Knowledge for Tomorrow

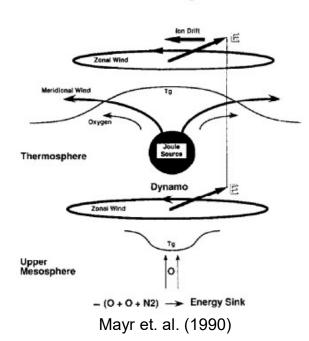
## Variations of total electron content with solar wind parameters at highlatitude ionosphere

P. Iochem<sup>1</sup>, C. Borries<sup>1</sup>, F. Davis<sup>1,2</sup>, S. Tasnim<sup>1</sup>, J. Vogt<sup>3</sup>

<sup>1</sup>German Aerospace Center (DLR), Institute for Solar-Terrestrial Physics, Neustrelitz, Germany <sup>2</sup>University of Applied Sciences Neubrandenburg, Neubrandenburg, Germany <sup>3</sup>Constructor University, Bremen, Germany

## Solar wind impact on high-latitude ionosphere

- Interaction between solar wind and magnetosphere drives the upper atmosphere.
- Changes in solar radiation, geomagnetic disturbances, and lower atmosphere forcing can all contribute to variability of maximum electron density of the F2 layer and the total electron content (TEC). (Cai et. al., 2021)
- Intense Joule heating in thermosphere results in upwelling of nitrogen-rich or oxygen-depleted air. Depleted O/N<sub>2</sub> air causing a decrease in electron density in polar regions due to a higher recombination process through ion-molecule exchange reaction with N<sub>2</sub> molecules (Ranjan et. al., 2023).
- There exist positive storm effects (electron density enhancements) and negative effects (electron density depletion) in the ionosphere (Borries et. al., 2015)



How does the thermosphereionosphere system respond to solar wind variation and how it changes with season and local time?



Auroral Region

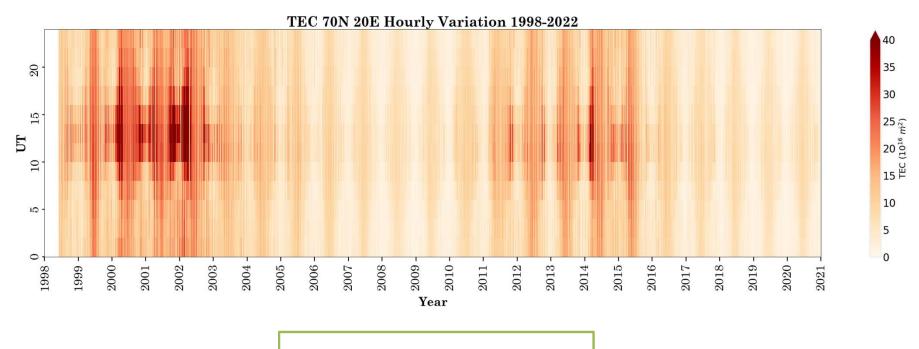
# **Objectives**

- Study the impact of solar wind on high-latitude ionospheric total electron content (TEC) on time scales of hours, seasons, years, and solar cycles.
- Reveal the differences in TEC and solar wind correlations.
- Study the response time of the ionosphere to solar wind variations.
- Explain the possible reasons and mechanisms behind the correlations.





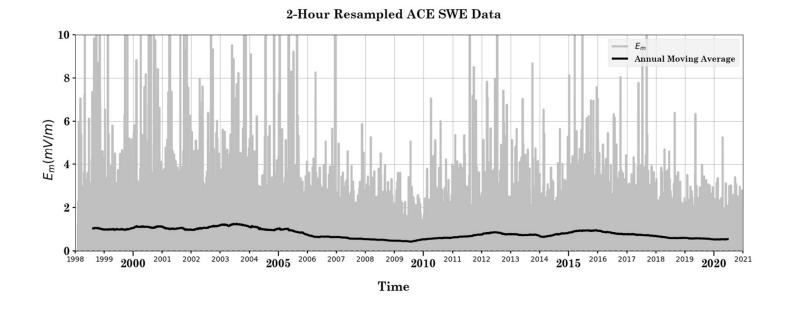
## Total Electron Content (TEC) and Daily Sunspot Number (SSN) Data

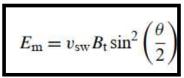


Data: 2 hours resolution IGS TEC Map Location: Near Tromsø/Norway, 70N 20E Time span: 23 years (1998-2021) Local time: UTC+1 (+2 daylight savings)



