age spectra that have not been well constrained from measurements; and (4) flexibility to be used with measurements across multiple instruments, platforms, and decades. We demonstrate the technique on aircraft and balloon measurements from the 1990s, the last period of extensive stratospheric in situ sampling, and several recent missions from the 2020s, and compare the results with previously published and modeled values.

1 Introduction

The age of air has been used as an observationally based diagnostic of atmospheric transport for a number of decades (e.g., Bischof et al., 1985; Woodbridge et al., 1995; Schoeberl et al., 2005; Engel et al., 2017; Ray et al., 2022). The stratospheric age of air is controlled by both the residual mean circulation and quasi-horizontal mixing (Waugh and Hall, 2002; Ploeger et al., 2014) and thus can give unique insight into these primary transport processes that govern

the distribution of radiatively and chemically important trace gases in the stratosphere. An ideal trace gas to calculate mean ages, the first moment of the age spectra, would have a linear growth rate and be essentially inert in the atmosphere. Since this ideal trace gas does not exist in nature, we use measured trace gases with long lifetimes and photochemical production or loss that can be accounted for and is in growth or decay. Most calculations of mean ages from in situ measurements have been based on a single trace gas with these characteristics, such as carbon dioxide $(CO₂)$ (Bischof et al., 1985; Schmidt and Khedim, 1991; Woodbridge et al., 1995; Boering et al., 1996; Andrews et al., 2001) or sulfur hexafluoride (SF_6) (Volk et al., 1997; Engel et al., 2008). In the case of $CO₂$, simultaneous measurements of methane (CH₄) and nitrous oxide (N_2O) are needed to adjust for CO_2 production and to account for the complication of interpreting the seasonal cycle of $CO₂$ for the youngest ages (Boering et al., 1996; Andrews et al., 2001).

Each age-of-air calculation technique from a single trace gas has complications that lead to uncertainty in the age values. These complications include accurately accounting for photochemical production or loss of a trace gas, such as the mesospheric loss of SF_6 (Loeffel et al., 2022; Garny et al., 2024), or accurately characterizing the boundary condition of a trace gas, whether at the tropopause or the Earth's surface (e.g., Engel et al., 2008). Comparisons between mean age estimates from different trace gas measurements have been difficult due to the sparsity of simultaneous measurements of multiple age-of-air species, but when available the differences have been shown to be significant (e.g., Leedham Elvidge et al., 2018). This uncertainty from the measurement-based estimates is compounded by the large disagreement among chemistry–climate models (CCMs) and reanalysis-based ages of air (Chabrillat et al., 2018; Ploeger et al., 2019; Abalos et al., 2021) that leaves us with a significant gap in our quantification of the stratospheric age of air and thus also the important processes that determine it and its changes over time.

One possibility to improve the calculation of age of air from observations is with the use of multiple simultaneously or near-simultaneously measured trace gases, which has been shown previously to have potential for calculating age of air with more fidelity compared to estimates from a single trace gas (e.g., Schoeberl et al., 2005; Ehhalt et al., 2007; Bönisch et al., 2009; Ray et al., 2022). The main advantage of using multiple trace gases is to reduce the uncertainties or biases caused by the characteristics of a single trace gas to define the age of air (e.g., the seasonal cycle of $CO₂$). A detail of these techniques is that each trace gas does not contribute equally to the determination of age of air; i.e., only the longest-lived trace gases that are in growth (or decay) can quantify the range of ages in the stratosphere (Podglajen and Ploeger, 2019). Thus, trace gases such as $CO₂$, $SF₆$ and carbon tetrafluoride (CF_4) , referred to as primary age trace gases, can be used to make independent estimates of the age, while shorter-lived trace gases such as trichlorofluoromethane (CFC-11) or chloromethane (CH₃Cl) can be used only in addition to the longer-lived trace gases to give information about the shortest ages (weeks to months) typically found in the tropopause region (Ray et al., 2022).

The main disadvantages of using multiple trace gases in an age-of-air calculation are (1) the previously mentioned sparsity of simultaneous measurements of the most suitable trace gases and (2) the detailed knowledge of boundary conditions for each trace gas that is required for accurate age-of-air estimates. The sparsity of simultaneous measurements is due to the different measurement techniques and sampling frequency for each trace gas and aircraft and balloon payloads that may not have included multiple instruments suitable for age-of-air trace gas measurements. The knowledge of boundary conditions for many trace gases, especially at either the tropical surface or tropopause, is limited by whether longterm measurements by a network are available. These disadvantages have confined age-of-air estimates with multiple tracers to limited times and locations (e.g., Luo et al., 2018; Ray et al., 2022).

In this work we demonstrate a new optimization technique to obtain mean age of air, along with several other transport diagnostics, from simultaneous measurements of multiple primary age trace gases. The technique is applicable to any primary age trace gases, and we use the examples of $SF₆$ and $CO₂$ here. This method is essentially a hybrid between the traditional single trace gas mean age estimates and the multiple trace gas techniques (e.g., Bönisch et al., 2009). As in Bönisch et al. (2009) and Ray et al. (2022), we leverage additional information beyond mean age obtained in the optimization of a solution from multiple trace gases. In addition, similar to those two studies, we use the Earth's surface as the boundary condition of age of air, since that is where we have long-term measurements of each trace gas. This method allows for transport characteristics of the upper troposphere and tropopause region to be included in addition to information about the latitudinal surface source regions. The use of SF⁶ measurements in the technique presented here is made possible by new estimates of SF_6 mesospheric loss (Garny et al., 2024) that are used in the optimization.

This paper focuses on the details of the new technique, with some example results using measurements from the 1990s and more recently from the 2020s. A follow-up paper will explore the differences in the age of air between these two time periods. The next section describes the data used in the examples shown in here. Section [3](#page-2-0) describes the method, and Sect. [4](#page-9-0) shows some results based on in situ measurements.

2 Data

For the stratospheric data in this study, we use aircraft- and balloon-based in situ measurements of SF_6 (Elkins et al., 1996; Moore et al., 2003; Hintsa et al., 2021), CO₂ (Daube et al., 2002), CH₄ and N₂O (Webster et al., 1994, 2001; Loewenstein et al., 2002) and balloon-based AirCore measurements of these four species (Baier et al., 2021; Li et al., 2023). AirCores (or larger-volume, dual-dimension Strato-Cores) are measured by two different analytical systems for trace gas species used in this analysis: by cavity ring-down spectrometry for CO_2 , N₂O and CH₄ and by a gas chromatograph coupled with an electron capture detector for SF_6 and N2O (Li et al., 2023). We use measurements from both the

1990s, the last time period of extensive stratospheric in situ sampling of age-of-air trace gases above 16 km, and from several recent missions and routine measurements in the 2020s. The 1990s ER-2 aircraft missions include ASHOE-MAESA (1994), STRAT (1995–96), POLARIS (1997) and SOLVE (1999–2000) and balloon-based data from Observations of the Middle Stratosphere (OMS, 1996–2000). The 2020s airborne missions include the Dynamics and Chemistry of the Summer Stratosphere (DCOTSS, 2021–22, ER-2) and Stratospheric Aerosol, processes, Budget and Radiative Effects (SABRE, 2023, WB-57) aircraft campaigns, as well as routine AirCore sampling (2021–24, Karion et al., 2010; Baier et al., 2021; Li et al., 2023). Mean ages have been previously calculated from the measurements in the 1990s, which allows us to compare the results shown here with those values (Volk et al., 1997; Andrews et al., 2001), but no mean ages have yet been published based on the 2020s measurements.

In order to use the method described in the next section we need simultaneous measurements of several trace gases, which has been exceptionally rare even when the aircraft or balloon platform carried all of the required instruments due to different sampling and calibration timing. In situ $CO₂$, as well as CH_4 and N_2O , have typically been measured at $∼ 1-5$ s resolution, while SF₆ has been measured at $∼ 1-$ 2 min resolution. In order to markedly increase the number of sample times with simultaneous mole fractions of the four trace gases considered here, we interpolated in time across measurement gaps of up to 100 s in one or more of the trace gases. This time gap corresponds to a horizontal distance of \sim 20 km at typical ER-2 aircraft flying speeds and a vertical distance of ∼ 0.3–0.5 km for typical balloon ascent and descent rates. For long-lived trace gases in the stratosphere such as those considered here, variability over the interpolated spatial scales is sufficiently small in almost all cases to not significantly affect the results.

