Scintillation Index Analysis of ship-borne GNSS Data recorded during the MOSAiC Expedition in the Arctic

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URSI AT-RASC, Gran Canaria, May 2024

Outline



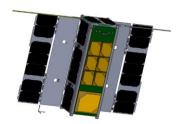
- GNSS Remote Sensing for Ionospheric Monitoring
- MOSAiC Expedition and GNSS Data in the Arctic
- Processing and Masking of Ship-based Data
- Results of Scintillation Index Analysis
- Conclusions



Motivation GNSS Remote Sensing

A: Low Earth Orbiter

Wickert et al. 2016 Semmling et al. 2016



B: Aircraft

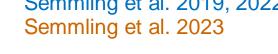
Semmling et al. 2014 Moreno et al. 2021



C: Research Vessels



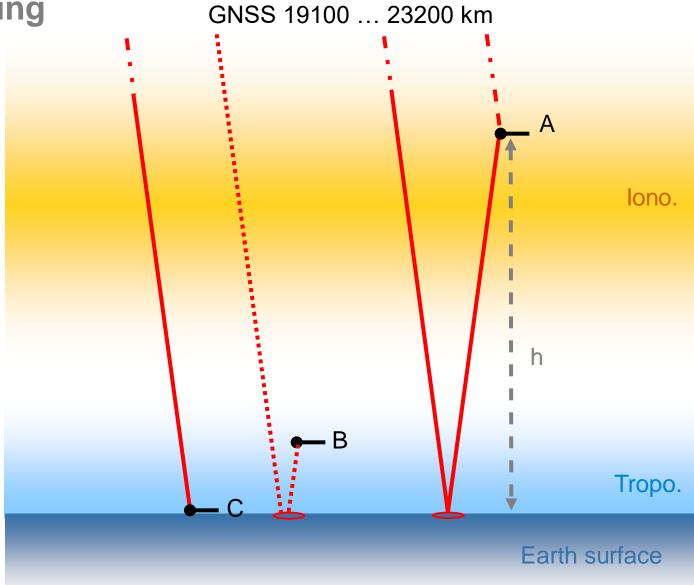
Semmling et al. 2019, 2022





sea surface altimetry sea state estimation sea-ice detection

water vapor estimation iono. scintillation detection

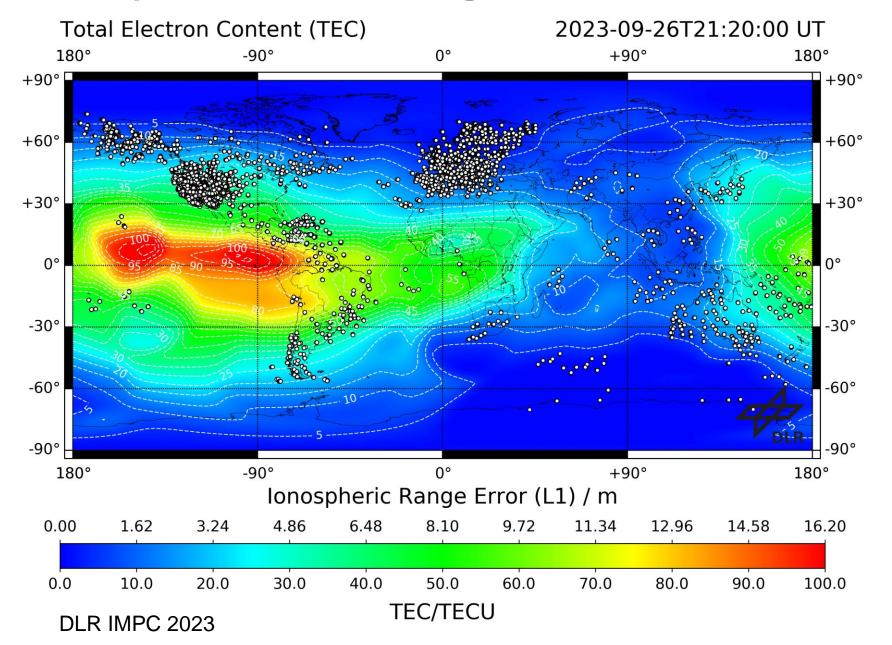


A: e.g. PRETTY, $h \sim 500 \text{ km}$

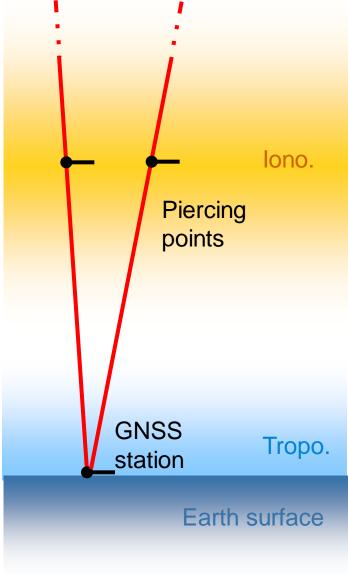
C: e.g. Polarstern, h ~ 25 m

B: e.g. HALO, h ~ 3500 m

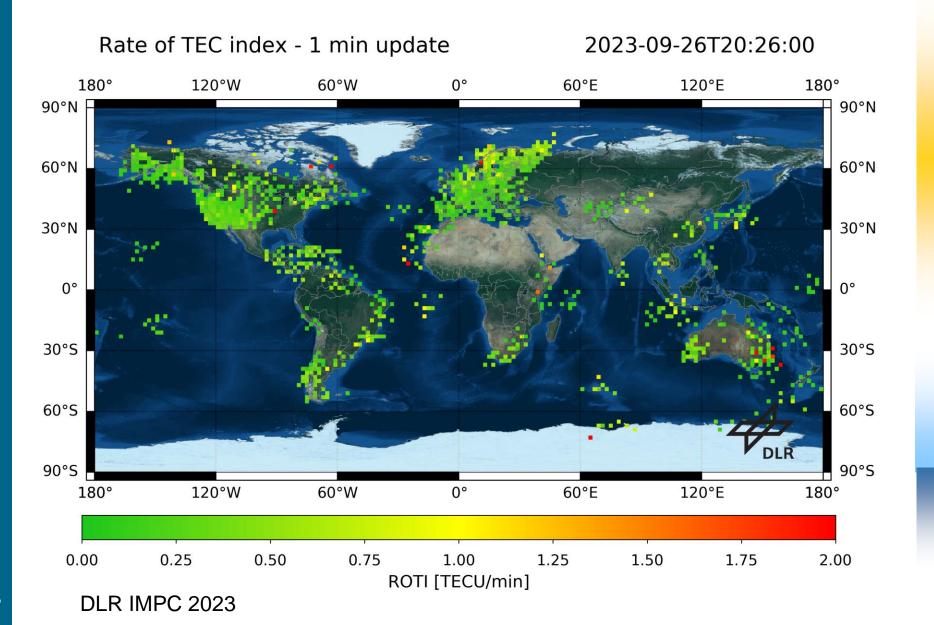
Ionosphere TEC Monitoring with GNSS



GNSS 19100 ... 23200 km



Ionosphere Disturbance Monitoring with GNSS



GNSS 19100 ... 23200 km lono. Piercing points **GNSS** Tropo. station

Earth surface

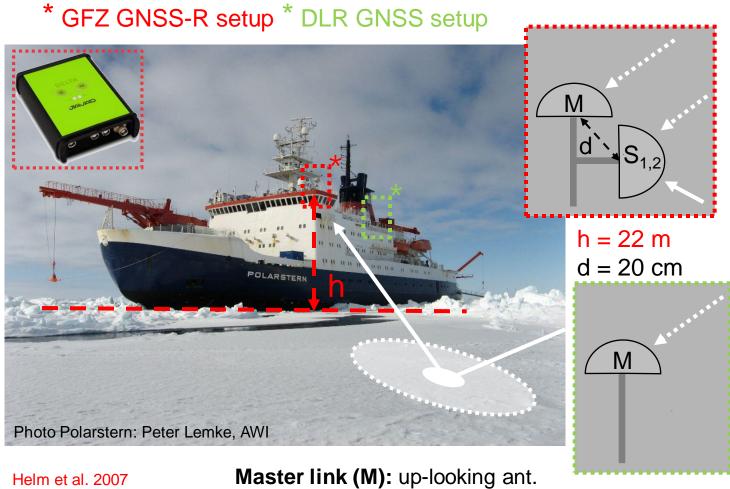


Can we benefit from ship-based data?

