STAND-ALONE COSMIC-RAY TOMOGRAPHY WITH SECONDARY PARTICLES

Muographers2023, Naples

Maximilian Pérez Prada Institute for the Protection of Maritime Infrastructures 21.06.2023





The growing success of cosmic-ray tomography applications is based on two

main measurement concepts: muon scattering and muon absorption



The growing success of cosmic-ray tomography applications is based on two main measurement concepts: muon scattering and muon absorption

- The scattering angle correlates with the density of the examined volume
- The absorption rate correlates with the density of the examined volume
- → Both methods allow a detailed
 3D volume reconstuction





A missing piece: secondary particles



The interaction of muons leave traces: secondary particles





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Image sources: Research Gate, Laboratory for Nuclear Technologies, ELENA

A missing piece: secondary particles

The interaction of muons and other cosmic-ray shower particles leave traces: secondary particles





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Image sources: Research Gate, Laboratory for Nuclear Technologies, ELENA

Let's put it to a test



Previous publication: MDPI Instruments, DOI 10.3390/instruments6040066

Scenario:

- 4 detector layers with perfect acceptance & efficiency: above, below, left & right of container
- Container: steel box (4 mm thick walls) with 20 ft ISO dimensions



Reconstruction procedure



Step 1: Selection and combination of different secondary particles and detection layers

 \rightarrow Possible discrimination between materials due to different secondary particle kinematics

	Photons	Neutrons	Electrons
Upper detector	<i>M</i> 1	M2	_
Sidewise detectors	M3	M4	_
Lower detector—production	M5	M6	_
Lower detector—absorption	M7	M8	M9

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Step 2: 3D voxel map creation

- \rightarrow Ray-tracing: trace back the secondary particles from their point of detection through the whole voxelized volume
- \rightarrow The more crossings (score) a voxel has, the higher its

density or atomic number

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Example 3D maps



One lead and one water block – empty container subtracted

Lead optimized combination



Water optimized combination



More Details:

Pérez Prada, M.; Barnes, S.; Stephan, M. Analysis of Secondary Particles as a Complement to Muon Scattering Measurements. *Instruments* 2022, *6*, 66. https://doi.org/10.3390/instruments6040066

Getting more realistic



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Getting more realistic



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- More realistic scenario includes:
 - Possible detector setup
 - \rightarrow Inclusion of realistic material thickness
 - to simulate impact on secondary particles



Getting more realistic



- Previous work was based on the assumption of perfect detection resolution and efficiency
- More realistic scenario includes:
 - Possible detector setup
 - \rightarrow Inclusion of realistic material thickness to simulate impact on secondary particles
 - Implementation of spatial resolution and detection efficiency
 - \rightarrow Only offline, no readout from the scintillator material is simulated \rightarrow No particle ID simulated, but using
 - GEN-truth information instead



Secondary particle detector setup





- 3 layers of plastic scintillator
- Plastic scintillator layer thickness: 50 mm
- Spacing between layers: 10 cm, 20 cm
- Detection efficiency per layer: 80%, 60%, 40%

Secondary particle detector setup



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- Plastic scintillator layer thickness: 50 mm
- Spacing between layers: 10 cm, 20 cm
- Detection efficiency per layer: 80%, 60%, 40%



 Distance between wavelength shifting (WLS) fibers: 10 mm, 20 mm, 30 mm

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Test setup & scenarios

- Cosmic-ray shower generation:
 - CRY library with all particles enabled and max. shower size of 30
 - 100M cosmic-ray shower from 10 m × 10 m
 plane (equivalent to ~30 min of scan time)



Source: CMS

A graphic representation of cosmic rays producing showers of particles



Test setup & scenarios

- Cosmic-ray shower generation:
 - CRY library with all particles enabled and max. shower size of 30
 - 100M cosmic-ray shower from 10 m × 10 m plane (equivalent to \sim 30 min of scan time)
- Test object:
 - Cube with volume of 1 m³ in the center of a simplified container
 - Material: water or lead
- Reconstruction:
 - Retuned the selection and combination of particle and detection layer for new setup
 - Voxel size in 3D map: 1 dm³



A graphic representation of cosmic rays producing showers of particles



Performance metrics



- To ensure a consistent measurement of the performance metrics, clustering is performed using the reconstructed density map as input
- Clustering method: set voxel with highest score as seed, loop over surrounding voxel and add them to the cluster, if its density score is at least 80% of the average score of the cluster

Performance metrics

- To ensure a consistent measurement of the performance metrics, clustering is performed using the reconstructed density map as input
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Cluster metrics:

- Reconstructed object density score (average)
- Reconstructed object size: volume, side lengths
- Reconstructed object shape: Chamfer distance* between reconstructed and GEN-truth object



* Chamfer distance: sum of the squared distances between nearest neighbor correspondences of two point clouds (X, Y) centered at origin