A critical part of the method is to have latitudinally varying and time-varying surface mole fractions of the key trace gases listed above. For the surface boundary condition of each trace gas, we use the NOAA Greenhouse Gas Marine Boundary Layer (MBL) reference product [\(https://gml.noaa.](https://gml.noaa.gov/ccgg/mbl/index.html) [gov/ccgg/mbl/index.html,](https://gml.noaa.gov/ccgg/mbl/index.html) last access: 9 December 2023; Lan et al., 2023) when available. This is a latitudinally resolved MBL product derived from weekly measurements of the four trace gases considered in this study. For time periods before the beginning of the NOAA MBL product we extend the time series based on different sources depending on the trace gas. For CO_2 before 1979 and N₂O before 2001, we use the CMIP6 reference time series from Meinshausen et al. (2017). For SF_6 before 1997, we use scaled Cape Grim mixing ratios based on Engel et al. (2008), which are very similar to the CMIP6 reference values. For CH_4 before 1983, we use annual average mixing ratios from Etheridge et al. (1998). For each trace gas we add on an average seasonal cycle and latitudinal gradient based on the earliest available 5 years of the NOAA MBL product.

3 Method

The mixing ratio, χ , of a trace gas, *i*, at a location, *x*, and time, t , can be expressed as follows:

$$
\chi_{\rm i}(x,t) = \int_0^\infty \chi_{\rm i0}(t-t') e^{-t'/\tau_{\rm i}(x,t')} G(x,t,t') \mathrm{d}t',\tag{1}
$$

where G is the age spectrum, t' is the transit time from a source region to the location x, $\chi_{\text{io}}(t-t')$ is the mixing ratio time series at the source region and τ_i is the path-dependent lifetime of the trace gas (Fig. [1\)](#page-3-0) (Schoeberl et al., 2000; Ehhalt et al., 2007). The age spectrum represents the distribution of transit times since an air parcel was last at the source region as shown schematically by the different colored lines in Fig. [1.](#page-3-0) The age spectra are assumed to have an inverse Gaussian functional form given by

$$
G(x, t, t') = \sqrt{\frac{\Gamma^2}{4\pi R t'^3}} \exp\left(\frac{-\Gamma^2 (t' - \Gamma)^2}{4R t'}\right),
$$
 (2)

where the mean age is given by $\Gamma = \int_0^\infty t' G(x, t, t') dt'$, the width of the spectrum is given by $\Delta^2 = \frac{1}{2} \int_0^\infty (t' (\Gamma)^2 G(x, t, t') dt'$ and the ratio of moments is given by $R =$ Δ^2/Γ (Hall and Plumb, 1994; Waugh and Hall, 2002).

The ratio of moments R , which is a measure of the width of the age spectra, is a largely unconstrained quantity, and thus we allow a range of values from 0.1–2.5 years for each possible Γ in the optimization calculation (Fig. S1 in the Supplement). Most studies of age of air have used values of R ranging from 0.7–1.25 years based on model estimates (e.g., Hall and Plumb, 1994; Volk et al., 1997; Engel et al., 2008). However, with a better understanding of the effect of the exponential tail of G for $t' > 10$ years, the model estimates of R have increased to values of 1.5 years or more, with considerable variability in the stratosphere (e.g., Diallo et al., 2012; Ploeger and Birner, 2016; Fritsch et al., 2020).

The source region for our calculation is the Earth's surface even though most previous stratospheric age-of-air studies have used the tropical tropopause since that is the primary entry region for air to the stratosphere above 380 K (e.g., Andrews et al., 2001; Engel et al., 2009). There are a number of reasons we choose to use the surface as the source region: (1) we have long-term measurements of all the trace gases at the surface, which is not the case for the tropical tropopause; (2) we can obtain information about the surface source latitude regions with our optimization technique; (3) most modeled mean ages use the surface as the source region; and (4) we can calculate ages in the troposphere and lowermost stratosphere with the same technique as the stratospheric overworld.

Following the techniques of Hauck et al. (2020) and Ray et al. (2022), we partition the age spectra into a part with

Figure 1. Schematic of the age of air with the surface as the source region. Different pathways to the sampled region (x) are represented by the colored lines with different surface latitudinal origins (y) and timescales of transport (t') represented by the age spectra (G) .

a tropical (30 \degree S–30 \degree N average) surface source, g_{TR} , and a part with a latitudinally varying surface source, g_V . The partitioning of the age spectra in this way is primarily due to a recognition that the extratropical upper troposphere–lower stratosphere (UTLS) can be influenced by local convective transport from the extratropical surface (e.g., Herman et al., 2017; Ray et al., 2022). A surface source that is only represented by a tropical average, as in most previous age-of-air studies, will not account for these extratropical surface influences that can be important for trace gases with large latitudinal gradients. The partitioned age spectra are expressed as

$$
G(\Gamma, \Delta, t') = g_{\text{TR}}(\Gamma, \Delta, t') + g_{\text{V}}(\Gamma, \Delta, t').
$$
\n(3)

The lowercase g is used to indicate age spectra that are non-normalized, which is in contrast to the uppercase G used to indicate total age spectra that are normalized. The age spectra g_V do not vary with latitude but are convolved with surface mixing ratio time series from varying latitudes as part of the optimization, as will be shown in the next section, whereas the age spectra g_{TR} are always convolved with the tropical average surface mixing ratios. Note that we have now expressed the age spectra as a function of Γ and Δ since those are the parameters in G we allow to vary in the convolutions with the $\chi_{\rm io}$ time series. The partitioned age spectra are expressed as fractions, f, of the \overline{G} as a function of t'.

$$
g_{\text{TR}}(\Gamma, \Delta, t') = f(t')G(\Gamma, \Delta, t')
$$
\n(4)

$$
g_V(\Gamma, \Delta, t') = (1 - f(t'))G(\Gamma, \Delta, t')
$$
\n(5)

The fraction f has an age dependence with $f(t' < t'_i) = 0$, $f(t' > t'_f) = 1$ and an exponential form between t'_i and t'_f , which are the transition ages between the latitudinally varying and purely tropical surface source regions. Based on Ray et al. (2022), where a value of $t_f' \approx 150$ d was found to be

optimal above 380 K, we use values of $t'_i = 0.4$ years and $t_f = 0.6$ years (Fig. S2). The Ray et al. (2022) study showed the importance of the extratropical surface as a source region for certain locations and seasons in the lower stratosphere, such as over North America during the monsoon season. But for other locations and seasons the tropical surface may be the most important source region for the youngest ages in the stratosphere, and the calculation allows for that possibility.

3.1 Convolutions

Convolutions of the age spectra with the surface mixing ratio time series for each trace gas are performed for the time period of available in situ measurements. The convolved mixing ratios, χ_i , are given by

$$
\chi_{\rm i}(\Gamma,\Delta,y_{\rm o},t) = \chi_{\rm iTR}(\Gamma,\Delta,t) + \chi_{\rm iV}(\Gamma,\Delta,y_{\rm o},t),\tag{6}
$$

where the subscript σ refers to a surface quantity, y_0 is a surface source latitude parameter, and

$$
\chi_{\text{ITR}}(\Gamma, \Delta, t) = \int_0^\infty \chi_{\text{ioTR}}(t - t', y_{\text{oTR}}) g_{\text{TR}}(\Gamma, \Delta, t') \, \text{d}t', \tag{7}
$$

$$
\chi_{\text{IV}}(\Gamma,\Delta,y_{\text{o}},t) = \int_0^\infty \chi_{\text{ioV}}(t-t',y_{\text{o}})g_{\text{V}}(\Gamma,\Delta,t')dt'.
$$
 (8)

The surface mixing ratio time series χ_{ioTR} are tropical averages from 30° S–30° N (y_{oTR}) in the case of χ_{iTR} or latitudinally varying averages in 10° intervals ranging from 60° S– 60° N (y_0) in the case of χ _i χ . Note that there is no latitudinal dependence for χ_{ITR} since y_{oTR} represents a single tropical average latitude region.