MOSAiC Expedition and GNSS Data in the Arctic

MOSAiC Expedition and Polarstern Setup





Slave links $(S_{1,2})$: side-looking ant.

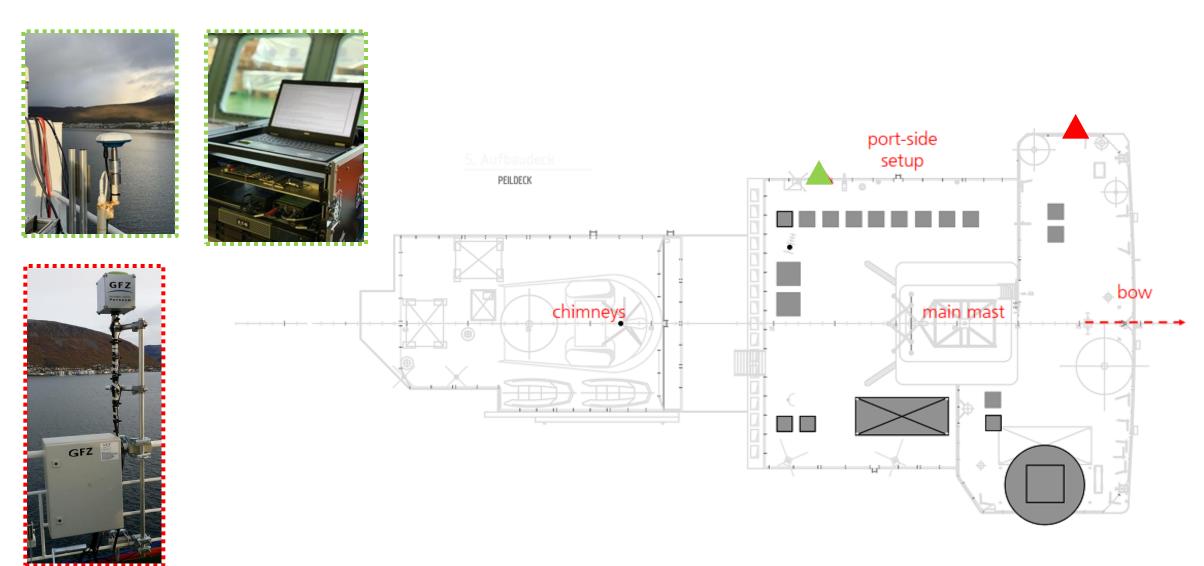
MOSAiC expedition: Sep 2019 - Sep 2020 2019 60°W 30°E

