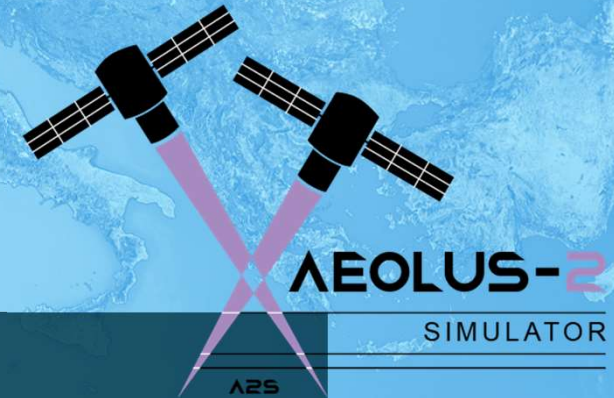


THE WAY TO A PERFORMANCE SIMULATOR FOR EPS-AEOLUS



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Theme: 03. Towards new operational programmes and preparedness studies

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Aeolus wind and aerosol-cloud observations

Objective:

Improve numerical weather prediction (NWP) and advance understanding of atmospheric dynamics and climate processes

Orbit:

polar, sun-synchronous
7 day repeat cycle with
111 orbits \approx 16 orbits / day

Geometry:

Altitude: 308 km
Angle: 35° (off-nadir)

Observations:

\approx 7000 line-of-sight (LOS) wind and aerosol/cloud optical profiles
(\approx 5-6 times more than radiosondes)

Horizontal resolution:

Raw data: 3-4 km
Mie wind: 10-20 km
Rayleigh wind: 90 km

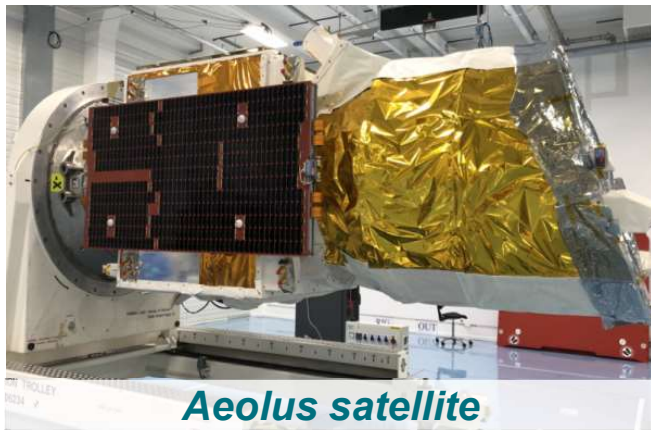
Vertical resolution:

Max. altitude: \approx 30 km
Number of bins: 24
Bin thickness: 0.25 – 2 km

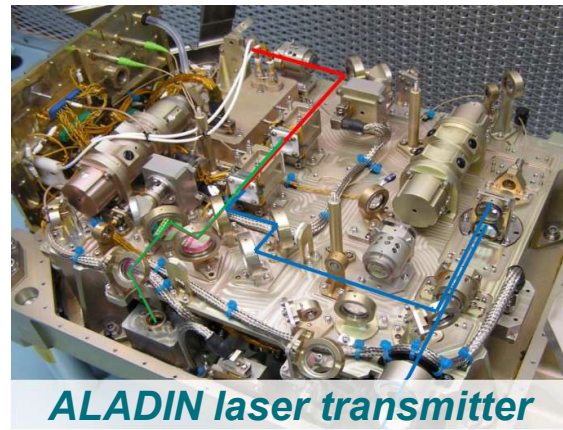
Wind requirements (HLOS)

random error: 1 – 2.5 m/s
systematic error: $<$ 0.7 m/s

ALADIN – a technological challenge



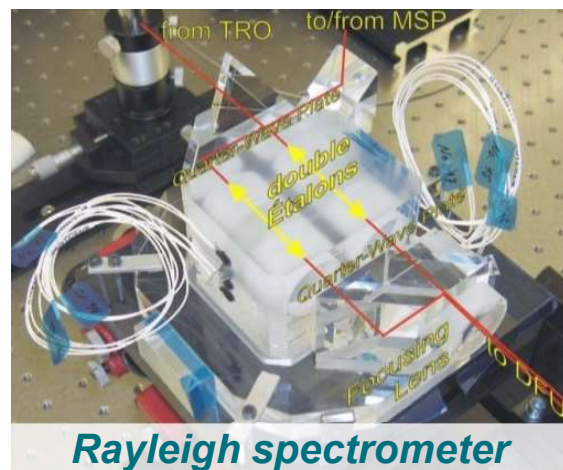
Aeolus satellite



ALADIN laser transmitter



Mie spectrometer



Rayleigh spectrometer

- **Launch** 22 August 2018, **nominal lifetime** of 3.5 years **exceeded** with operation until 5 July 2023, and **assisted re-entry** on 28 July 2023
- **First European lidar in space** after 20 years of development challenges and **first wind lidar in space**
- **First high-power, ultraviolet (UV) laser** in space (@ 354.8 nm) with stringent requirements on frequency stability (<7 MHz rms)
- **Doppler wind lidar principle** – straightforward but incredibly small wavelength shift
- **Challenging direct-detection** approach, due to need for winds from broad-bandwidth molecular Rayleigh backscatter up to lower stratosphere