Examples of χ _{SF6} and χ _{CO₂} as a function of Γ for early November 1994 are shown in Fig. [2.](#page-4-0) The dependence of χ_i on y_0 and R are shown separately to reveal the Γ range most

Figure 2. Convolutions of SF_6 (a, c) and CO_2 (b, d) surface mixing ratios with a range of age spectra and mean ages. These convolutions could be compared to stratospheric measurements of each trace gas in early November 1994 (NH fall). The dependence of the convolutions on surface source latitude is shown in (a, b) , while the dependence on the ratio of moments is shown in (c, d) . The grey shading represents the full range of possible mixing ratios, with both source latitude and ratio of moments allowed to vary. The dashed green line represents convolutions with only a tropical surface source.

sensitive to these parameters; χ_i varies most strongly with y_0 for Γ < 2 years (Fig. [2a](#page-4-0) and b). This is expected due to the functional form of g_V , which peaks for $t' < 0.5$ years. Even for $\Gamma = 2$ years, g_V is a significant portion of the total age spectrum G (Fig. S2). But for $\Gamma > 2$ years the contribution of g_V to G decreases rapidly and thus the dependence on y_0 does so as well. The values of t'_i and t'_f are of course not known precisely and likely vary with time and location, as shown for a limited time and region in Ray et al. (2022). Changes in the assumed timescales of transport from y_0 will change the χ_i distributions in Fig. [2](#page-4-0) somewhat, but we do not expect moderate timescale changes to change our results. One reason for this is the frequent opposite dependence of the two trace gas convolutions on y_0 , meaning that even if the χⁱ distributions were broadened based on different assumed transport timescales, the optimal solutions will be confined to the relatively narrow overlap region of the distributions. We do not explore the effects of the values of t'_i and t'_f any further in this work, but it is an interesting aspect that would be helpful to understand better.

The surface latitudinal gradient of χ_{CO_2} is known to have substantial seasonal variability with larger values of χ_{CO_2} in the NH during boreal winter and a reversed latitudinal gradient in boreal summer [\(https://gml.noaa.gov/ccgg/mbl/,](https://gml.noaa.gov/ccgg/mbl/) last access: 1 June 2024), whereas the latitudinal gradient of $\chi_{\rm SF6}$ has essentially no seasonal variability with larger values in the NH. For NH fall, the contrast between the χ _{SF6} and χ _{CO₂} dependence on y_{oV} for $\Gamma < 2$ years is stark, with larger values of χ _{SF₆} for more northern values of y_{0} and the opposite for χ_{CO_2} . This is due to the opposite latitudinal surface gradients of these trace gases in the preceding NH summer, shown by the minimum of χ_{CO_2} for the most northern values

of y_0 at $\Gamma = 0.3$ years. The reversed latitudinal dependence of χ _{SF6} and χ _{CO₂} persists to $\Gamma \approx 2$ years since g_V remains influenced by the NH summer for all values of Γ . Values of χ_{CO_2} for April have the opposite latitudinal dependence (Fig. S3) to that in November due to presence of the opposite latitudinal surface gradient of CO₂ in NH winter compared to summer. For $\Gamma > 3$ years, the values of χ_i have negligible dependence on y_{oV} and correspond to those using only χ_{ioTR} (dashed green lines in Fig. [2\)](#page-4-0). Thus, for older mean ages the values of χ_i are similar to previous studies that used the average of Mauna Loa and Samoa station data (e.g., Andrews et al., 2001).

The dependence of χ _{SF₆} and χ _{CO₂} on R for $y_0 = 10^\circ$ N contains a significant spread for nearly all values of Γ but especially for $\Gamma > 2$ years, in contrast to the dependence on y_0 (Fig. [2c](#page-4-0) and d). In NH fall, χ _{SF6} and χ _{CO₂} have the opposite dependence on R for $\Gamma > 1$ year, while for NH spring the dependence is the same for both trace gases (Fig. S3c and d). This again reflects the seasonal change of the $CO₂$ surface gradient and how the shape of the age spectra can make the convolutions sensitive to a particular season.

A noticeable feature of the NH fall convolutions is the wide range of possible values of χ_i for the same Γ , especially χ_{CO_2} for Γ < 2 years as shown by the grey shading (Fig. [2\)](#page-4-0). The implication of this is that unless there is high confidence in y_0 and R for a particular air mass, the uncertainty in Γ from either SF_6 or CO_2 alone can be significant. But in combination, the intersection of χ_{SF_6} and χ_{CO_2} can provide a much narrower range of possible Γ , which we show below to not only reduce the uncertainty in Γ compared to using a single trace gas but also reveal optimal values of y_0 and R.

3.2 Adjustments for photochemistry

Before two trace gases can be used in an optimization of age of air, their photochemical lifetimes (Eq. [1\)](#page-2-1) and thus potential adjustments to the measured mixing ratios (χ_i^*) must be considered. A reduction in $\chi^*_{\text{CO}_2}$ has traditionally been made before Γ is calculated due to the conversion of CH₄ to CO₂ in the atmosphere (e.g., Boering et al., 1996; Daniel et al., 1996). This adjustment requires simultaneous measurements of CH_4 or N₂O along with CO_2 and has assumed that the difference between the surface average $CH₄$ at the time of the measurement and the CH⁴ measured in the stratosphere is entirely due to the oxidation of CH_4 to CO_2 . In the case where a simultaneous $\chi_{CH_4}^*$ was not available and $\chi_{N_2O}^*$ was, then a CH_4-N_2O relationship has been used to approximate $\chi^*_{\text{CH}_4}$ (e.g., Andrews et al., 2001).

While the adjustment of $\chi_{\text{CO}_2}^*$ due to CH₄ oxidation is typically small since CH⁴ surface mole fractions have been \sim 1.7–1.9 ppm over the past few decades, the effect on Γ estimates can be significant based on the nonlinear shape of the χ_{CO_2} distributions shown in Fig. [2b](#page-4-0) and d. Furthermore, it has been shown in modeling studies that not all of the oxidized CH₄ is converted into $CO₂$ (Boucher et al., 2009; Shindell et al., 2017). Although the average conversion percentage in these two studies was 61 % and 88 %, respectively, essentially all of the conversion in the stratosphere ends up as CO2. Based on these results, we use a 95 % average conversion of CH_4 into CO_2 , and we also increase the uncertainty of $\chi^*_{\text{CO}_2}$ by 5 % of the CH₄ loss to account for some uncertainty in the conversion percentage.

Another aspect of the $CO₂$ adjustment due to CH₄ oxidation that we reevaluate is the use of the surface average CH⁴ mixing ratio at the time of measurement to subtract from $\chi_{\text{CH}_4}^*$. The mixing ratios that should be used are the convolution values χ_{CH_4} since these represent the actual boundary conditions of CH₄. The values of χ _{CH₄} have a larger dependence on $y_{0}y$ (Fig. S4) than either the Γ or R. The difference between the surface average $CH₄$ at the time of stratospheric measurement used in previous studies and values of χ_{CH_4} is typically 0.02–0.03 ppm but can be greater than 0.05 ppm for Γ < 1 year. On its own, this difference and the subsequent change in the adjustment of $\chi_{CO_2}^*$ are small, but as mentioned earlier they can have a nonlinear effect on the age calculation. The use of χ C_{H₄} is also simply the correct technique to estimate the entry or initial CH₄ mixing ratios for any measurement location, especially in an optimization with multiple tracers.

For $SF₆$ it has been known for some time that mesospheric loss has an impact on age-of-air calculations with this trace gas (Hall and Waugh, 1998; Hall et al., 1999; Andrews et al., 2001). The old age bias in Γ from SF_6 was thought to be largely confined to air masses in or near the winter polar vortex (e.g., Ray et al., 2017) and not in the midlatitudes (Engel et al., 2009; Ray et al., 2014). Based on recent studies, we now understand that the old age bias in Γ from SF₆ has a latitudinal and temporal dependence on the mixing ratio of $SF₆$ and that a Γ bias correction can be found and applied (Loeffel et al., 2022; Garny et al., 2024). This is an important advancement in the use of SF_6 as an age-of-air tracer, and without it an optimization of the kind described here would not be possible.

