

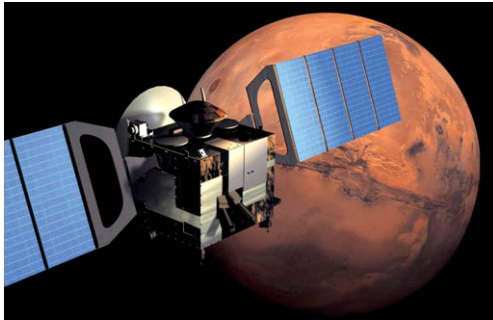
Patterns in textured dust storms in Mars North Pole

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Observations & Data from Mars Express



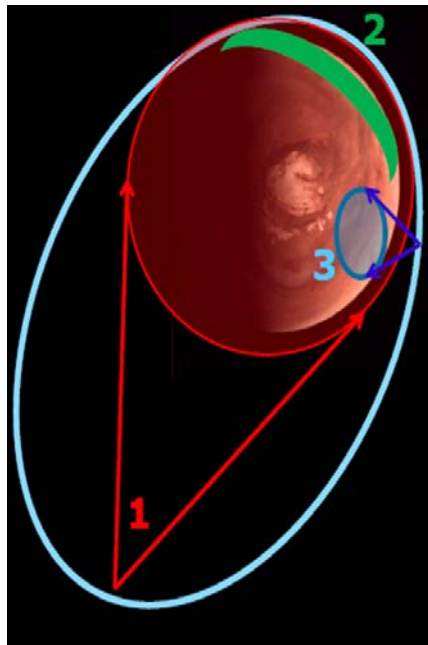
Mars Express (ESA)

Polar orbit:

Pericenter ~ 300 km

Apocenter ~ 10,000 km

Period ~ 7.5 hr



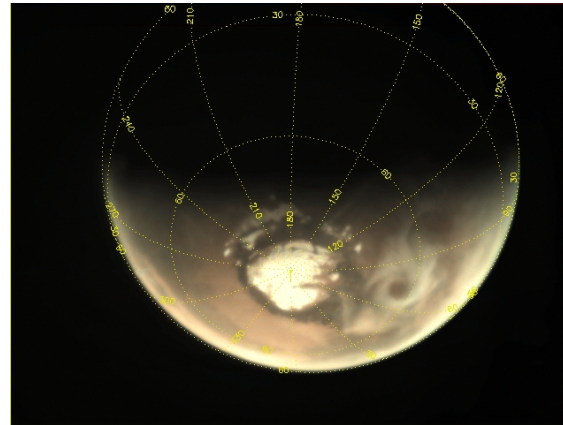
Visual Monitoring Camera VMC



FOV: $40^\circ \times 31^\circ$

Bayer RGB (COLOR)

Wavelength: 400-650 nm



See E. Ravanis et al. presentation

High Resolution Stereo Camera HRSC



2 stereo channels (S1, S2)

2 photometry channels (P1, P2)

1 nadir channel (ND)

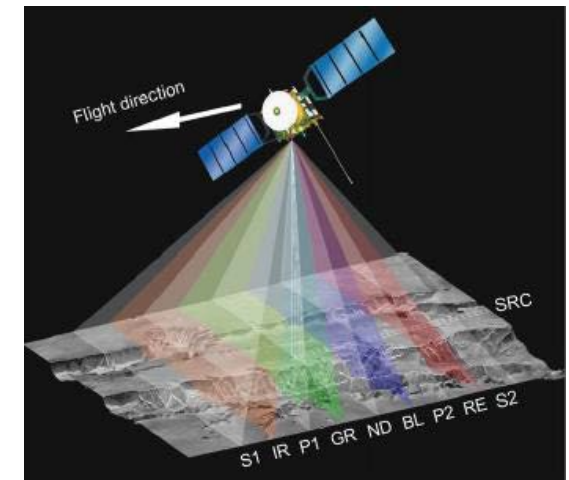
BL (blue channel 440 nm)

GR (green channel 530 nm)

RE (red channel 750 nm)

IR (near infrared channel 970 nm)