Doppler equation: $\Delta f / f_0 \approx 10^{-8}$

relative Doppler shift: $\Delta f = 2f_0 \frac{v_{LOS}}{c}$

→ 1 m/s (LOS) ~ 5.64 MHz ~ 2.37 fm

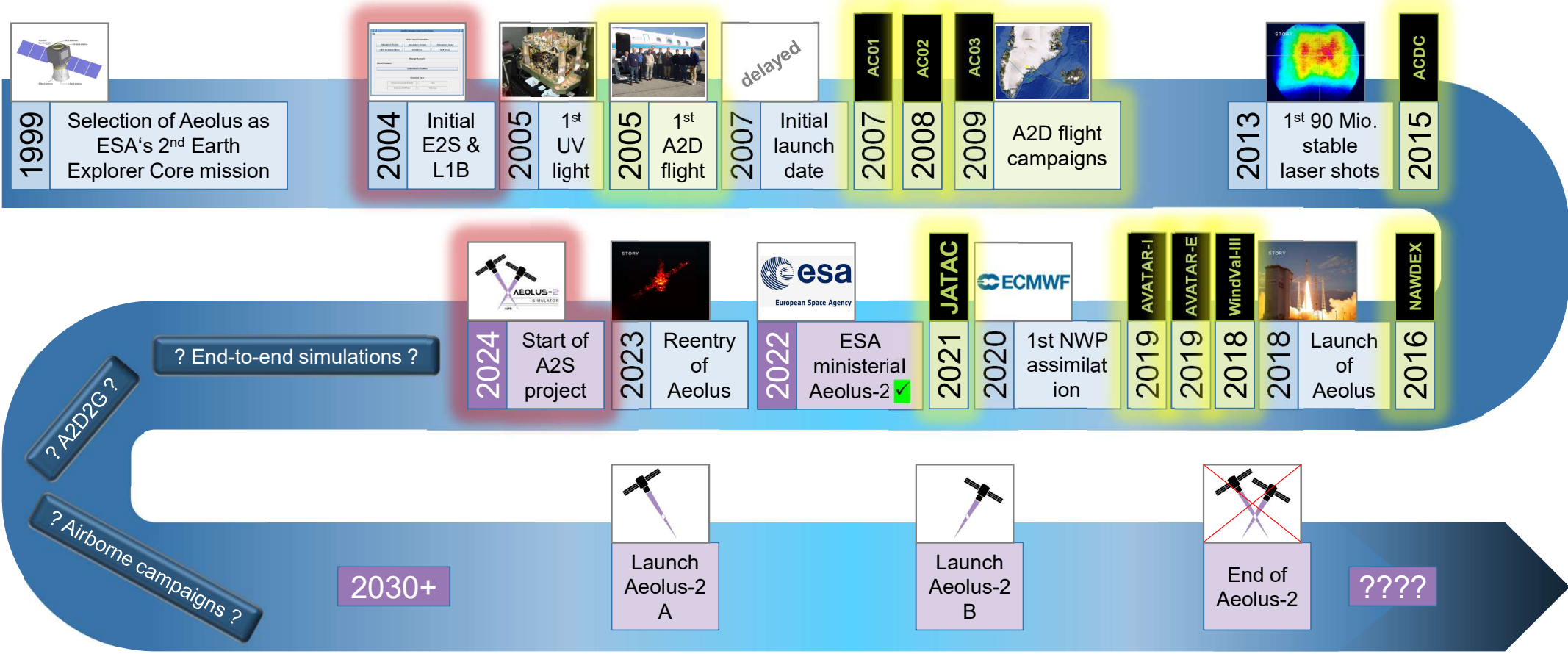
Fig. credits Airbus DS

Aeolus – History and future

... accompanied by the End-to-End Simulator (E2S) & the Aeolus Airborne Demonstrator (A2D)



→ Pre-launch: > 100 recommendations derived for Aeolus instrument alignment, operation, retrieval algorithms and CalVal

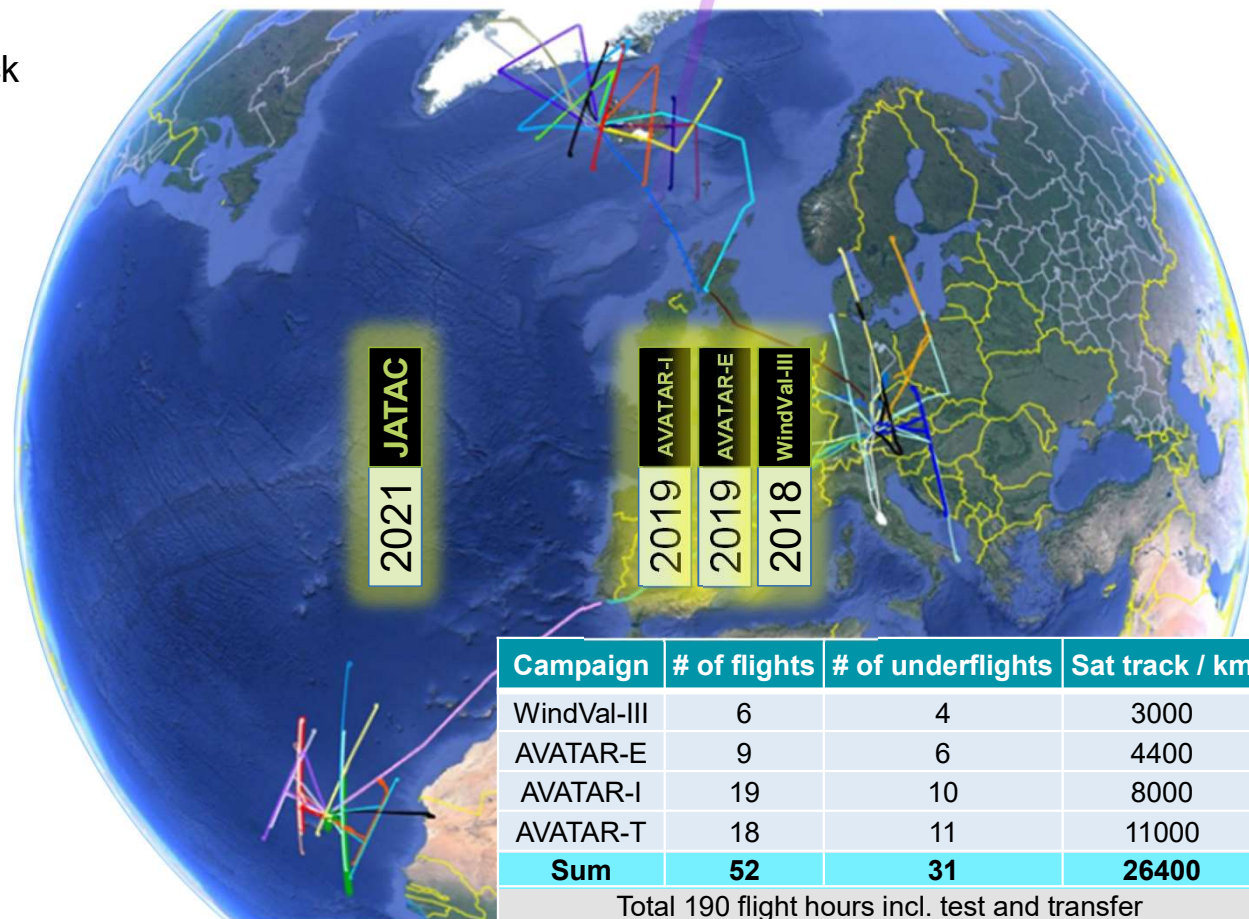
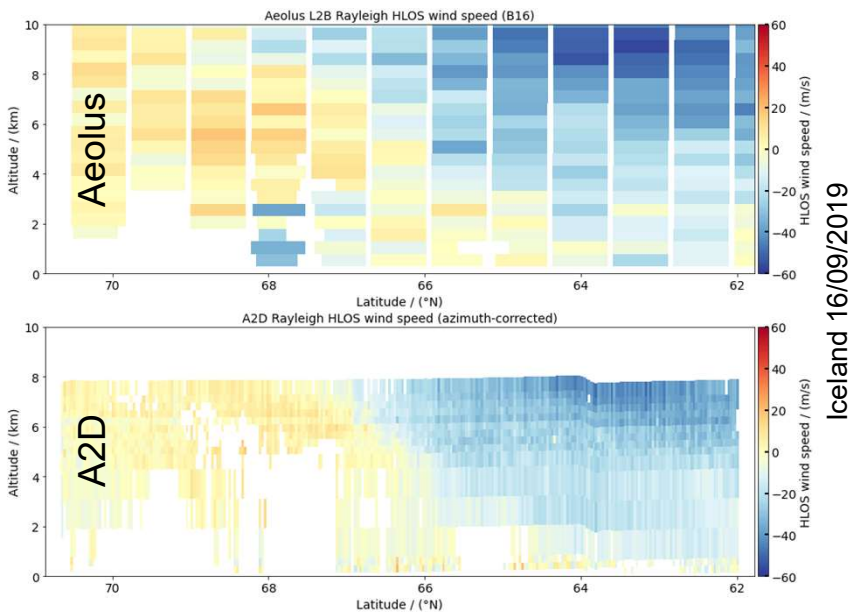


Aeolus Airborne Validation Campaigns After Launch



- 4 airborne campaigns employing the A2D
 - 52 flights with 26400 km along the Aeolus track
- high resolution and high quality data