We make $SF₆$ adjustments based on a correction technique from the work of Garny et al. (2024) that parameterizes the expected Γ bias as a function of Γ and time. We convert this Γ bias into a mixing ratio adjustment via the SF₆ tropospheric growth rate over the previous year and add this estimated mesospheric loss to $\chi_{SF_6}^*$. There are uncertainties inherent to this technique, primarily due to the relatively new Γ bias estimates that are based on a single model thus far. We do not explicitly add an additional uncertainty to the adjusted values of $\chi^*_{SF_6}$ (denoted as $\chi^*_{aSF_6}$), but we do investigate possible uncertainties in the adjustments through a scaling method that is described below.

3.3 Optimization

We now look at an example of a specific measurement time and place to demonstrate the optimization technique. The ex-

Figure 3. Zoomed in convolutions similar to Fig. [2](#page-4-0) but also including the simultaneous measured SF₆ and CO₂ mixing ratios from the ASHOE-MAESA ER-2 flight on 4 November 1994 at 18:27 UTC (dashed black lines, with uncertainties indicated by the dotted black lines). The adjusted mixing ratios and uncertainties are indicated by the dashed and dotted blue lines, respectively. The range of solutions for SF_6 alone ($\chi_{SF_6}(\Gamma_{SF_6})$) are indicated by the magenta symbols in (a), and the range of solutions for CO₂ alone ($\chi_{CO_2}(\Gamma_{CO_2})$) are indicated by the orange symbols in (b). The values of χ_{SF_6} with the CO₂ solutions ($\chi_{SF_6}(\Gamma_{CO_2})$) are shown by the orange symbols in (a), and likewise the values of χ_{CO_2} with the SF₆ solutions ($\chi_{CO_2}(\Gamma_{SF_6})$) are shown by the magenta symbols in (b). The intersection of solutions that agree with both trace gases within uncertainties are indicated by the green symbols and labeled $\chi_{SF_6}(\Gamma \cap)$ and $\chi_{CO_2}(\Gamma \cap)$. Dashed lines are as described in the caption of Fig. S5.

ample is from an ER-2 flight in early November 1994 during the ASHOE-MAESA campaign at a time when simultaneous $\chi_{SF_6}^*$ and $\chi_{CO_2}^*$ measurements were available. The range of χ_{CO_2} and χ_{SF_6} values is the same as shown by the grey shading in Fig. [2](#page-4-0) since those values are relevant for the November 1994 time period. The photochemically adjusted χ_i^* values, denoted as $\chi_{ai}^*(\Gamma)$, have a dependence on Γ since for SF₆ the older the Γ is, the more loss will have occurred (as described above). Likewise, for CO_2 the $\chi_{CH₄}$ values are dependent on Γ . The difference between the χ_i^* and $\chi_{ai}^*(\Gamma)$ values for both trace gases are shown in Figs. S5 and [3.](#page-6-0) The uncertainty of $\chi_{ai}^*(\Gamma)$ is indicated by χ_{eai}^* , but for the two species used here only $CO₂$ has a change in the size of the uncertainty from the original measurement due to the uncertainty in CH⁴ production.

A range of possible mean ages, surface source latitudes and ratios of moments $(\Gamma, y_0 \text{ and } R)$ can be found for each trace gas based on agreement of the convolutions (χ_i) with the adjusted measured mole fractions ($\chi_{ai}^*(\Gamma)$) within the uncertainty range ($\pm \chi_{\text{eai}}^*$). For the example shown here, the values of Γ compatible with $\chi_{aCO_2}^*(\Gamma)$ range from 0.1–2.6 years, while for SF_6 the Γ range is 2.4–2.9 years (Figs. S5 and S6). Figure [3](#page-6-0) zooms in on the range of Γ and χ_i around where both trace gases have similar Γ solutions. The χ_{SF_6} values that agree with $\chi_{\text{aSF}_6}^*(\Gamma)$ within $\pm \chi_{\text{eSF}_6}^*$ are shown by the magenta symbols and labeled $\chi_{SF_6}(\Gamma_{SF_6})$ in Fig. [3a](#page-6-0), and for CO_2 this is shown by the orange symbols and labeled $\chi_{\text{CO}_2}(\Gamma_{\text{CO}_2})$ in Fig. [3b](#page-6-0). The Γ resolution of our calculation was 0.05 years,

which is why there are distinct columns of symbols in the figure.

Since any optimal solutions must be compatible with both trace gases, we need to look at how the possible Γ , Δ , y combinations from one tracer affect the χ_i values of the other tracer. The values of χ_{SF6} with CO₂ solutions are shown by the orange symbols in Fig. [3a](#page-6-0) and labeled $\chi_{SF_6}(\Gamma_{CO_2})$, and likewise the values of χ_{CO_2} with SF₆ solutions are shown by the magenta symbols in Fig. [3b](#page-6-0) and labeled $\chi_{\text{CO}_2}(\Gamma_{\text{SF}_6})$. Most of the $\chi_{\rm SF_6}(\Gamma_{\rm CO_2})$ values are greater than the $\chi_{\rm aSF_6}^*$ values within $\pm \chi_{\rm eSF_6}^*$, while most of the $\chi_{\rm CO_2}(\Gamma_{\rm SF_6})$ values are less than the $\chi_{aCO_2}^{\text{on}}(\Gamma)$ values within $\pm \chi_{\text{eaCO}_2}^{\text{on}}(\Gamma)$. However, there is a region of overlapping solutions for both trace gases, as indicated by the green symbols and labeled $\chi_{SF_6}(\Gamma_{\cap})$ and $\chi_{\text{CO}_2}(\Gamma_{\cap})$, where Γ_{\cap} represents the intersection of possible Γ , Δ , y combinations.

The intersection of solutions is the basis of the optimization used to find a single best solution for each set of measurements. We first find normalized differences expressed as $\delta \chi_i = |\chi_i - \chi_{\text{ai}}^*| / \chi_{\text{eai}}^*$ (Fig. S6). A combined difference quantity, $\delta \chi_{\cap}$, for the intersection of solutions is defined as the average of the $\delta \chi_i$ values (Fig. S7),

$$
\delta \chi_{\cap} = (\delta \chi_{\text{CO}_2}(\Gamma_{\cap}) + \delta \chi_{\text{SF}_6}(\Gamma_{\cap})) / 2. \tag{9}
$$

From $\delta \chi_{\cap}$, we define weighting functions W_{\cap} = $(\max(\delta \chi_{\cap}) - \delta \chi_{\cap}) / \sum \delta \chi_{\cap}$ to be multiplied by the set of transport parameters within the intersection of solutions to find the optimum values

$$
\Gamma_{o} = W_{\cap} \Gamma_{\cap} \Delta_{o} = W_{\cap} \Delta_{\cap} y_{o} = W_{\cap} y_{\cap}. \tag{10}
$$

Thus, for each measurement location there is a single set of transport parameters Γ_0 , Δ_0 , y_0 that are optimized for χ_{ai}^* of both trace gases (Fig. S8).

Note that the optimization can be performed for a single trace gas by following the same general procedure described above. The range of possible transport parameters is larger for a single trace gas compared to an optimization with two or more trace gases, and thus the uncertainty of the results will be larger, and there will be a higher probability that the optimized transport parameters are "incorrect". In the example measurement location used here, if the optimization were performed only on CO_2 , Γ_0 would be much younger than the result with both SF_6 and CO_2 .

3.4 Offsets to measured mixing ratios

In some cases, the intersection of possible solutions for SF_6 and $CO₂$ within uncertainties is small or there is no intersection at all. This lack of overlap could be due to various reasons, such as an inconsistent calibration between surface and atmospheric measurements or an inaccurate correction for photochemical loss or production of a trace gas. Since there can be only one set of transport parameters in the real atmosphere at any one time and place, a lack of an intersection of solutions thus implies an offset in one or both trace gases at the measurement location is necessary to perform the optimization.