1-layer vs. 3-layer detection: lead block

20000

19000

Density Score

16000

- 15000

-2

 $^{-1}$

0 + Imi

1

2

1-layer



- 100% eff., perfect res.
- 1-layer: •

6500

6250

6000

- 5750 -Density Score

5250

5000

4750

-2

-1

0 + lmi

1

2

3-layer, 10 cm

Particle is detected and perfectly reconstructed from 1st layer only

3-layer: •

> Particle is detected only if it hits all layers and reconstructed by fitting the layer hits

GEN-truth 3-layer, 10 cm spacing 3-layer, 20 cm spacing 1-layer 17829 5668 Score 5896 ---Size [m³] 1,210 0,942 1,388 1,332 dx [m] 1,1 1,1 1,1 1,1 dy [m] 1,0 1,0 1,0 1,0 dz [m] 1,1 1,2 1,9 1,8 Chamfer 4,8 12,1 11,2 _ _ _

2.0

1.5 1.0

0.5

-0.5

-1.0

-1.5

-2.0

-4 -3

-2

-1

0 Y [m]

1

2

3

4

Z [m] 0.0

- Setup has an acceptance of ~33%
- Significant size increase in z-direction
- Slight increase in Chamfer distance

20

2.0

1.5

1.0

0.0 -0.5

-1.0

-1.5

-2.0

-4

-3 -2

-1

0 Y [m]

1

2

3

4

Z [m] 0.5

1-layer vs. 3-layer detection: water block

5600

5400

5200

0005 - 00

4600

4400

4200

-2

-1

+ Imi

0

1

2

1-layer



- 100% eff., perfect res.
- 1-layer:

- 2900

- 2800

2700

- 2600 ឫ

Density

2400

- 2300

-2

 $^{-1}$

+ Imi

0

2

3-layer, 20 cm

Particle is detected and perfectly reconstructed from 1st layer only

3-layer:

Particle is detected only if it hits all layers and reconstructed by fitting the layer hits

GEN-truth 3-layer, 10 cm spacing 3-layer, 20 cm spacing 1-layer 5182 2762 2755 Score ---Size [m³] 1,210 1,050 1,078 1,066 dx [m] 1,1 1,1 1,0 1,0 dy [m] 1,0 1,0 1,0 1,0 dz [m] 1,1 1,4 1,7 1,7 Chamfer 4,6 6,4 8,0 _ _ _

2.0

1.5

1.0

0.5

0.0

-0.5

-1.0

-1.5

-2.0

-4 -3

-2

-1

0 Y [m]

1

2

3

4

Z [m]

- Setup has an acceptance of ~53%
- Small size increase in zdirection
- Slight increase in Chamfer distance

2.0

1.5

1.0

0.5

0.0

-0.5

-1.0

-1.5

-2.0

-4

-3 -2

-1

0 Y [m]

1

2

3

4

Z [m]

Impact of detection efficiency: lead block

100% Eff.

6500

6250

6000

sity Score

- 5500 อี

5250

5000

4750

-2

 $^{-1}$

+ Imi

0

1

2



10 cm spacing, perfect resolution

20% Eff.

- 1400

- 1350

1300

- 1220 -Isity Score

- 1200 อี

- 1150

- 1100

1050

-2

 $^{-1}$

+ Imi

0

1

2

- For simplicity, all particles in all energy regimes have same efficiency
- Cor. Score corrects for the loss in effective scan time

	100% Eff.	80% Eff.	60% Eff.	40% Eff.	•
Cor. Score	5896	6020	6475	7367	
Size [m³]	1,388	1,371	1,117	0,627	
dx [m]	1,1	1,1	1,1	1,1	•
dy [m]	1,0	1,1	1,1	1,1	
dz [m]	1,9	2,0	1,7	1,6	•
Chamfer	12,1	14,1	5,3	7,6	

2.0

1.5

1.0

0.5

0.0

-0.5

-1.0

-1.5

-2.0

-4

-3 -2

-1

0 Y [m]

1

2

3

4

[m] Z

- Significant increase in density score with lower efficiency
- Size reduction in z-direction
- Slight decrease in Chamfer distance

2.0

1.5

1.0

0.5

0.0

-0.5

-1.0

-1.5

-2.0

-4

-3 -2

-1

0 Y [m]

1

2

3

4

Z [m]

Impact of detection efficiency: water block

100% Eff.

2.0

1.5

1.0

0.0

-0.5

-1.0

-1.5

-2.0

-4

-3 -2

-1

0 Y [m]

1

2

3

4

Z [m] 0.5

23

2900

2800

- 2700

- 2600 ឫ

- 2500 อี

- 2400

- 2300

2200

-2

-1

+ Imi

0

1

2

sitv



10 cm spacing, perfect resolution

50% Eff.