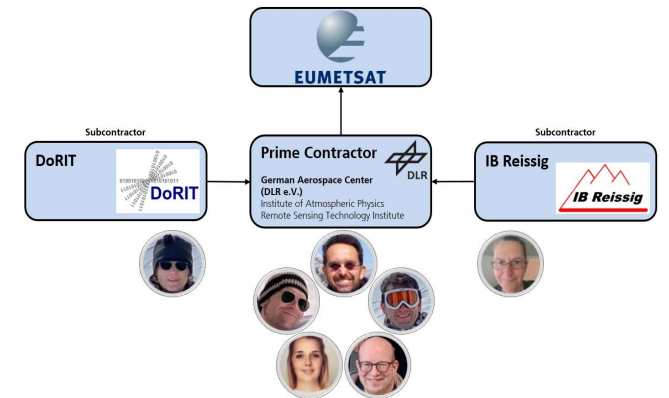
Rayleigh channel



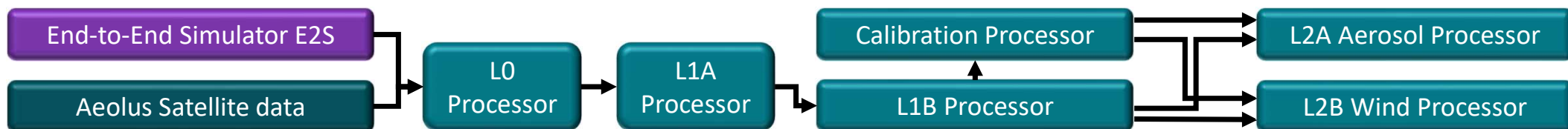
The Aeolus-2 Simulator (A2S) Study



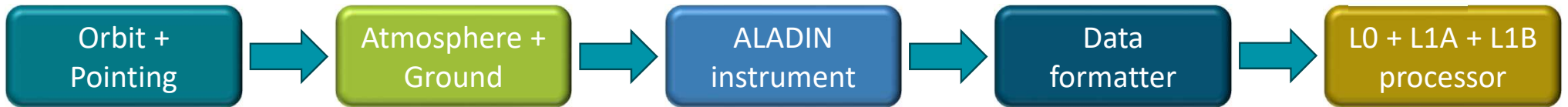
- The **purpose** of this study (Jan. 2024 – Apr. 2025):
 - **Re-configure** the Aeolus End-to-End Simulator (**E2S**)
 - **Assess the radiometric performance** of the Aeolus Rayleigh channel (Mie channel more complicated) by comparing simulated and measured signals
 - Run **simulations** representative of the planned **EPS-Aeolus** mission performance (without updating the E2S software code).
 - Investigation of a Dual **Michelson** interferometer (DMI) to assess the influence of **Mie contamination on the Rayleigh** channel + **accuracy** of the correction



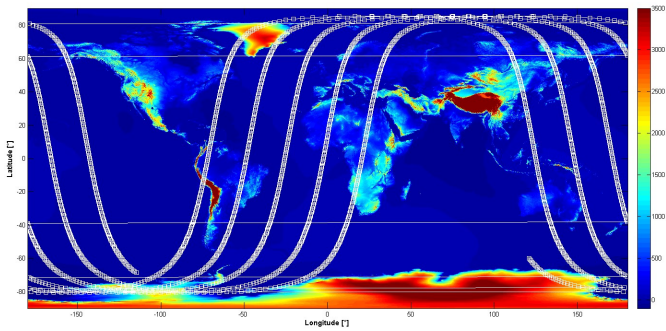
- The great **potential** of the Aeolus End-to-End Simulator (**E2S**):
 - The E2S was **used extensively** in combination with the Aeolus L1B, L2A and L2B processors before launch for performance simulations, algorithm sensitivity studies and functional testing **over 15 years**.
 - Significant work was invested in the E2S to simulate atmosphere and the ALADIN instrument **close to reality** before launch
 - After launch, the E2S was used for **estimation** of initial signal loss (DLR) and the **investigation** of noise in the L2A retrievals (Météo France, ESA-ESTEC) and is now used in the functional **testing** during the processor delivery (DoRIT, KNMI)
 - Functional **updates** of the E2S (compatible with L1B processor) were provided every 6 months during the operational phase of Aeolus.



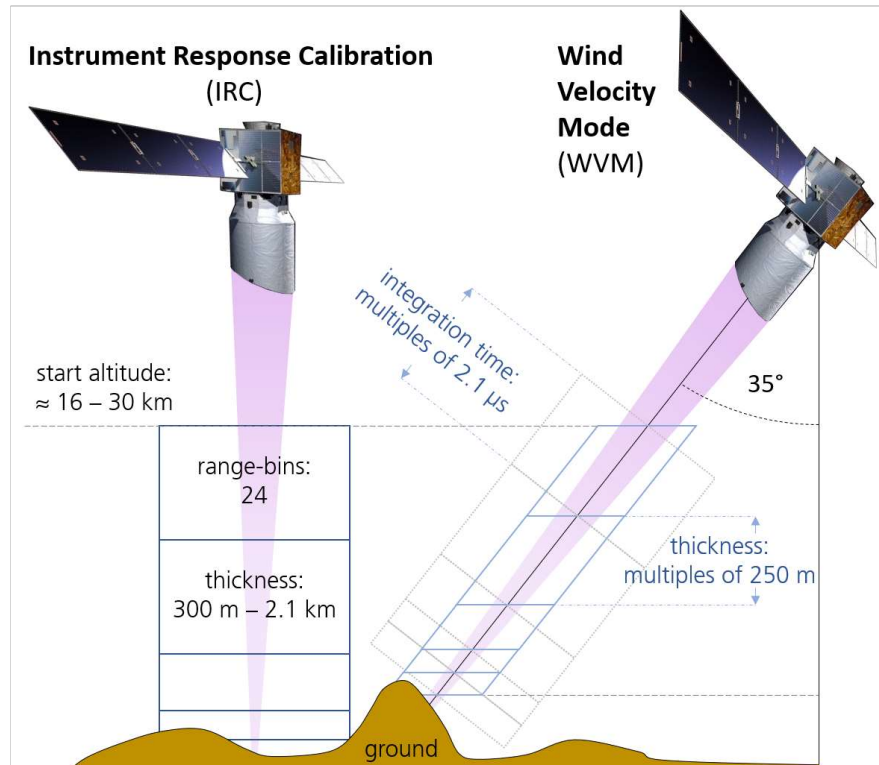
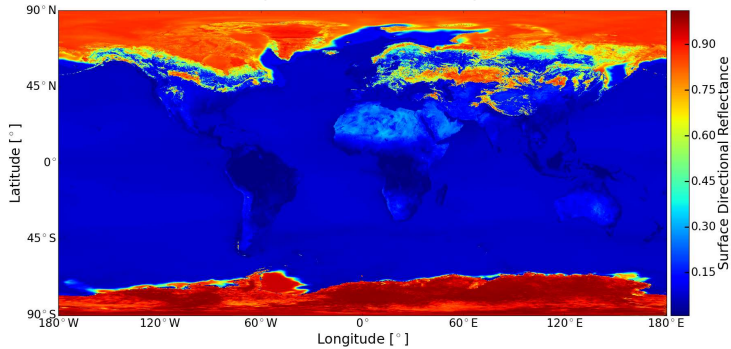
The Aeolus End-to-End Simulator (E2S)



Aeolus orbit and DEM



UV surface reflectance (ADAM database)



E2S parameters (selection)



Atmosphere and Ground

wind	altitudes
temperature	cloud thickness
pressure	cloud backsc. & extinct.
molec. backsc. & extinct.	DEM
aeros. backsc. & extinct.	albedo