To account for the possibility that an offset of either trace gas is needed to obtain a solution set, we ran an ensemble of optimizations (see the Supplement for further details) that swept over a grid of mixing ratios with positive and negative offsets from the measured value of each trace gas for each measurement location. Note that we offset the adjusted (χ_{ai}^*) mixing ratios and use the same uncertainties (χ_{eai}^*) . The optimization performed for each member of the offset ensemble produces a unique set of transport parameters $\Gamma_0(s)$, $\Delta_0(s)$, $y_0(s)$ and a minimum value of the combined normalized difference $\delta \chi_{\cap \min}(s)$, with the dependence on s added to denote the offset ensemble member (Fig. S9).

There are a number of possible ways to select the best member of the ensemble and thus the best overall set of solutions. We chose to use an optimization method that is consistent with that used for each member of the ensemble as described above, and the details are included in the Supplement. Most of the offset values are small, especially for $CO₂$, but there are certain measurement times and locations where the optimal offset of $SF₆$ is significant. Fig-ures [4](#page-8-0) and [5](#page-9-1) show SF_6 and CO_2 average optimal offset values as a function of normalized N_2O for the 1990s and 2020s data, respectively. Normalized N_2O (χ_{N_2On}) is defined as $\chi^*_{N_2O}(t)/\overline{\chi_{N_2O}O}(t)$, i.e., the measured atmospheric N₂O mixing ratio divided by the global mean surface N_2O mixing ratio at the time of measurement. N_2O has often been used as a quasi-vertical coordinate when displaying in situ data due to the compact relationships of long-lived trace gases and Γ

with N_2O (e.g., Plumb and Ko, 1992; Andrews et al., 2001). We use normalized N_2O in this figure and others in order to account for the growth rate of N_2O over time, enabling the comparison of compact relationships with N_2O over years or decades.

The offset values are also normalized by the global average surface mixing ratios of each trace gas and displayed as a percentage. There are several features of note in Figs. [4](#page-8-0) and [5.](#page-9-1) The first is that $\chi_{\text{CO}_{28}}$ (where the subscript s refers to an offset value) is essentially zero, < 0.01 % compared to an annual growth rate of 0.3 %–0.5 %, for all values of N_2O for both the 1990s and 2020s data. This follows from the symmetric nature of the offset ensemble results in $CO₂$ as shown by the example in Fig. S9. This implies that there is no advantage in the optimization calculation to offset $CO₂$ either positively or negatively in order to better agree with any set of $SF₆$ solutions. Further implications are that the calibration between $CO₂$ measured in situ in the atmosphere and the surface have remained consistent over several decades and that the photochemical adjustment of $CO₂$ does not appear to have a significant bias.

A second feature of note is the size of χ _{SF6s} for the 1990s data compared to that from the 2020s. In the 1990s, χ_{SF_6s} ranges from -9% to $+7\%$ at a time when the annual growth rate of SF₆ was \sim 5 % yr⁻¹, while in the 2020s χ _{SF6}s ranges from -2% to $+1\%$ at a time when the annual growth rate was ~ 3.5 % yr^{-1} . Thus, the χ _{SF6s} values translate to Γ shifts from -1.8 to $+1.4$ years in the 1990s and -0.5 to $+0.3$ years in the 2020s. As mentioned earlier, these $\chi_{SF,6s}$ values could be due to an inaccurate SF_6 photochemical loss adjustment that we apply in the optimization based on the Garny et al. (2024) study. If the photochemical loss was estimated to be larger than reality, $\chi_{aSF_6}^*$ would be too large, and the optimization would compensate by producing negative χ _{SF6s} values to agree with $CO₂$ and vice versa. The fact that there is essentially an equal amount of positive and negative χ _{SF6s} values, outside of the SOLVE mission that we will discuss below, and that the 2020s χ _{SF6s} values are mostly near zero suggests that there is not a systematic error in the photochemical loss adjustments. There can be certain locations and times when the adjustment could be too large or small but this is likely due to known limitations of the simplified adjustment used here, such as in the region of the polar vortex where the gradients of $SF₆$ loss are large.

A third feature, notable in Fig. [4,](#page-8-0) is the nearly linear dependence and large negative values of $\chi_{SF,6}$ as a function of N_2O during the SOLVE mission. These large negative values could be due to either assumed photochemical loss that is much too large or a calibration issue with $\chi^*_{SF_6}$ during the mission. Based on the outlier nature of $\chi_{\rm SF,65}$ during SOLVE compared to other missions, as well as the size of the photochemical loss error that would be necessary to cause these values ($> 100\%$ for $\chi_{N_2\text{On}} > 0.3$), it is unlikely the assumed photochemical loss is the primary cause. There was a known SF_6 calibration scale change during SOLVE

Figure 4. Offset values (χ_{is}) divided by the global average surface mixing ratios of each trace gas at the time of measurement and binned by normalized N₂O (χ _N, χ _{Dn}) for SF₆ (a, c) and CO₂ (b, d). The top row (a, b) shows normalized offset values binned separately for four different aircraft missions in the 1990s, and the bottom row (c, d) shows normalized offset values averaged over seven OMS balloon flights in the 1990s. Essentially all of the data represented in this figure are from the stratosphere at altitudes up to ∼ 32 km.

[\(https://espoarchive.nasa.gov/archive/browse/oms/Balloon\)](https://espoarchive.nasa.gov/archive/browse/oms/Balloon) that resulted in the value of the Volk et al. (1997) delay term from the surface to the tropical tropopause $(\delta \Gamma)$ being changed from $+0.8$ years as used in previous missions to -0.8 years in SOLVE. A negative value of $\delta\Gamma$ is obviously non-physical but was necessary to calculate reasonable Γ values due to the relatively large $\chi^*_{SF_6}$ compared to the surface values. This change in the value of $\delta\Gamma$, even considering the different source region of the surface compared to the tropical tropopause, is enough to explain most of the negative χ _{SF6}s values from the optimization. Thus, this analysis clearly points out a significant high bias in $\chi^*_{SF_6}$, especially at the highest altitudes and latitudes (lowest χ_{N_2On} , measured on the ER-2 during the SOLVE mission.

3.5 Single trace gas optimizations

In order to utilize the large number of in situ measurements with no simultaneous or near-simultaneous measurements of a second age-of-air trace gas, we also perform the optimization described here on a single trace gas. A single trace gas optimization is not ideal, but we make use of the relatively tight $\Gamma - \chi_{N_2On}$ and $R - \chi_{N_2On}$ correlations derived from the two trace gas optimizations to constrain those parameters for a single trace gas optimization (Figs. [7](#page-11-0) and S10). Based on the measured χ_{N_2On} at any location, we choose a range of possible Γ and R consistent with the correlations based on other missions during a similar time period and perform the optimization only within the limited range of each parameter. Thus, our technique requires a near-simultaneous measurement of N_2O with any single age-of-air trace gas.

Figure 5. The same as Fig. [4](#page-8-0) but for two DCOTSS aircraft missions in the 2020s (a, b) and for routine AirCore balloon flights in the 2020s (c, d).

The lack of simultaneous measurements can be due to a mismatch in instrument sampling rates, as is the case for $CO₂$ and $SF₆$ mentioned in Sect. [2,](#page-1-0) or the lack of measurements of two different age-of-air trace gases on a particular airborne platform or mission. The single trace gas optimization is often most desirable with $CO₂$ since there are usually nearly 2 orders of magnitude more measurements available during a mission compared to SF_6 . Since χ_{CO_2s} is negligible for all of the measurements considered in this study, we can reasonably perform the optimization with $CO₂$ alone without scaling. In the 2020s, χ _{SF₆s} is also minimal, so we can perform the single trace gas optimization with SF_6 , e.g., with data from the SABRE mission in 2023 when $CO₂$ measurements were not available.