## Merging Electric Field as Solar Wind Coupling Function





Kan & Lee (1979)

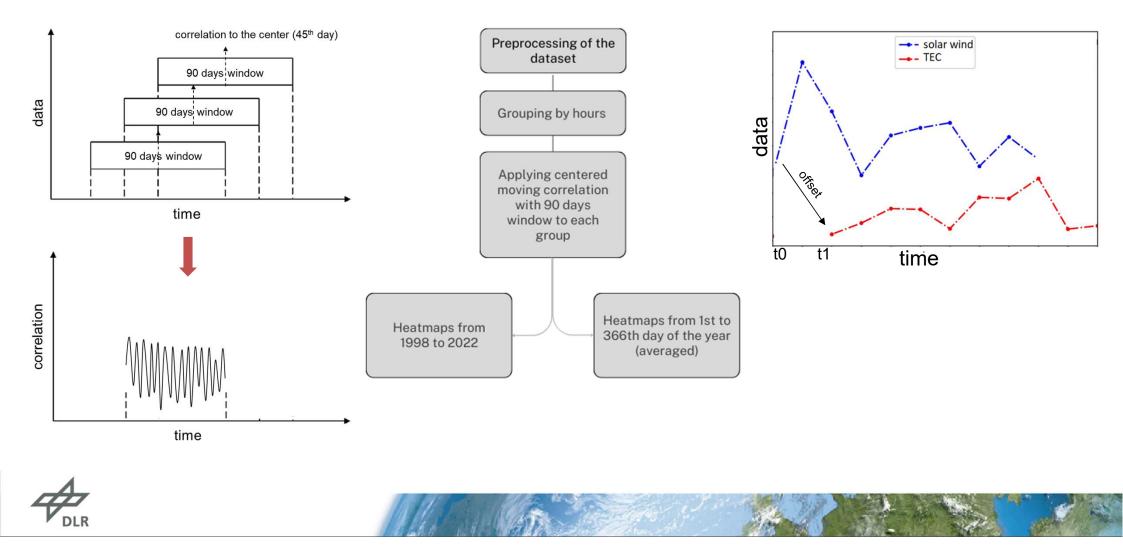
 $B_T = \sqrt{B_y^2 + B_z^2}$  Magnitude of IMF in yz plane  $V_{SW}$  = solar wind speed  $\Theta$  = Angle between z direction and projection

of IMF in yz-plane

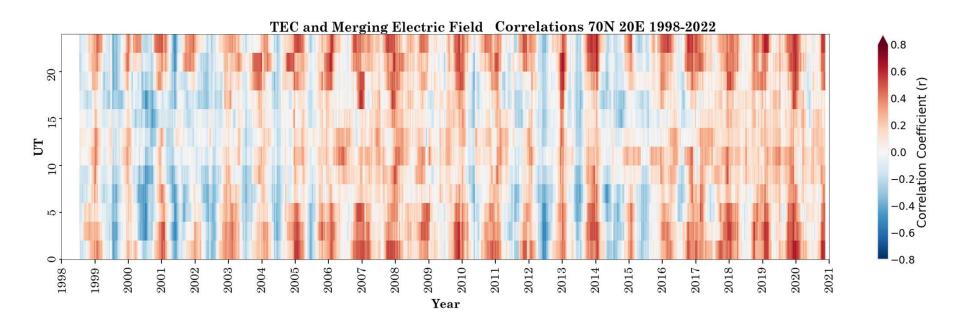
Data: 2 hours resampled ACE SWE data Time span: 23 years (1998-2021)



## Method: Centered Moving Correlation with 90 Days Window and Delay



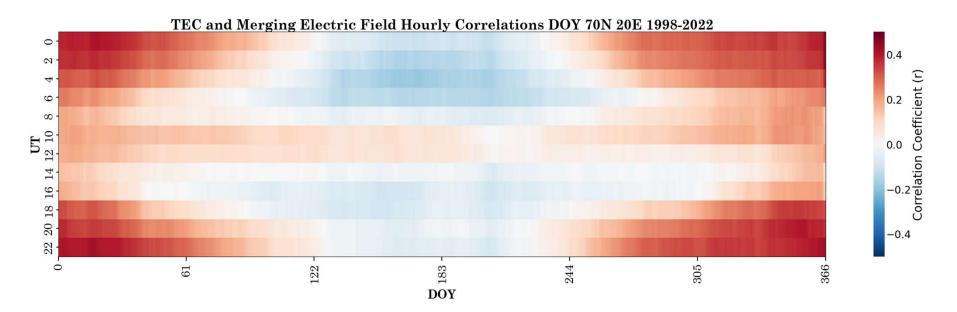
# Temporal Evolution of the TEC and *E<sub>m</sub>* Correlations