Satellite and Structure

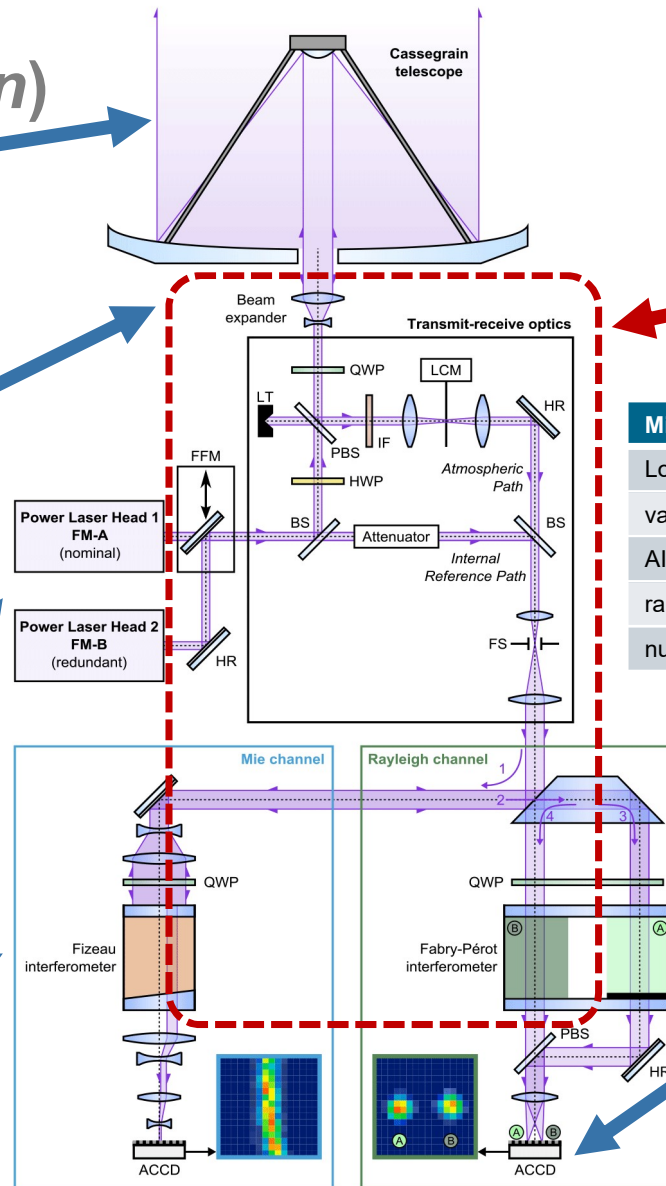
telescope aperture
receiver FOV
pointing errors
transmission transmitter chain
transmission receiver chain
pointing errors
nominal off-nadir angle

Laser

wavelength	pulse energy
linewidth	integration times

Fizeau spectrometer

peak transmission	fringe tilt
FWHM / FSR	geometrical factor



Only single transmission values!
 → „weakest“ part of the simulator
 → no simulation of angular dependent illumination

Miscellaneous

Look-Up-Table (= onboard DEM)	constants
various noise sources	master clock rate
AISP defaults parameters	delay times,
range-dependent bias	frequency arrays
number of pulses & measurement	sub-resolution

Fabry-Pérot spectrometer

peak transmission (direct reflected)
spectral spacing
FWHM / FSR
gaussian defects

ACCD

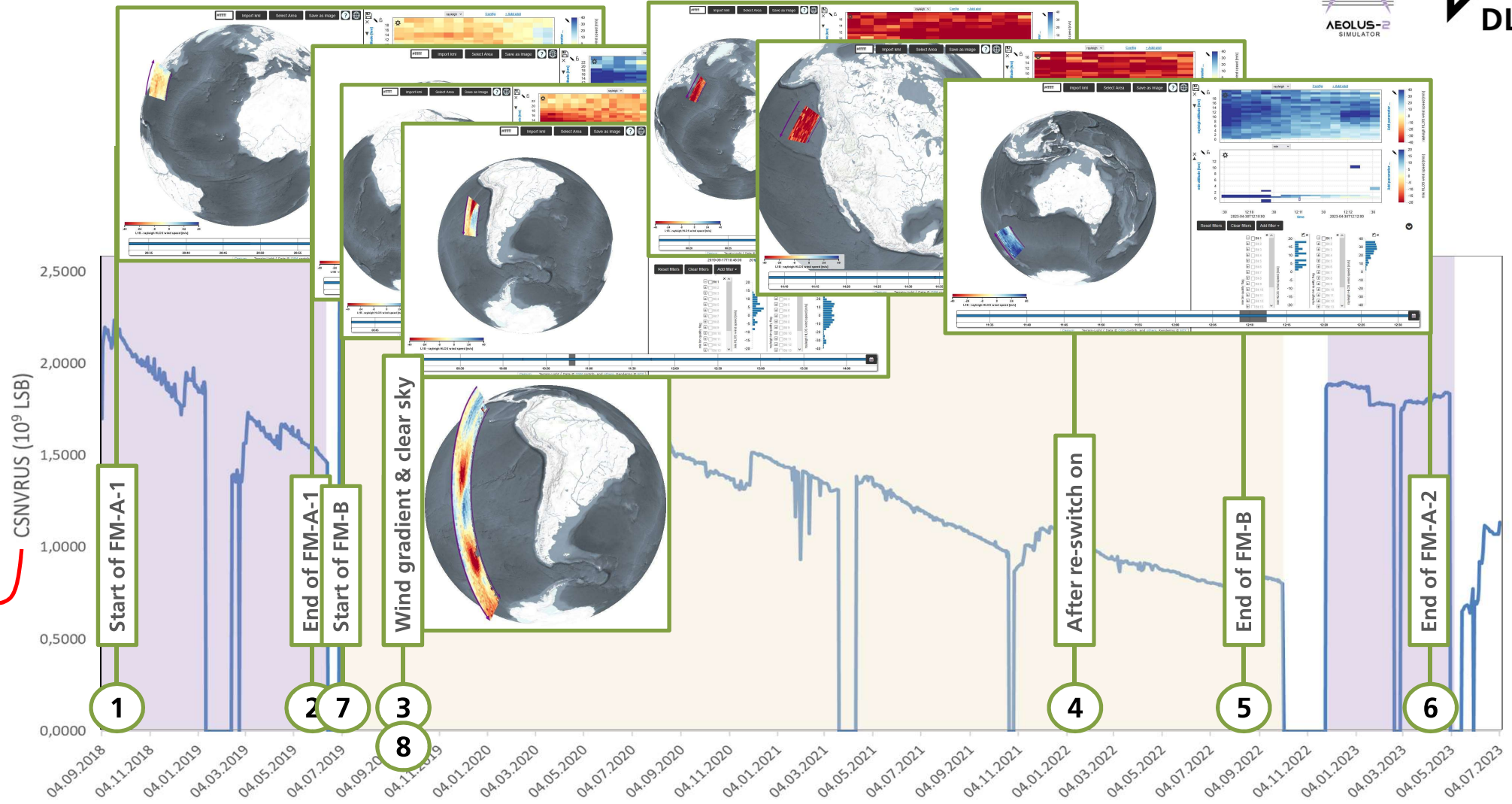
quantum efficiency
pixel characterisation (size, noise, ...)
tripod transmission
radiometric gain (LSB/e ⁻)
time in memory zone

Lux et al. (2021) AMT

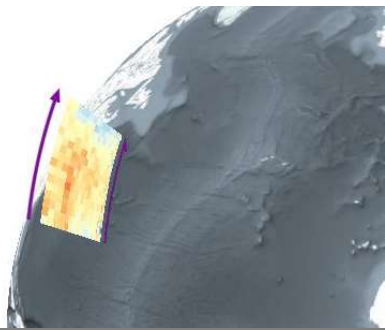
Scenarios for comparison of Aeolus vs. Simulation