4 Results

4.1 Results for the 1990s

The 1990s was the last time period of extensive in situ measurements from both aircraft and balloon platforms of longlived trace gases and Γ calculated from CO₂ and SF₆ (denoted Γ_{CO_2} and Γ_{SF_6}) in the stratosphere at altitudes above 16 km. The Γ_{CO_2} and Γ_{SF_6} values from this time period have been used extensively over the past several decades to compare to model output, especially the use of the "wing plot" of Γ in the 20 km altitude region as a function of latitude (Fig. [6\)](#page-10-0) (Waugh et al., 1997; Hall et al., 1999; Andrews et al., 2001; Waugh and Hall, 2002; Strahan et al., 2011; Diallo et al., 2012; Chabrillat et al., 2018; Dietmuller et al., 2018; Ploeger et al., 2019). The general features of Fig. [6](#page-10-0) are similar to versions of this plot shown in previous publications. The seasonally resolved averages reveal that the spread of

Figure 6. Mean age at 20 km altitude as a function of latitude for aircraft measurements from the 1990s, commonly referred to as the "wing plot". The averages for different seasons, as well as for all times at each latitude, are shown separately.

the mean ages in most latitude bins is driven by expected seasonal variability, most noticeable in the NH subtropics and polar regions, rather than an uncertainty on the measurements or mean age estimates. The latitudinal averages in Fig. [6](#page-10-0) also have missing values and some outlier points, both of which are expected from aircraft sampling even during this peak stratospheric sampling period.

The compact $\Gamma - \chi_{N_2On}$ relationship has also been commonly used as a transport metric and method to compare in situ data and model output (e.g., Hall et al., 1999; Andrews et al., 2001; Strahan et al., 2011; Birner et al., 2020). As men-tioned in Sect. [3.5,](#page-8-1) we utilize the $\Gamma_{o} - \chi_{N_2On}$ relationship in our optimizations for cases with a single trace gas measurement, and we also can compare the relationships found here with those from previous Γ techniques.

Figure [7](#page-11-0) shows $\Gamma - \chi_{\text{N}_2\text{On}}$ correlations for all of the 1990s aircraft and balloon data. Each set of correlations comes from essentially the same sampled latitudes, altitudes and seasons for each measurement platform. The correlations include measurements from the tropics to polar latitudes and nearly all seasons for the aircraft data, demonstrating the remarkable compactness of the correlations. The Γ_{SF_6} bias of up to2 years due to SF_6 photochemical loss at that time is seen by comparing the Γ_{SF_6} and Γ_{CO_2} . Values of Γ_0 typically fall between the values of Γ_{SF_6} and Γ_{CO_2} . We clearly expect Γ_0 to be younger than most of the Γ_{SF_6} values since we account for photochemical loss, while we also expect Γ_0 to be older than $\Gamma_{\rm CO_2}$ by at least 3 months since the source region of our calculation is the surface compared to the tropical tropopause for the previous calculation. To compare more directly to previous mean age estimates we could, for instance, subtract 3 months from Γ_0 , but the point of Fig. [7](#page-11-0) is to show that the mean ages calculated in this study are different from previous estimates, which is to be expected for the reasons discussed above.

The aircraft Γ_{CO_2} from the 1990s are based on $\chi_{N_2O}^*$ rather than $\chi_{\text{aCO}_2}^*$ for values of $\chi_{\text{N}_2\text{On}} > \sim 0.8$ (Andrews et al., 2001), which explains the tight relationship and lack of variability in Γ_{CO_2} in this range in Fig. [7a](#page-11-0). This relationship was based on measurements in the tropical tropopause region during this time. While this Γ_{CO_2} – $-\chi_{\text{N}_2\text{On}}$ relationship was valid for the particular time and place of the original measurements, a fixed relationship of this sort does not allow detection of changes in the relationship over time or variation in latitude. The Γ_0 values in this study are derived independently of any $\Gamma - \chi_{N_2On}$ relationship and yet produce very compact curves for both the aircraft and balloon data as seen in Fig. [7.](#page-11-0) There is at least an order of magnitude more aircraft measurements compared to balloon measurements resulting in the relatively compact relationships in Fig. [7a](#page-11-0) compared to those in Fig. [7b](#page-11-0).

Average distributions of the transport parameters derived from the optimizations with the 1990s aircraft measurements are shown in Figs. [8](#page-11-1) and S11. The two figures reveal the difference in spatial coverage of where there were near simultaneous measurements of SF_6 and CO_2 (Fig. S11) compared to $CO₂$ alone (Fig. [8\)](#page-11-1). We focus here on the $CO₂$ optimizations shown in Fig. [8](#page-11-1) since the more extensive number of measurements and spatial coverage gives a nearly complete picture of the average transport features in the upper troposphere and lower stratosphere.

One of the most notable features of the distributions in Fig. [8](#page-11-1) is that they cover a range of latitudes and altitudes including the troposphere, lowermost stratosphere and stratospheric overworld nearly from pole to pole. There have been no previous Γ estimates from in situ measurements that span this spatial range. This is partly due to the original concept of age of air as a timescale since an air parcel crossed the tropical tropopause into the stratospheric overworld (Kida, 1983). Thus, many previous Γ estimates from in situ data have focused only on the stratospheric overworld with the tropical tropopause as the entry region (e.g., Boering et al., 1996; Volk et al., 1997; Andrews et al., 2001; Engel et al., 2008). Subsequent studies have focused on the lowermost stratosphere with a modified age-of-air concept to account for multiple tropopause entry regions (e.g., Bönisch et al., 2009; Hauck et al., 2019, 2020). Interhemispheric tropospheric age of air from the NH surface to the SH surface has also been estimated from in situ surface measurements (Waugh et al., 2013; Holzer and Waugh, 2015). The Ray et al. (2022) study calculated Γ from the troposphere through the stratospheric overworld but just from one aircraft mission in the NH midlatitudes.

In general, the Γ_0 distributions look as expected, with the youngest ages in the troposphere, near the tropopause and in the tropical stratosphere and the oldest ages at the highest latitudes and altitudes. Values of Γ_0 at the tropopause vary with latitude, from \sim 0.5 years in the tropics to 1–1.5 years in the extratropics. The extensive in situ measurements of key long-lived trace gases CO_2 , SF_6 , N_2O and CH_4 in the 1990s

Figure 7. Mean age vs. normalized N₂O for aircraft (a) and balloon (b) measurements from the 1990s. The results from this study are shown in blue, those based on previously published $CO₂$ are in orange and those from SF₆ are in green. The aircraft measurements are from the four ER-2 missions shown in Fig. [4a](#page-8-0). The uncertainties represent the standard deviation of the mean ages within each N_2O bin. The functional form from Andrews et al. (2001) translated to normalized N_2O is shown as the solid orange line in both plots.

Figure 8. Equivalent latitude vs. altitude distributions of mean ages (a), age spectra ratios of moments (b) and surface source latitudes (c) from the CO_2 optimizations averaged over the 1990s aircraft missions. The grid spacing is 0.5 km in altitude and 5° in latitude.

provides a nearly complete view of the Γ_0 distributions at that time.

The distribution of R_0 shown in Fig. [8b](#page-11-1) is a unique contribution of this study. The R_0 distribution mirrors that of Γ_0 with the smallest values near 1 year in the troposphere and tropical stratosphere, increasing to values > 1.5 years at high latitudes and altitudes. The only other previously published distributions of R come from model output. Hall and Plumb (1994) found values from 0.4–0.7 years throughout the stratosphere from global climate model output at that time. Based on this study, Engel et al. (2008) used a value of $R = 0.7 \pm 0.5$ years in their analysis of Γ trends from balloon measurements of $CO₂$ and $SF₆$. In contrast, Volk et al. (1997) used a value of $R = 1.25 \pm 0.5$ years to calculate mean ages from SF_6 measurements based on the model output from Waugh et al. (1997). Thus, the value of R has varied widely among models and has essentially been unconstrained by observations. More recent modeling studies have revealed the sensitivity of R to the long age tail of G (e.g., Ploeger and Birner, 2016) and have found values of R up to 1.7 years in the extratropical lower stratosphere with appropriate extension of the age spectra tail (Fritsch et al., 2020). The magnitude and distribution of the R values shown in Fig. [8b](#page-11-1) compare well to those in Fritsch et al. (2020, Fig. [6b](#page-10-0)) and provide some of the first extensive validation of modeled values of R throughout the lower stratosphere.