- Positive correlations during winter nighttime
- Negative correlations during summer daytime solar max



# Average Annual Variation of TEC and *E<sub>m</sub>* Correlations

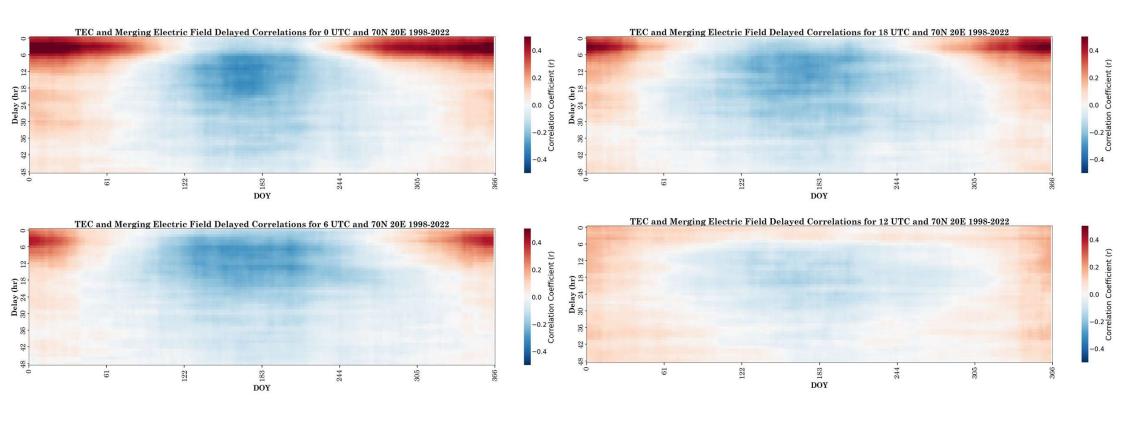


Positive-negative correlation effect

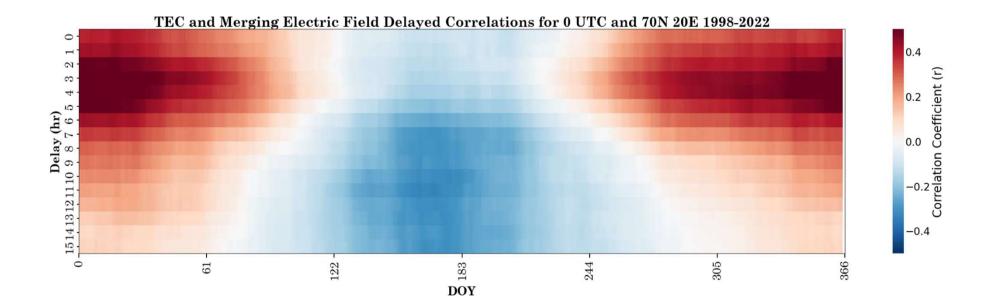
- Convection (winter-transport of plasma)
- Recombination (summer-reduction of plasma)



# Response Time of TEC to Solar Wind $(E_m)$









# Summary

- Seasonal and local time dependency of solar wind in high latitude ionosphere
- Merging electric field correlation analysis with Total Electron Content
  - 1. During winter (summer) it is strongly positive (mostly negative during daytime)
  - 2. Positive winter effect plasma transport (convection)
  - 3. Negative summer effect recombination (upwelling of molecules and more intense heating)
- Response of ionospheric TEC to solar wind:
  - 1. Shorter-positive winter effect
  - 2. Longer- negative summer effect





#### References

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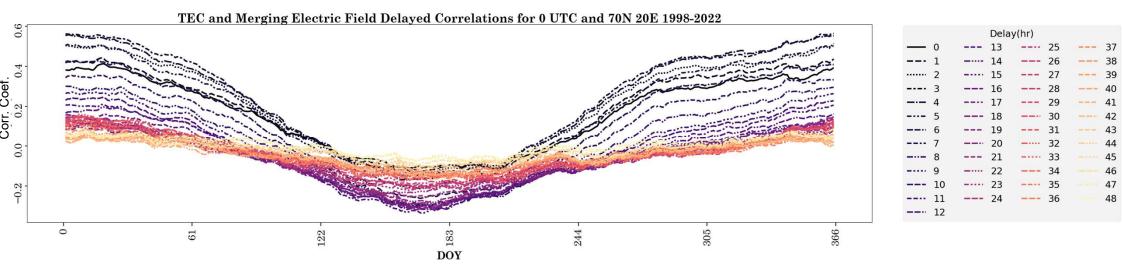
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- Richmond, A. D. (n.d.). Joule Heating in the Thermosphere Space Physics and Aeronomy. Space Physics and Aeronomy, Volume 4: Upper Atmosphere Dynamics and Energetics, 4. <u>https://cedarscience.org/sites/default/files/workshops/2022/Richmond2020preprint.pdf</u>