> **Cruising Periods:** speed > 1 m/s **Drifting Period:** speed < 1 m/s

Semmling et al. 2013 Kriegel et al. 2017

MOSAiC Expedition and Polarstern Setup



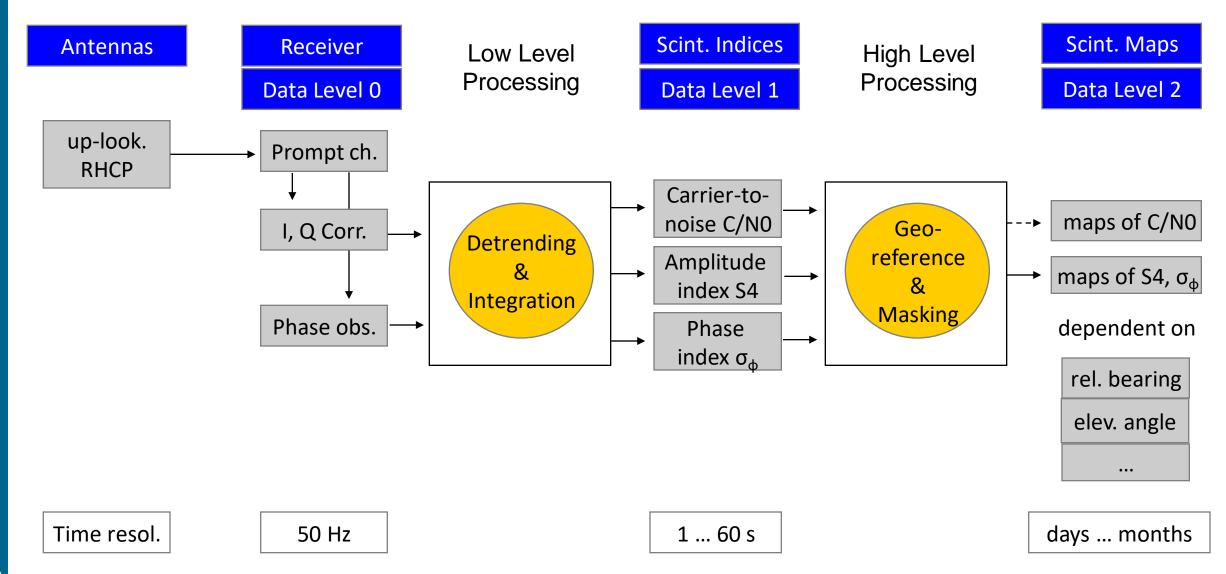




Processing and Masking of Ship-based Data

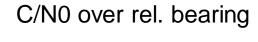
High-rate GNSS Data Processing

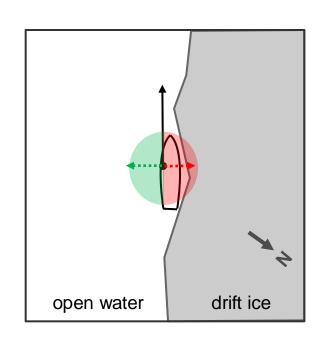




Limits of Visibility from the Ship

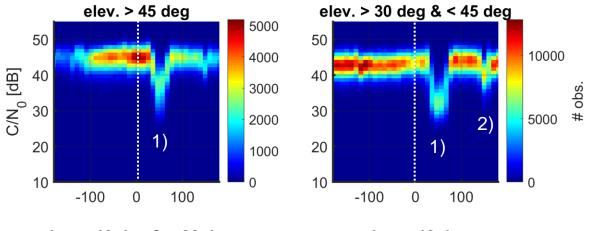


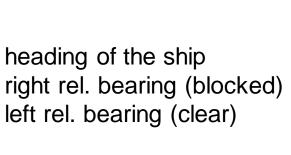


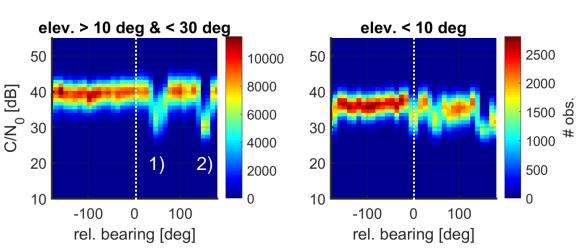


clear view to port-side

left rel. Bearing: -180° to 0°







- 1) ship's main mast
- 2) ship's chimney

Sep 2019 ... Sep 2020

Semmling et al. 2023

4.....

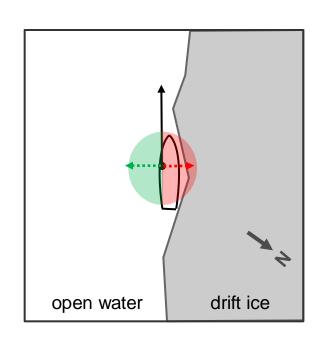
Limits of Visibility from the Ship



1500

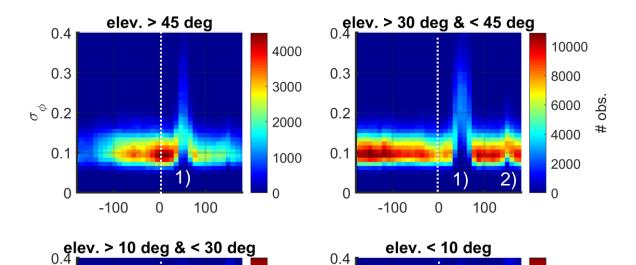
1000 sqo #

500



clear view to port-side

left rel. Bearing: -180° to 0°



0.4

0.3

0.2

-100

0

rel. bearing [deg]