- 1550

- 1500

1450

- 1400 - 1400 - Density Score

1300

- 1250

-2

 $^{-1}$

+ Imi

0

2

- For simplicity, all particles in all energy regimes have same efficiency
- Cor. Score corrects for the loss in effective scan time

	100% Eff.	80% Eff.	60% Eff.	40% Eff.
Cor. Score	2762	2962	3255	3983
Size [m³]	1,078	1,132	1,079	0,551
dx [m]	1,0	1,1	1,1	1,1
dy [m]	1,0	1,1	1,0	1,1
dz [m]	1,7	1,8	1,8	1,5
Chamfer	6,4	10,9	7,1	7,1

2.0

1.5

1.0

0.5

-0.5

-1.0

-1.5

-2.0

-4

-3 -2

-1

0 Y [m]

1

2

3

4

Z [m] 0.0

- Significant increase in density score with ower efficiency
- Size reduction in z-direction

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Impact of detection resolution: lead block





- 10 cm spacing, 100% effciency
- Resolution due to scintillator thickness: 50 mm
- Resolution due to WLS fiber grid spacing: 10 mm, 20 mm, 30 mm
- For simplicity, all particles in all energy regimes have same resolution

	Perfect Res.	10:50 mm Res.	20:50 mm Res.	30:50 mm Res.	•
Score	5896	5726	5573	5266	
Size [m³]	1,388	1,357	1,352	1,419	
dx [m]	1,1	1,0	1,0	1,2	•
dy [m]	1,0	1,0	1,0	1,2	
dz [m]	1,9	2,0	2,0	1,9	
Chamfer	12,1	16,3	17,7	11,5	

- Small reduction of density score with higher resolution
- Small increase in size,
 mostly in x- and y direction with very high
 resolution

Impact of detection resolution: water block



• 10 cm spacing, 100% effciency

30:50 mm Res.

- 2600

- 2500

- 2400 - 2400 - 2400 Density Score

- 2200

- 2100

-2

 $^{-1}$

+ Imi

0

2

- Resolution due to scintillator thickness: 50 mm
- Resolution due to WLS fiber grid spacing: 10 mm, 20 mm, 30 mm
- For simplicity, all particles in all energy regimes have same resolution

- Small reduction of density score with higher resolution
- No significant statement about the correlation between size and resolution possible
- Slight increase in Chamfer distance



	Perfect Res.	10:50 mm Res.	20:50 mm Res.	30:50 mm Res.
Score	2762	2697	2640	2505
Size [m³]	1,078	1,053	0,988	1,050
dx [m]	1,0	1,1	0,9	1,1
dy [m]	1,0	1,0	1,0	1,0
dz [m]	1,7	1,7	2,0	1,8
Chamfer	6,4	5,6	9,5	10,7

 $^{-1}$

0 Y [m]

1

2

3

4

Impact of layer spacing



Calculate ratio 10 cm / 20 cm layer spacing scenario for density score and size

Water	Perfect Eff. & Res.	80% Eff.	60% Eff.	40% Eff.	10:50 mm Res.	20:50 mm Res.	30:50 mm Res.
Score	1,003	1,014	1,020	1,017	0,996	0,961	0,926
Size [m³]	1,011	1,154	1,255	1,611	1,021	1,035	1,087

Lead	Perfect Eff. & Res.	80% Eff.	60% Eff.	40% Eff.	10:50 mm Res.	20:50 mm Res.	30:50 mm Res.
Score	1,040	1,039	1,066	1,040	1,020	1,007	0,967
Size [m³]	1,042	1,011	0,882	1,161	0,996	0,988	1,065

- \rightarrow Only small difference between spacing scenarios for lead block, no difference for water block
- \rightarrow Only small or no significant difference for efficiency and resolution variations
- \rightarrow Allows for some variability in the optimal detector layout



Calculate ratio lead / water block scenario for density score and size

10 cm layer spacing	Perfect Eff. & Res.	80% Eff.	60% Eff.	40% Eff.	10:50 mm Res.	20:50 mm Res.	30:50 mm Res.
Score	2,135	2,032	1,989	1,849	2,123	2,111	2,102
Size [m³]	1,288	1,211	1,035	1,138	1,289	1,368	1,351

20 cm layer spacing	Perfect Eff. & Res.	80% Eff.	60% Eff.	40% Eff.	10:50 mm Res.	20:50 mm Res.	30:50 mm Res.
Score	2,057	1,983	1,904	1,809	2,072	2,016	2,013
Size [m³]	1,250	1,382	1,473	1,579	1,322	1,434	1,379

 \rightarrow Only small or no significant difference for efficiency and resolution variations

 \rightarrow Material discrimination is possible and stable for realistic efficiencies and resolutions

Final remarks

- The usage of secondary particles provides a promising and complimentary approach for cosmic-ray tomography
- The discussed detector conditions simulate realistic material budget, resolution and efficiency
- Even with a more realistic detection setup, our stand-alone method can succesfully reconstruct simple geometric objects located inside a container and discriminate their material



- Further work will try to optimize the setup, considering a wider range of detector materials
- Further planed improvements: machine learning based reconstruction, automized material parameter scan, ...