The surface source latitude distribution range in Fig. [8c](#page-11-1) is also a unique result, with Ray et al. (2022) being the only known previous study to have calculated this quantity from trace gas observations but over a more limited time and location. The latitudes shown in the figure are averages of y_0 scaled by $\sum g_V$, which could be considered an extratropical fraction of air. We scale y_o this way to show where the extratropical surface source latitudes matter in the optimization, which is primarily for locations where $\Gamma < \sim 3$ years. For Γ older than this, the extratropical surface contribution to the optimization is negligible, while the source latitude is the Equator. The most notable feature in the source latitude distribution is the hemispheric symmetry with surface latitudes from the same hemisphere as the location of χ_i^* except for in the tropical stratosphere. There will of course be seasonal variability in this quantity, which we will not describe further here. This quantity could also be used to compare to surface source region attributions in the stratosphere from model output (e.g., Yan et al., 2021).

Distributions of the three transport quantities from the optimizations based on OMS balloon measurements in the 1990s (Fig. [9\)](#page-13-0) show the more limited sampling, primarily in latitude, but more extensive vertical coverage compared to the aircraft. The vertical scale on these plots is potential temperature rather than altitude since this emphasizes the stratospheric overworld more than the tropopause region. The balloon measurements extend up to 900 K (34 km altitude), much higher than the \sim 550 K (22 km) level that can be reached by aircraft. No other in situ platform can fly that high, which is of particular value in the tropics.

The OMS balloon Γ_0 and R_0 values generally agree with those from the aircraft measurements in similar regions during this time period. In the tropics above 600 K, Γ_0 and R_0 are nearly constant at values of \sim 3 and 1.5 years. In the northern extratropics, Γ_0 reaches 5 years at 900 K and 40° N and 6 years in a deep region of the polar vortex. Values of R_0 peak at ∼ 1.7 years and are nearly constant above 600 K throughout the extratropics. This is in contrast to model estimates that have peak values in the extratropical lower stratosphere and decreasing values up to 10 hPa (Fritsch et al., 2020). The surface source latitudes have similar features to those from the aircraft measurements in the NH. The tropical lower stratosphere has a layered structure of alternating NH and SH influence (Fig. [9c](#page-13-0)). There is some indication of this in the aircraft data as well, but the averaging over all seasons shown in Fig. [8c](#page-11-1) has reduced the amplitude.

Similar to the comparisons in Fig. [7,](#page-11-0) we can also compare the differences in the spatial distributions of Γ_0 with $\Gamma_{\rm SF_6}$ and Γ_{CO_2} (Fig. [10\)](#page-13-1). In the comparisons with Γ_{CO_2} (Fig. [10a](#page-13-1)), most locations have a difference of -0.3 to -0.5 years, which represents the transport time between a source region at the tropopause for Γ_{CO_2} and the surface for Γ_{o} . There are some locations with larger differences of −1 year, especially in the SH. The region of these larger age differences in the SH coincides with where Γ_0 values are ∼ 1–3 years and the surface source latitudes have SH influence (Fig. [8\)](#page-11-1). The most negative differences were from measurements taken in October (not shown), which has a similar latitudinal dependence in χ_{CO_2} as in November (Fig. [2b](#page-4-0)). At this time of year, a SH surface source results in larger χ_{CO_2} values compared to the Mauna Loa–Samoa average and thus a relatively old Γ compared to assuming only a tropical source.

The $\Gamma_{\rm SF_6}$ values are only younger than $\Gamma_{\rm o}$ in the tropics and southern midlatitudes (Fig. [10b](#page-13-1)). The tropics have the youngest mean ages and thus the smallest correction for SF_6 ; thus, as with $CO₂$, the archived mean ages will be younger in this region due to the tropopause vs. surface boundary conditions. In the northern extratropics the mesospheric loss of SF⁶ and subsequent old age bias overcomes the 3–6-month negative age bias due to the tropopause vs. surface source region difference, resulting in positive mean age differences of up to 1.5 years. This distribution of old age bias generally follows that from the model output of Garny et al. (2024), which was the source of the age bias estimates used in our optimization.

4.2 Results for the 2020s

Only in recent years have simultaneous measurements in the stratospheric overworld of the four trace gases utilized in the optimization become available again. This is an unfortunate gap in the monitoring of the stratospheric circulation over the past 2 decades. With the DCOTSS and SABRE aircraft

Figure 9. Similar to Fig. [7](#page-11-0) but for OMS balloon CO₂ optimizations averaged over the 1990s. Here, the vertical coordinate is potential temperature. Note that the mean age range in (a) and the latitude scale have been changed from Fig. [8](#page-11-1) to better match the balloon sampling and that the theta scale is cut off for the surface source latitudes. Grid spacing is $10 K$ in potential temperature and 5° in latitude.

Figure 10. Gridded average differences between the archived aircraft $CO₂$ (a) and $SF₆$ (b) mean ages minus the optimized mean ages in this study. The optimized mean ages are from the single trace gas technique with $CO₂$ (a) and $SF₆$ (b) in order to maximize the number of comparison data points for each trace gas. Grid spacing is the same as in Fig. [8.](#page-11-1)

missions and AirCore balloon flights, we now have the ability to calculate age of air from in situ measurements in the current stratosphere. The new measurements were the inspiration for this study since a consistent technique is necessary to compare the old and new measurements and age of air. Here we briefly describe the derived mean ages, which have not previously been published from this data, and other transport quantities from the three measurement campaigns in the 2020s. A follow-up study will examine the differences between the 1990s and 2020s ages in detail. Briefly, the main differences are younger mean ages at nearly every χ_{N_2On} value for the 2020s compared to the 1990s, especially when

Figure 11. Distributions of mean age vs. normalized N_2O from DCOTSS (blue), which took place during the summer; SABRE (green), which took place during the winter; and AirCore (red for summer, purple for winter) measurements. The latitude and potential temperature sampled ranges of each mission are shown in Fig. [12.](#page-15-0)

comparing the aircraft mission data. The younger mean ages are highly suggestive of a strengthened Brewer–Dobson circulation (BDC) over the past 2 decades.

The distributions of $\Gamma_0 - \chi_{N_2On}$ for the three different mission data sets are compact and agree well on average, with slightly older Γ_0 from AirCore primarily from the winter sampling (Fig. [11\)](#page-14-0). The summer average AirCore mean ages agree within uncertainties with the DCOTSS mission, which took place during the NH summer. The distribution of Γ_0 in latitude and theta shows the different sampling between the three missions (Fig. [12\)](#page-15-0). DCOTSS used the ER-2 and focused on the summer NH stratosphere, which resulted in sampling of relatively young Γ_0 up to nearly 500 K and at all latitudes. SABRE used the WB-57 and took place in NH winter with a focus on the polar region where relatively old Γ_0 were sampled, although only SF₆ was measured during this mission. AirCore uses balloons flown predominantly from Boulder, CO, but in various seasons, with the oldest Γ_0 above 500 K in the midlatitudes. Distributions of R_0 and surface source latitudes for each mission are shown in Figs. S12 and S13 in the Supplement.

NOAA's AirCore program is ongoing and is likely to expand to different latitudes as new recovery techniques for this system are developed, but regardless it is a valuable addition to the monitoring of the stratospheric circulation. It is unclear what future aircraft missions will occur with the requisite payload to use this optimization technique and calculate age of air. However, it is unlikely that any period of aircraft sampling of the stratosphere from pole to pole as in the 1990s will occur again.

5 Conclusions

In this study we describe and demonstrate a new technique to calculate the first and second moments of the age-of-air distribution and surface source latitudes based on aircraft and balloon in situ measurements of SF_6 , CO_2 , CH_4 and N_2O throughout the upper troposphere, lowermost stratosphere and stratospheric overworld. This technique depends crucially on the recent quantification of the time-dependent bias in $SF₆$ -derived mean age due to mesospheric loss by Garny et al. (2024) and the NOAA Greenhouse Gas MBL Reference product (Lan et al., 2023) as a time-varying and latitudinally varying surface source for each of the trace gases mentioned above. The main advantages of this new technique are as follows.

- 1. The use of multiple trace gases within this method allows for a single, optimized age-of-air estimate to be found for each measurement location while minimizing the impact of various uncertainties associated with age of air from single trace gases.
- 2. The method allows for the calculation of age of air from the upper troposphere through the stratosphere.
- 3. The method provides the flexibility to utilize measurements from different instruments and platforms across decades and account for calibration offsets.