100

8000

6000

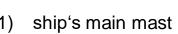
4000

2000

 σ_{ω} over rel. bearing



heading of the ship right rel. bearing (blocked) left rel. bearing (clear)



0.3

-100

rel. bearing [deg]

100

ნ 0.2

ship's chimney



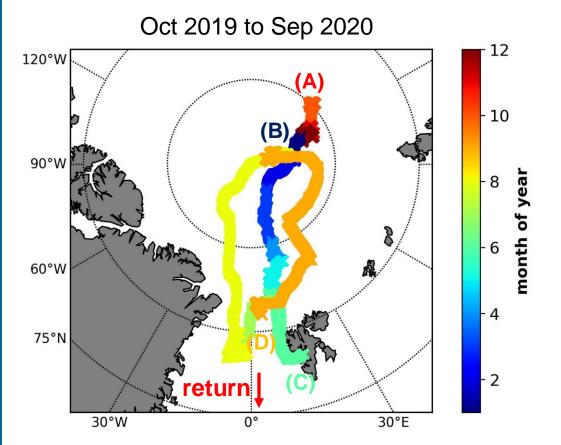


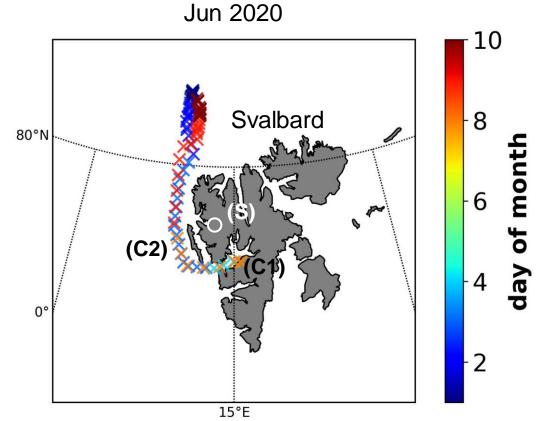
Results of Scintillation Index Analysis

How does a ship-based setup perform compared to a station setup?

Track of R/V Polarstern (PS)







(A) Oct 1, 2019

(B) Jan 15, 2020

(C) Jun 5-8, 2020

(D) Sep 30, 2020

before drift in ice near Svalbard before return

(C1) noon Jun 4 to afternoon Jun 8(C2) night Jun 8

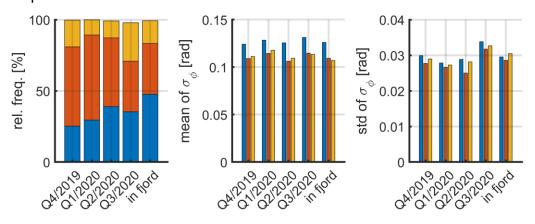
calm sea, inside fjord high sea state, outside fjord

(S) Ny-Alesund station operated by Univ. of Oslo

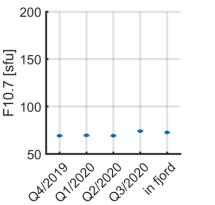
σ_{ω} statistics comparing PS with station data



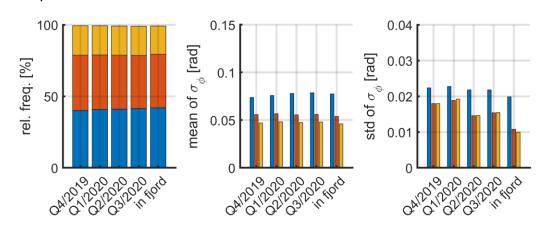
 σ_{ω} on PS during MOSAiC – all weak (< 0.3 rad)



Solar radio flux – permanently low



σ_{ϕ} at Ny-Alesund station – all weak (< 0.3 rad)