Clear Sky Normalized Valid Rayleigh Useful Signal



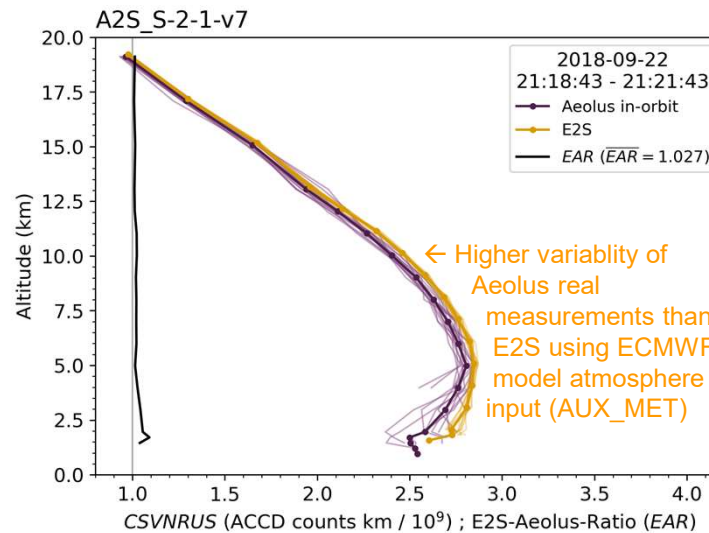
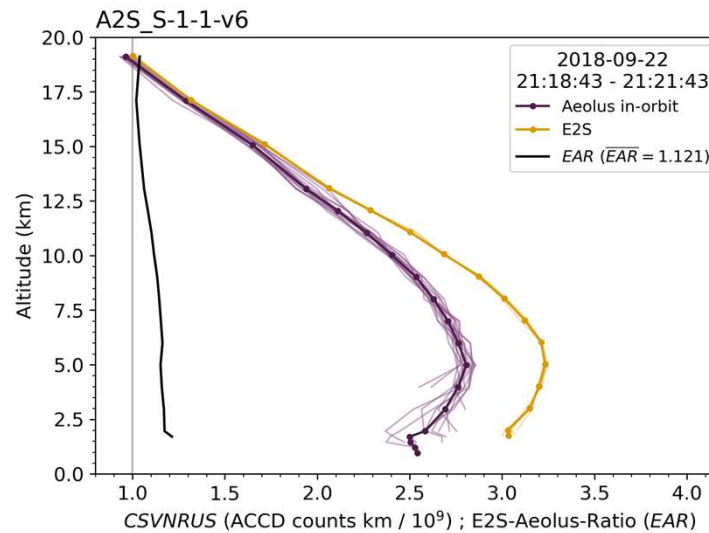
Comparison of Rayleigh useful signal



Scenario

1

	FM-A-1
Date	22.09.2018
NRT Baseline (L1B Version)	1B14 (07.12)
Pulse energy / mJ	62
Pulses / Measurements	19 / 30



• U.S. Standard Atmosphere

- constant wind profile
- median aerosol profile*
- adapted dynamic range, noise, ...



- + atmosphere from ECMWF: wind, temperature and pressure → and derived molecular backscatter and extinction

- + transmission values



• Simulator-to-Aeolus-ratio ≈ 1

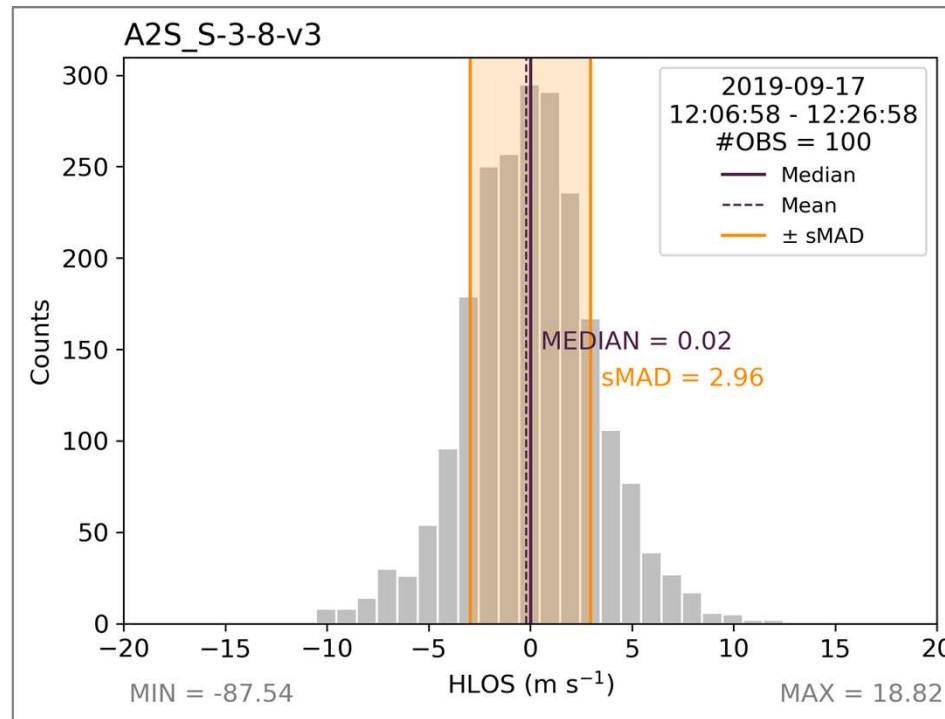
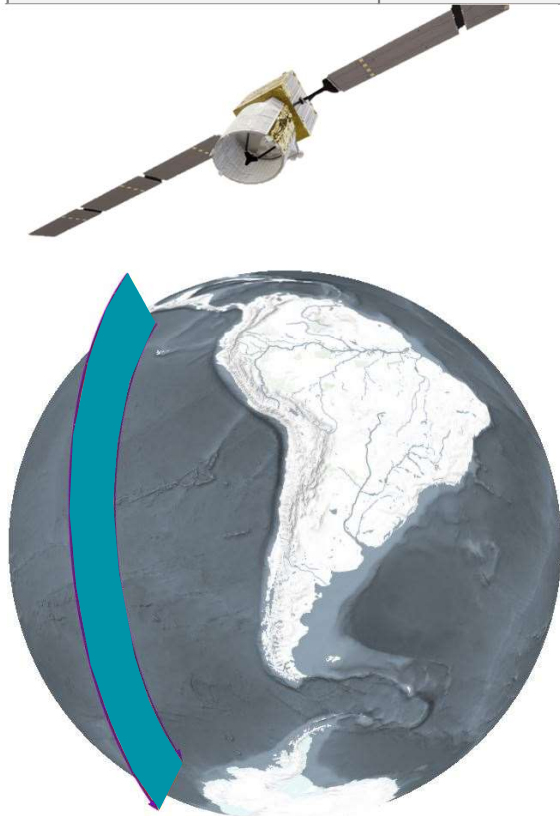
- Only very weak altitude dependence

* ESA, Reference Model of the Atmosphere (1999)

Comparison of Rayleigh wind random error



8	Laser	FM-B
	Date	17.09.2019
NRT Baseline (L1B Version)		1B12 (07.10)
Pulse energy / mJ		61.5
Pulses / Measurements		19 / 30



• U.S. Standard Atmosphere

- constant wind profile (0 m/s)
- median aerosol profile*
- adapted dynamic range, noise, transmissions, ...