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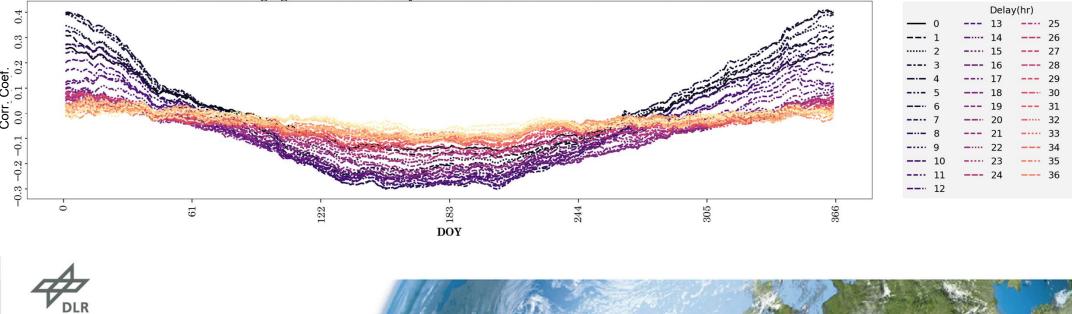
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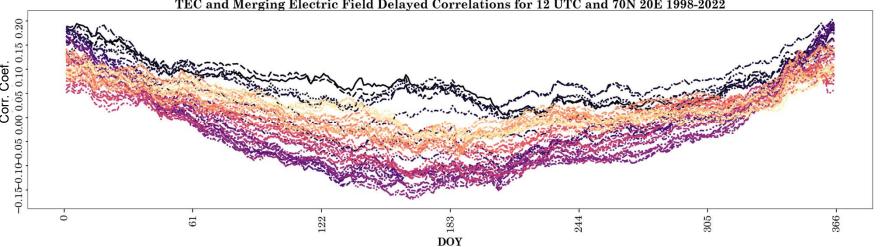
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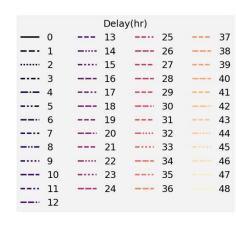
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TEC and Merging Electric Field Delayed Correlations for 6 UTC and 70N 20E 1998-2022







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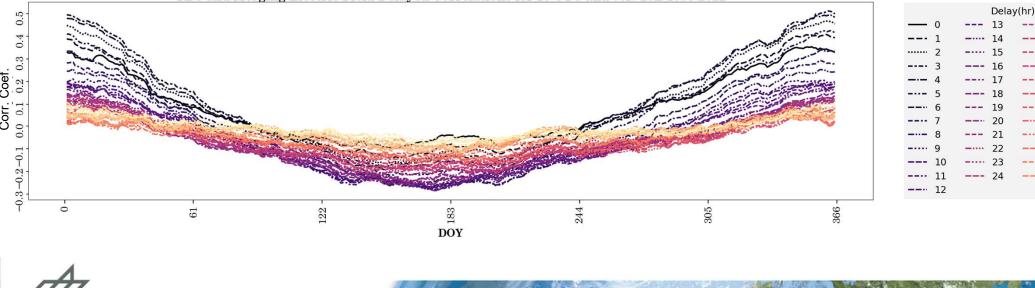
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TEC and Merging Electric Field Delayed Correlations for 18 UTC and 70N 20E 1998-2022