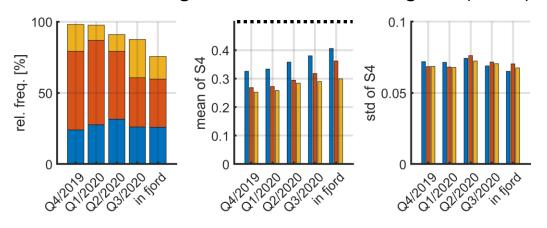
Remarks:

- In general, all σ_{ϕ} on PS and at the station are classified weak scint.
- σ_{ϕ} on average significantly higher on PS than at the station (0.05 rad to 0.1 rad)
- almost no changes over the year

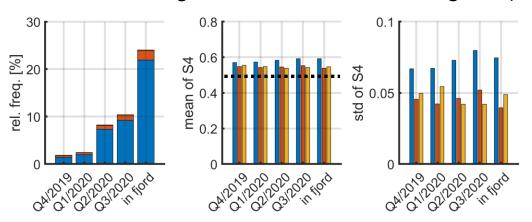
S4 statistics comparing PS with station data



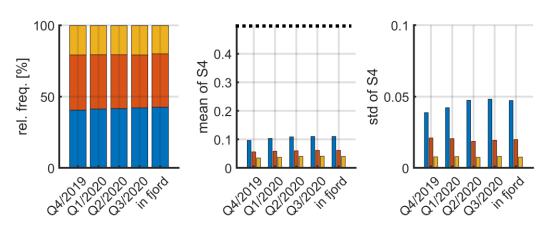
S4 on PS during MOSAiC – weak regime (< 0.5)



S4 on PS during MOSAiC – moderate regime (>= 0.5)



S4 at Ny-Alesund station – all weak (< 0.5)



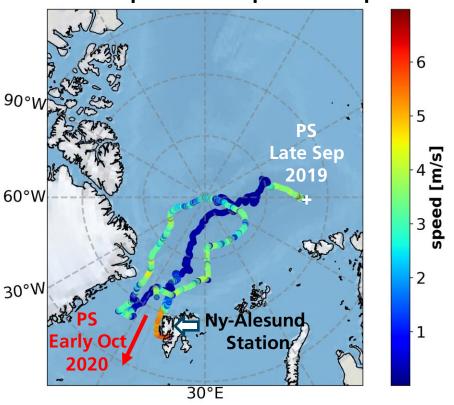
Remarks:

- weak part dominating (> 75%) of S4 for PS
- moderate part for PS mainly at low elev. (10°-30°), later in 2020 (Q2, Q3) and in fjord period
- S4 on average significantly higher for PS (0.25 ...
 0.55 rad) than for station (0.05)
- reason for increased S4 at PS?