The technique we use is based on the convolution method of estimating age of air and builds on studies such as Bönisch et al. (2009), Ray et al. (2017), Leedham Elvidge et al. (2018), and Ray et al. (2022). One of the primary aspects of how we define the age of air is that it is from the Earth's surface rather than the tropopause as has been commonly used in previous studies. The main reason we choose to use the Earth's surface is that this is where we have long-term measurements of the trace gases used here, which is necessary for any age-of-air calculation. Of course, all age-of-air studies have used surface measurements to generate a source region time series for the trace gas(es) of interest, but for those that used the tropopause as a source region, assumptions about how surface trace gas time series translate to the tropopause need to be made.

There are benefits and complications to the choice of the surface as the age-of-air source region. Complications include (1) significant latitudinal gradients in most trace gases that are in growth, which makes defining a single source region time series difficult, and (2) conceptualizing what age of air from the surface means compared to the more traditional age since an air parcel passes through the tropopause. We address the first complication by partitioning the surface source region time series of each trace gas into two parts, one averaged over all tropical latitudes similar to the standard Mauna Loa–Samoa average (e.g., Boering et al., 1996; Andrews et al., 2001) and a second part averaged over 10°

Figure 12. Equivalent latitude vs. potential temperature distributions of mean age from the DCOTSS (a), SABRE (b) and AirCore (c) missions. Note that the mean age scale and latitude range are different from Figs. [8](#page-11-1) and [9](#page-13-0) to emphasize the sampled region during these missions. Grid spacing is the same as in Fig. [9.](#page-13-0)

latitude bins from 60° S–60° N that is a free parameter in the calculation (e.g., Ray et al., 2022). The latitudinally varying surface time series is only considered as a source region for the youngest ages of air, on the order of months, primarily because for older ages the measured trace gases lose sensitivity to the surface latitudinal gradients due to mixing. In other words, unlike an idealized tracer in a model that can be emitted and tagged from a certain surface latitude and followed throughout any trajectory and timescale in the atmosphere (e.g., Orbe et al., 2015), real trace gas measurements can only be identified as having been emitted from a certain surface latitude range for a limited time. Conceptually, this means most of the ages in the age spectra in our calculation are from a tropical average surface source similar to that from almost every previous age-of-air study but without any assumed constant transport time from the surface to the tropopause. The youngest ages in the age spectra have a flexible surface source latitude that allows us to utilize any latitudinal emission information that persists in the measured trace gases.

Thus, the complications from the use of the surface as a source region can be turned into a benefit in that additional transport information can be obtained from the trace gas measurements. This benefit is only realized with the use

of multiple trace gases since any single trace gas can have too many ambiguities in surface sources combined with transport timescales to constrain the latitudinal source region. The more in situ trace gas measurements are available, the more potential there is to constrain the surface source region (e.g., Ray et al., 2022). However, most aircraft and balloon in situ payloads have not included instruments that measured more than the four age-of-air trace gases considered here, and often only one or two were measured.

Further benefits of a surface source region include consistency with age of air from CCM output and the ability to calculate age of air at any location in the atmosphere. This last benefit applies primarily to locations in the upper troposphere and above since the shortest transport timescales in the troposphere (days to weeks) require a range of shorterlived trace gases to define (e.g., Luo et al., 2018). A single technique to calculate age of air in the upper troposphere, tropopause region, lowermost stratosphere and stratospheric overworld can provide new insight into transport in these regions and connect various airborne measurements with different sampling limitations.

A further unique aspect of this study is the offset ensemble as part of the optimization procedure. The assumption behind this technique is that there are many uncertainties in calculating the age of air that are not accounted for in the measurement errors alone. The additional sources of uncertainty can be identified and added together to create a total uncertainty that can be quite large (e.g., Engel et al., 2008). But the offsetting of individual trace gases is a method to identify systematic errors that can often be attributed to a specific issue, such as the $SF₆$ calibration during the SOLVE mission in the 1990s. The offset optimization accounts for the calibration issue while still calculating mean ages that agree well with values from missions throughout the 1990s. Without the offset method, measurements from SOLVE, for instance, would not be usable in the optimization because there would be no overlap between solutions for SF_6 and CO_2 . For other missions and locations there may be a minimal set of overlapping solutions due to an inaccurate assumption of photochemical loss for instance. As long as one of the trace gases, in this case $CO₂$, requires negligible scaling across essentially all the available measurements, the scaling method can identify and account for systematic errors in the measurements or certain aspects of the age calculation. It may be asked why $SF₆$ is used at all from the 1990s if the offset values can be large. The benefits of adding SF_6 are many, but primarily it constrains the younger mean ages (Γ < 3 years) that are not well constrained by $CO₂$, as shown by the example in Figs. [3](#page-6-0) and S8, which then allows better constraint on the ratio of moments.

We demonstrate the technique on the extensive in situ measurements from the 1990s and the more limited recent measurements in the 2020s. The 1990s average distributions of mean ages, age spectra width and surface source latitudes from the upper troposphere to the lower stratosphere and nearly pole to pole provide a unique view of average transport at that time. The optimized mean ages calculated here are compared to the archived values previously calculated from $CO₂$ and $SF₆$ measurements independently. In many regions the optimized mean ages are similar to the archived values from $CO₂$, accounting for the surface vs. tropopause source difference. An additional factor that could contribute to older mean ages is the variable surface source latitudes that often result in larger boundary condition $CO₂$ mixing ratios compared to the Mauna Loa–Samoa average. The optimized mean ages are younger than the previously calculated values from SF_6 everywhere but in the tropics and SH due to the SF_6 mesospheric loss and mean age bias that was not previously accounted for.

The age spectra ratios of moments and surface source latitude distributions from both time periods are unique contributions from this study. The ratios of moments have not been well constrained by observations, and the distributions shown here generally agree well with recently published values from CCM output. The surface source latitudes have a number of interesting features that are beyond the scope of this study to discuss but demonstrate the potential information available from this technique. We also leave comparisons of the results between the two time periods to a followup study.

The recent aircraft missions and more systematic balloon flights in the stratosphere provided the primary motivation for this study, yet these new data are limited in their ability to detect decadal changes in the stratospheric circulation due to the NH-only sampling. It has been known for quite some time that the average difference in mean age between the tropics and extratropics on an isentrope can reveal the strength of the vertical component of the BDC (Neu and Plumb, 1999; Linz et al., 2017). Yet, the last in situ measurements in the tropics above 550 K of the four trace gases used in this study to calculate mean ages were made in the 1990s. The result of this lack of tropical measurements is that we have to rely on mean age trends from the somewhat more extensive in situ measurements in the NH to infer multi-decadal trends in the BDC (e.g., Engel et al., 2008, 2017). This single-hemisphere view is an incomplete picture, partly due to hemispheric shifts in the BDC that can occur on decadal timescales (e.g., Ploeger and Garny, 2022). Satellite measurements of long-lived trace gases can be used to calculate age of air (e.g., Haenel et al., 2015), but thus far no satellite has measured two age tracers such as $CO₂$ and $SF₆$ necessary to perform the optimization described here. Thus, in situ measurements provide an important complement to satellite measurements, and more consistent in situ sampling of the tropical and extratropical stratosphere (e.g., Moore et al., 2014) is necessary to better understand and model the ongoing stratospheric circulation changes.

Code availability. The IDL software used to perform the data analysis and make the figures in this study is available from [https://csl.noaa.gov/groups/csl8/modeldata/data/Ray_etal_2024/](https://csl.noaa.gov/groups/csl8/modeldata/data/Ray_etal_2024/IDL%20Code/) [IDL%20Code/](https://csl.noaa.gov/groups/csl8/modeldata/data/Ray_etal_2024/IDL%20Code/) (Ray, 2024a).

Data availability. The processed data supporting this study are available from [https://csl.noaa.gov/groups/csl8/modeldata/data/](https://csl.noaa.gov/groups/csl8/modeldata/data/Ray_etal_2024/Data/) [Ray_etal_2024/Data/](https://csl.noaa.gov/groups/csl8/modeldata/data/Ray_etal_2024/Data/) (Ray, 2024b).

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