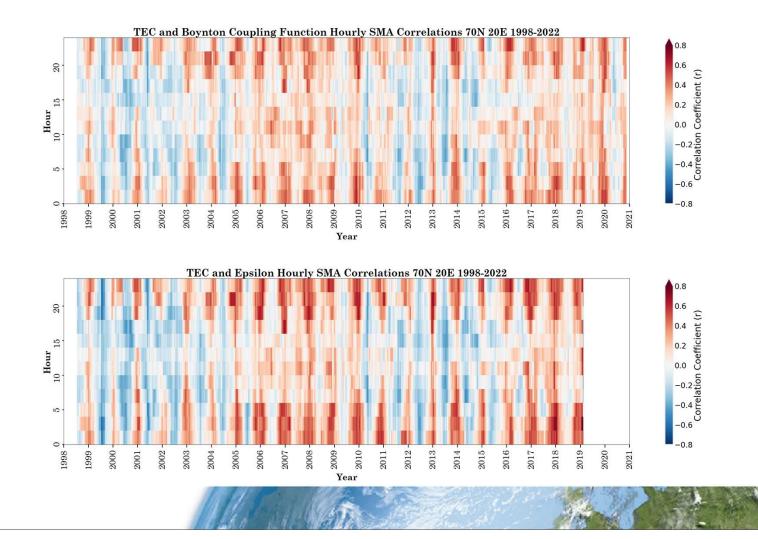


TEC and Merging Electric Field Delayed Correlations for 12 UTC and 70N 20E 1998-2022

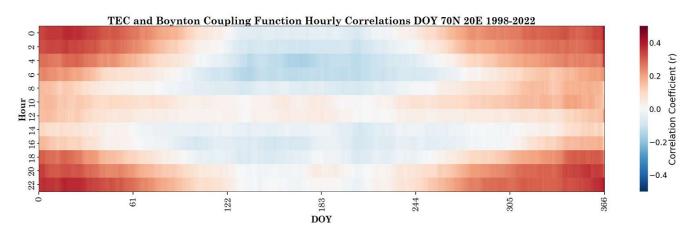
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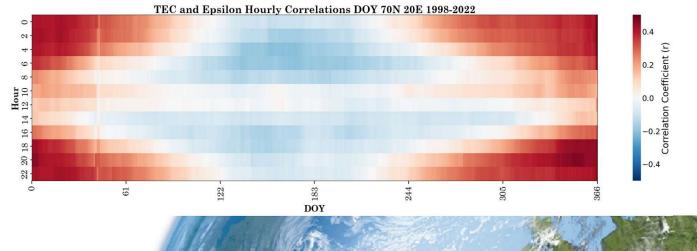
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#### **Temporal Evolution of the TEC and Other Coupling Parameters Correlations**

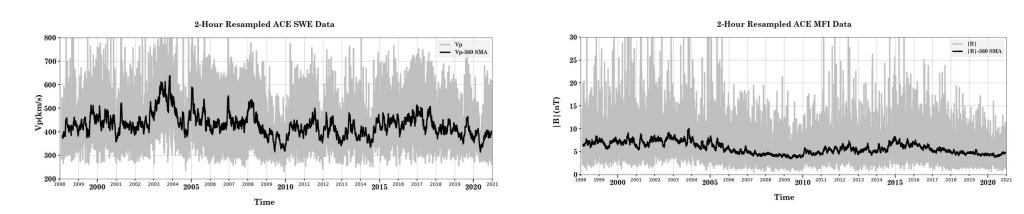


# Average Annual Variation of TEC and Other Coupling Parameters Correlations

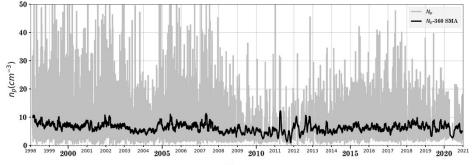




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2-Hour Resampled ACE SWE Data



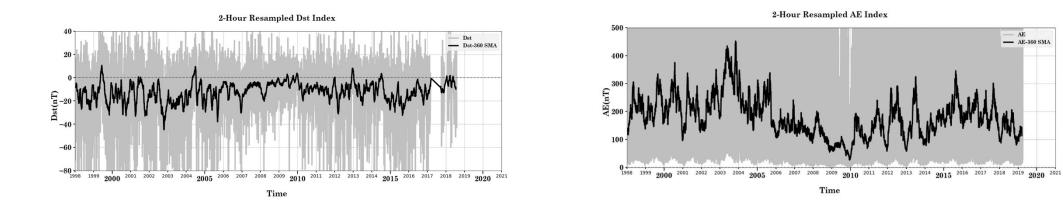




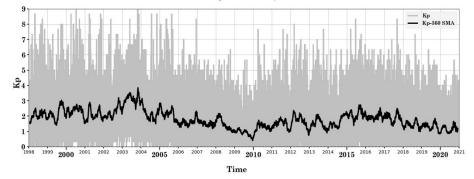


**Solar Wind Parameters** 

**Geomagnetic Activity Parameters** 



3-Hour Resampled Planetary-K Index







---- AE ----- AE-360 SMA