Movement of Polarstern

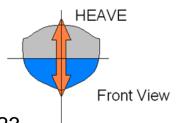




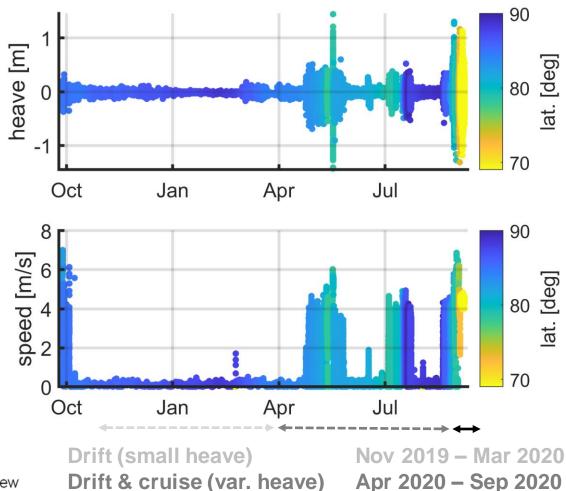


Cruising Periods: speed > 1 m/s

Drifting Period: speed < 1 m/s



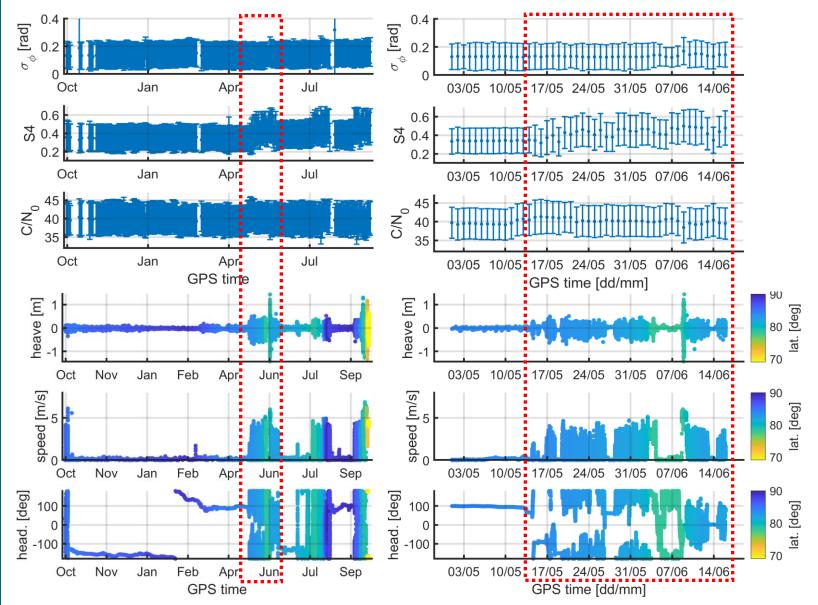
MOSAiC expedition: Polarstern 2019/2020



Wikipedia 2023 (Public Domain)

PS results in relation to ship's movement

Ship near Svalbard





	Jun 6	Jun 7	Jun 8	Jun 9
av. σ_{ϕ} [rad]	0.14	0.13	0.14	0.17
av. S4	0.41	0.48	0.47	0.50
av. C/N0 [dB]	40.0	40.6	40.8	38.4
max. heave [m]	< 0.2	< 0.2	< 0.2	> 1.0

Remarks:

- night June 8-9 ship leaves fjord
- high sea state outside leads to increased heave of the ship
- scint. indices slightly increased

Summary of Results at High Elevations



	Feb 2020	Mar 2020	Jun 2020	Sep 2020	Oct 2020	Mar 2020	Jun 2020
Days of obs.	28	31	30	30	5	31	30
Av. Speed [m/s]	0.1+-0.1	0.2+-0.1	0.7+-1.2	0.9+-1.3	4.4 +-0.3	-	-
Av. Heave [dm]	-0.2+-0.1	-0.1+-0.4	-0.1+-0.8	0.0+-0.5	-0.5 +-2.3	-	-
High elev. Indices							
Av. S4 (% to ref.)	0.26 (100)	0.25 (95)	0.29 (113)	0.32 (122)	0.31 (120)	0.04 (15)	0.04 (16)
Av. σ(φ) [rad] (%)	0.12 (100)	0.11 (92)	0.11 (93)	0.11 (92)	0.12 (99)	0.05 (41)	0.05 (38)
Av. C/N0 [dBHz]	43+-2	44+-2	44+-3	43+-3	44+-3	51+-1	50+-1

S4 scint. is ...

Weak 0 ... 0.5

Moderate 0.5 ... 0.8

Strong > 0.8

 $\sigma(\phi)$ [rad] scint. is ...

Weak 0 ... 0.4

Moderate 0.4 ... 0.7

Strong > 0.7

Polarstern Setup during MOSAiC

Ny-Alesund Station on Svalbard

Conclusions

- GNSS remote sensing from a ship requires adapted processing (mask out ship structure disturbance)
- Baseline phase noise (σ_{ϕ}) and amplitude fluct. (S4) are increased compared to station obs., however, not to severe level
- Carrier-to-noise ratio (C/N0) is decreased (by 6 ... 7 dB) compared to station obs., it means reference signal intensity is lower on a ship than at the station
- S4 increased (by $\sim 20\%$) in periods of increased heave (std 0.5 ... 2m), σ_{φ} is not
- sensitivity to moderate and strong iono. scintillation expected, more data needed.

Acknowledgements

Support from MOSAiC team
G. Spreen, L. Kaleschke, R. Ricker, A. Tavri
Logistics at AWI & Crew of R/V Polarstern
Werkstatt and IT staff at DLR and GFZ

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Thank you for your attention.