- **Bias:** almost **0 m/s** reached
(mainly by adapting the calibration and noise parameters)

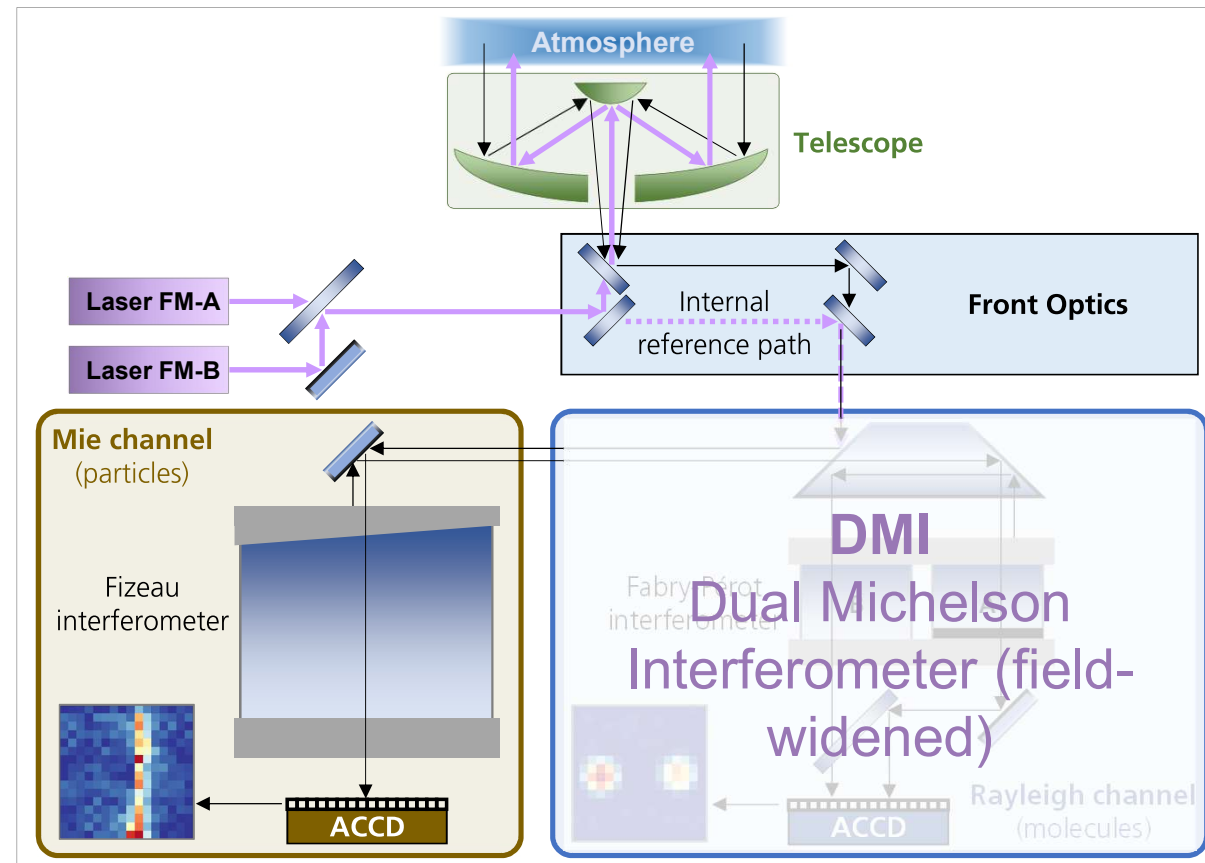
- **Random error (HLOS):**
 - **E2S:** (vs. median from 0-profile)
sMAD ≈ **3.0 m/s**
 - **Aeolus:**
sMAD ≈ 3.4 m/s
(from ECMWF O-B, corrected for B-error)

* ESA, Reference Model of the Atmosphere (1999)

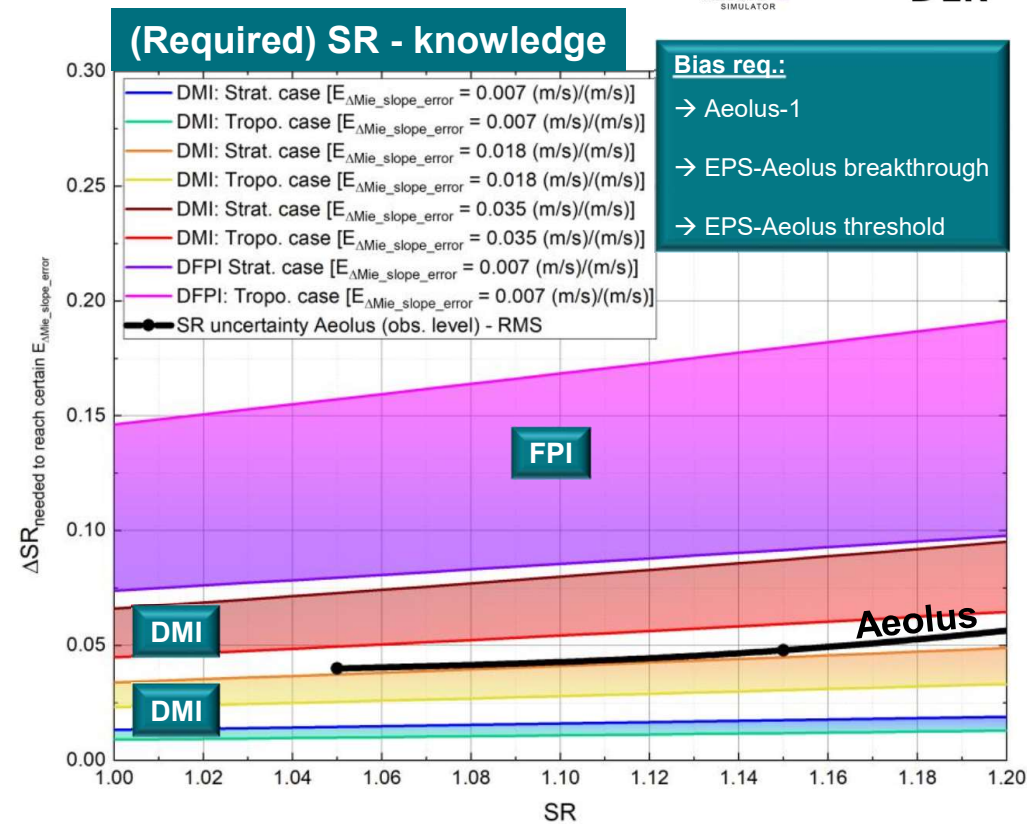
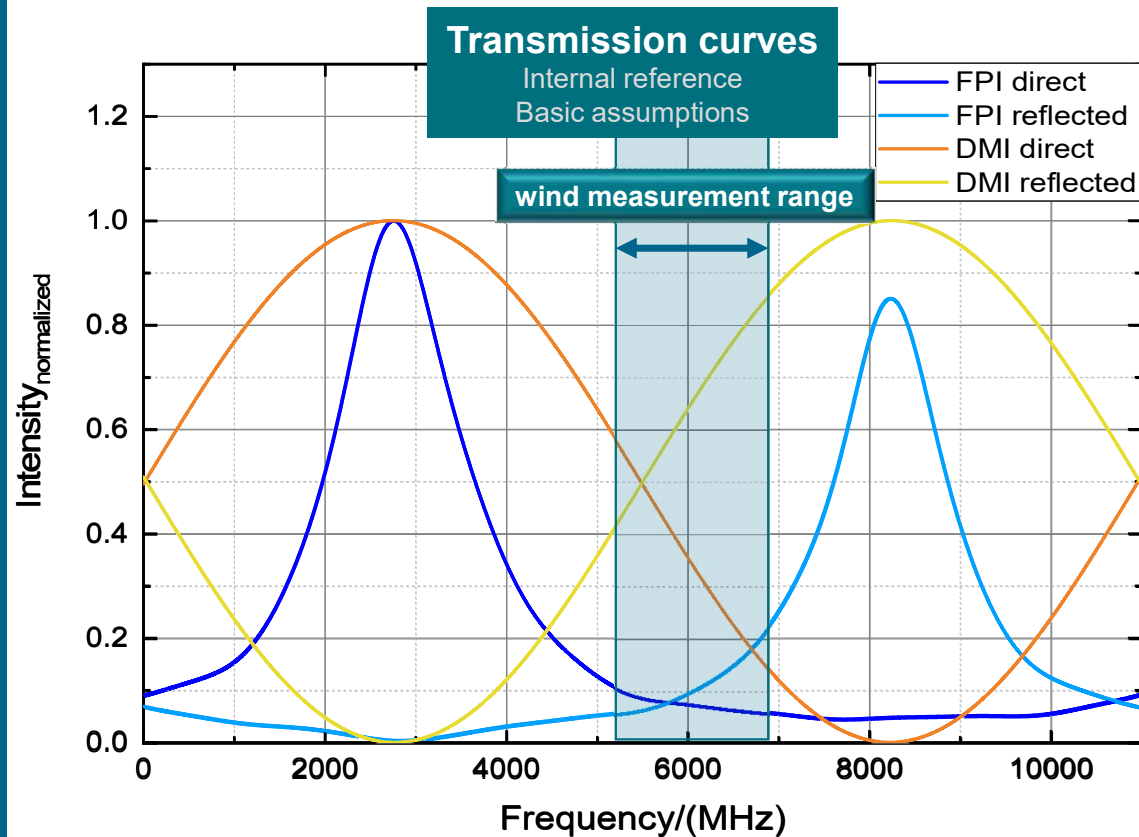
Dual Michelson Interferometer (DMI) & Mie cross-talk



- **Mie contamination on Rayleigh** signal is a significant error contributor for $1.3 < SR < 2.0$ (depending on atmospheric T)
- → In regions where Mie SNR is too low to derive accurate Mie winds (**intermediate SR regime**) a “correction” of Rayleigh winds is needed (critical for DMI)
- → Either **correction** of contamination or **flagging** of observations
- Both approaches need determination of **scattering ratio** from EPS-Aeolus data (low accuracy from NWP)
- **EPS-Aeolus** will provide **backscatter information** → Influence of Mie contamination can be corrected.
- But: **How accurate** does the backscatter information need to be to meet the EPS-Aeolus wind **error requirements**?
- → **Tasks:**
 1. Model the DMI (transmission & responses)
 2. Determine the wind bias and assess a potential error correction scheme
 3. Compare to Aeolus DFPI



DMI vs. Fabry-Pérot-Interferometer (FPI)



- Mie contamination plays only a minor role for the FPI
- With Aeolus-like SR-knowledge the Mie sensitivity error would be $\approx 0.018 \text{ m/s} / (\text{m/s}) \rightarrow 1.8 \text{ m/s}$ on 100 m/s range
- The error wind speed dependent

Summary



- DLR has \approx **20 years of experience** with airborne wind lidars
- **Good correlation** achieved for Rayleigh clear air **signal** profiles simulated by the E2S and measured by Aeolus
- **Reasonable** results achieved for L1B Rayleigh wind **random errors** (3 m/s vs. 3.4 m/s) for a first case
- **Next step**: simulate **EPS-Aeolus** performance with E2S by updating the respective (and known) parameters and tuning to required wind random error specification and comparing performance of Aeolus to (expected) EPS-Aeolus
- **Future E2S improvements**: RSP spots, Fizeau illumination, Rayleigh solar background simulation, end-to end verification, ...
- A representative **DMI simulator** (as one option planned for EPS-Aeolus) was developed in the current study.
- The DMI **performance** was investigated and compared to the Aeolus FPI performance
- **Final step**: Validate real EPS-Aeolus measurements by E2S simulations and A2D(2G) measurements

The funding for this study was provided by  **EUMETSAT**

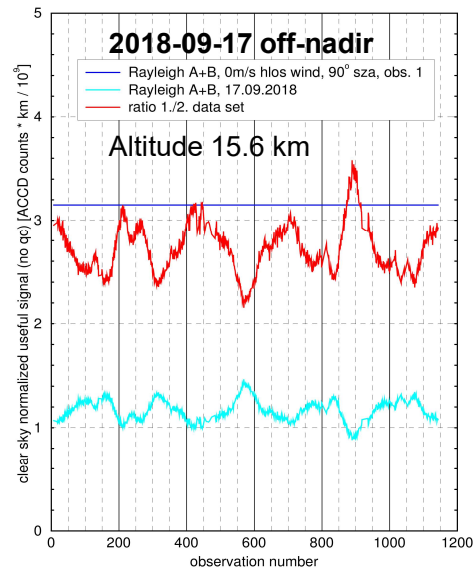
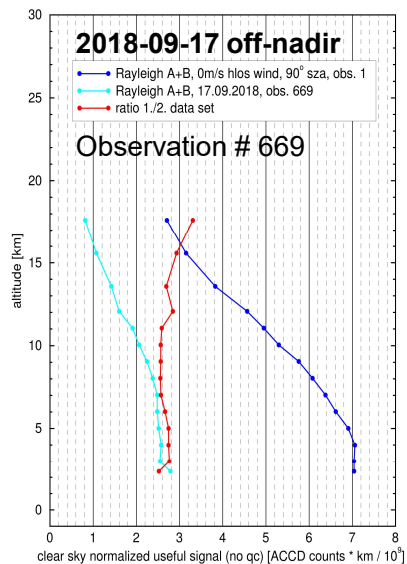
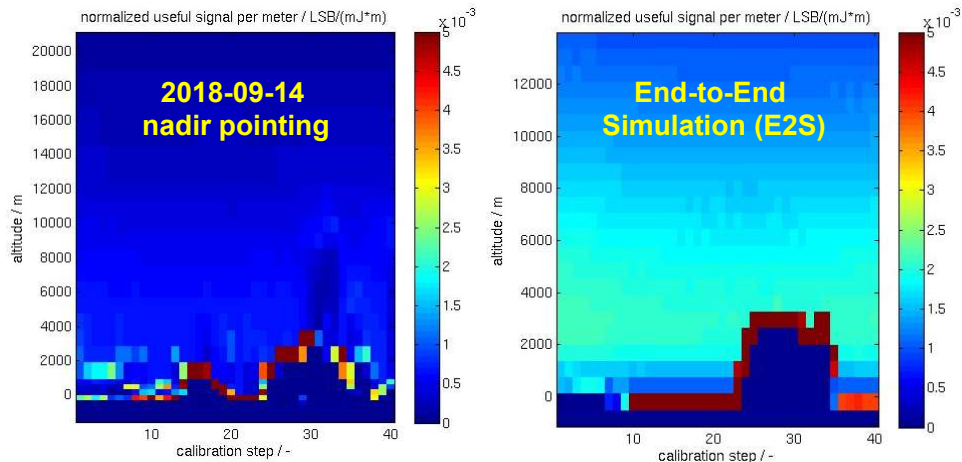


Thank
you for
your
attention!