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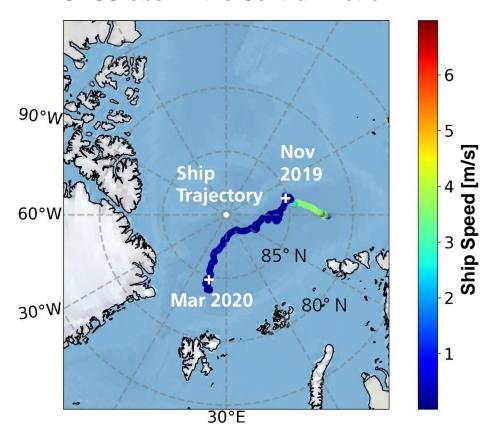


Appendix

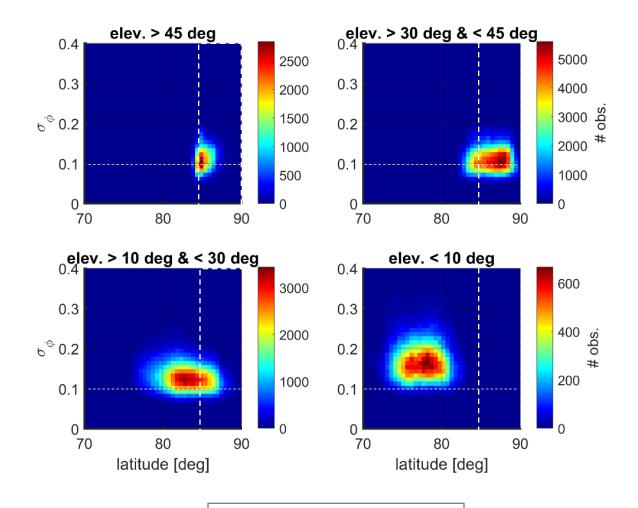
Drift - High Arctic - Winter



GNSS obs. in the Central Arctic



 σ_{ω} over lat. at IPP (height 350 km)



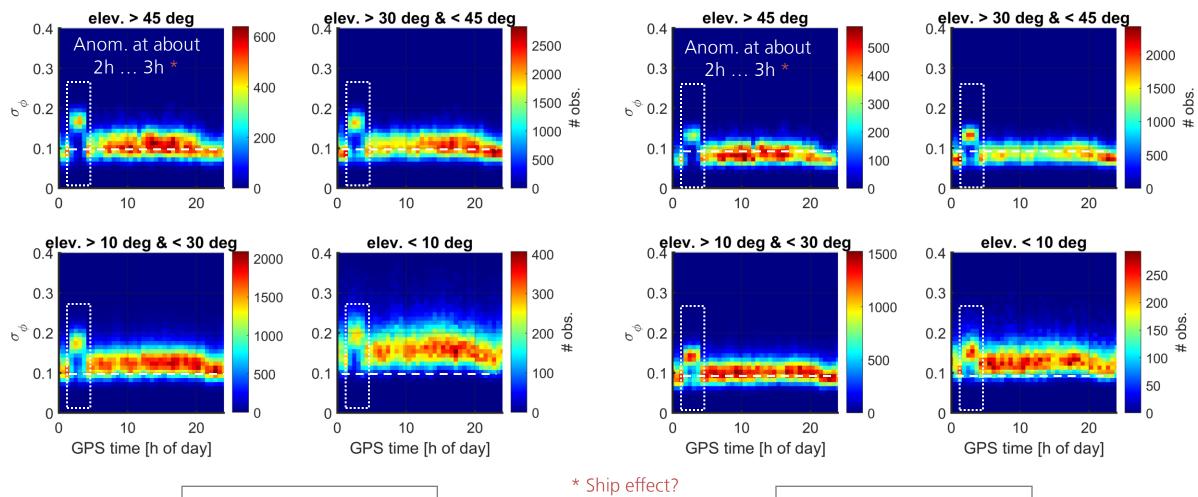
Nov 2019 ... Mar 2020

Drift - High Arctic - Winter









Nov 2019 ... Mar 2020

Nov 2019 ... Mar 2020