BACK UP SLIDES

Radiometric performance assessment – Aeolus vs. E2S



- Use of **molecular Rayleigh backscattering** above clouds for instrument radiometric performance verification
 - only depending on atmospheric density (temperature)
 - low uncertainty
- Different approaches** (nadir / off-nadir viewing) and different tools at different teams show a factor of about 2.5 to 3 **lower Rayleigh signal** levels compared to pre-launch expectations (derived from End-to-End Simulations using default settings)
- Signal levels for high-albedo (ice) **ground returns** are even lower by factor 2.5 to 5.0 (Mie/Rayleigh)
- A factor 2.5 – 3.0 lower atmospheric signal signal increases **wind random error** by a factor of 1.6 – 1.7
- Signal loss potentially caused by a combination of beam clipping, characteristics of the telescope and the transmit-receive optics, and atmospheric turbulence.

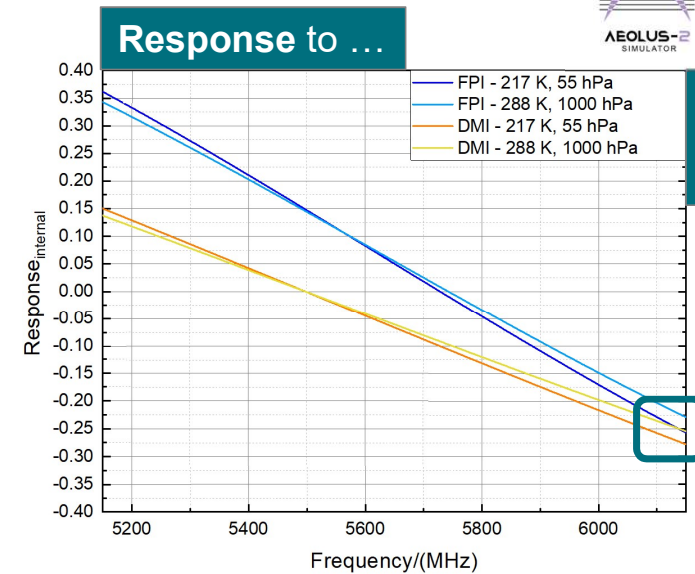
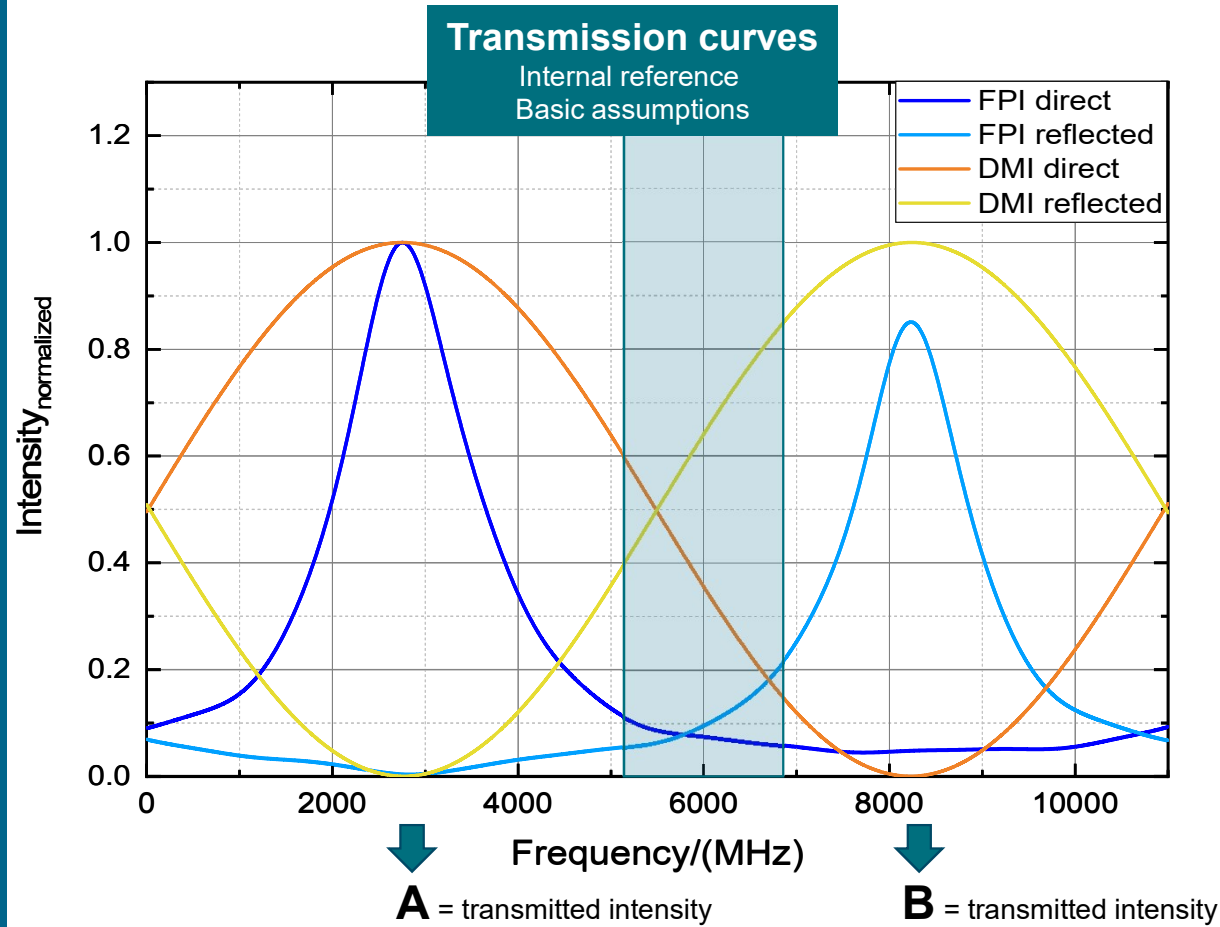
Limitations of the current version of the Aeolus End-to-End Simulator (E2S)



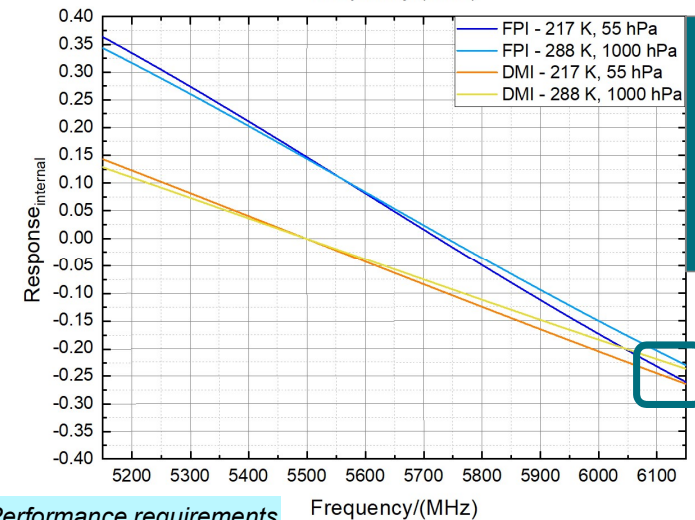
- As the **E2S has not been updated** since the launch of Aeolus → Feed the **knowledge gained from in-orbit operations** back into the simulator and produce realistic signals depicting real Aeolus observations
- Usage of **EO-CFI** for orbit and viewing geometry of EPS-Aeolus
- No **DEM Look Up Table** existing yet for EPS-Aeolus
- Not representing the full **complexity** of the Fizeau and Double Fabry-Perot interferometers
- **DMI** can only be integrated in the future
- No standard option to insert information about **depolarization** by particles
- Reliance on the ADAM **albedo** map and the inherent uncertainties.
- Inability to simulate certain types of **bias sources** (e.g. primary mirror temperature)

DMI vs. Fabry-Pérot-Interferometer (FPI)

$$R = \frac{A-B}{A+B}$$



broad-band
molecular
signal



broad-band
molecular
signal
+
Mie
contamination