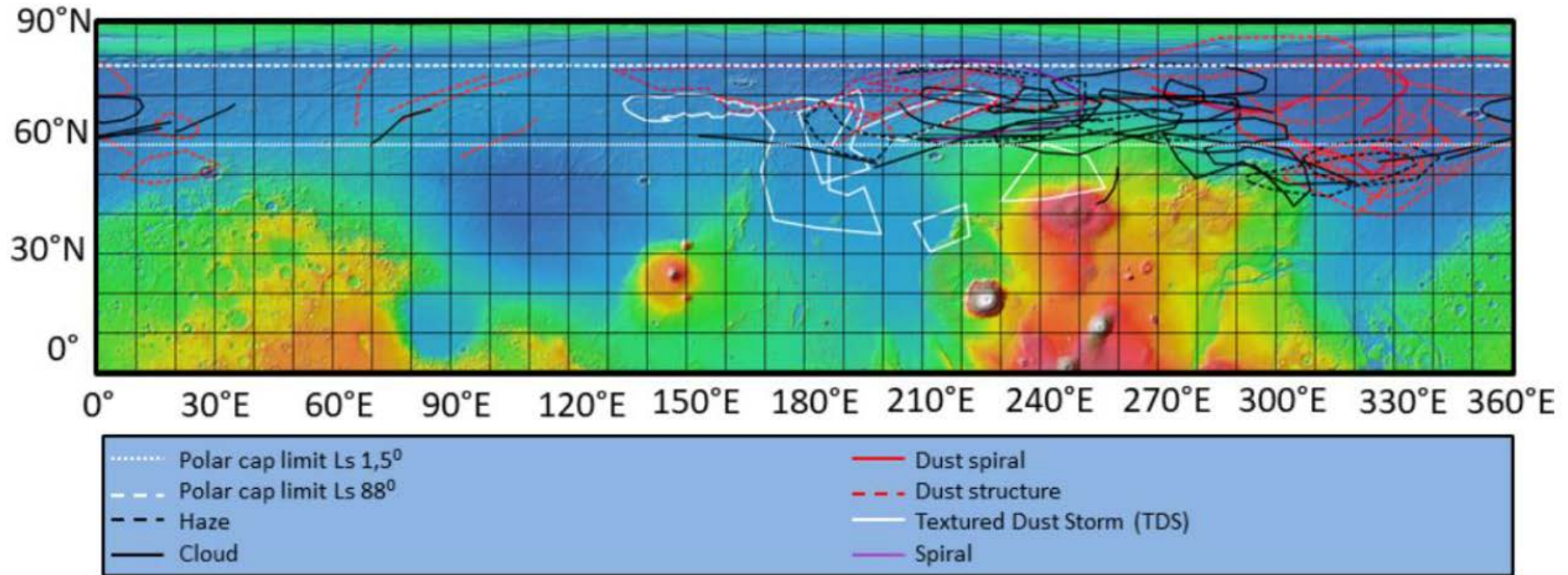
Maximum resolution = 10 m/pixel



Dynamical activity in Mars North Polar region in spring time

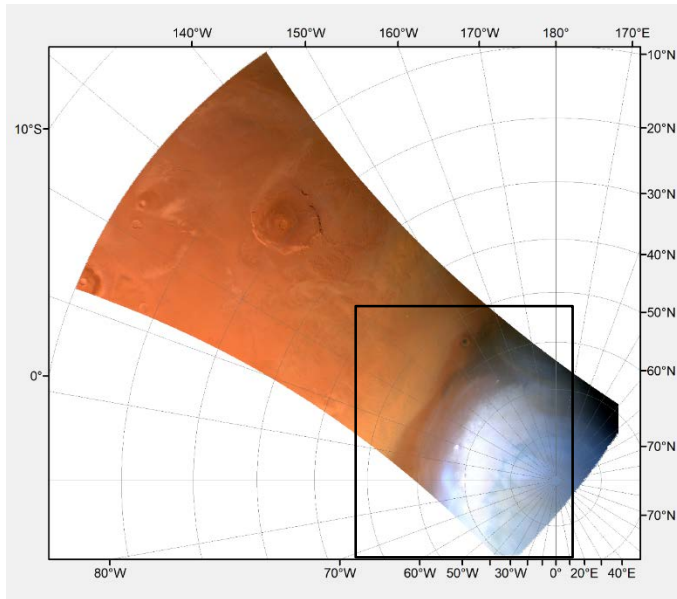
Period this study: 3 March – 17 July 2019
Ls = 350° (MY 34) - 54° (MY 35)
Aphelion is at Ls = 71°
Latitudes: 45°N - 90°N

Features: clouds and Hazes and a variety of local dust storms* (shapeless, arcs, fronts, spirals)

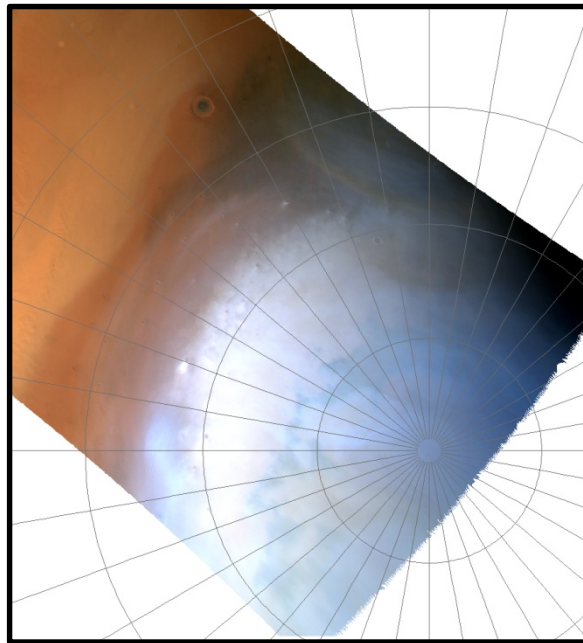
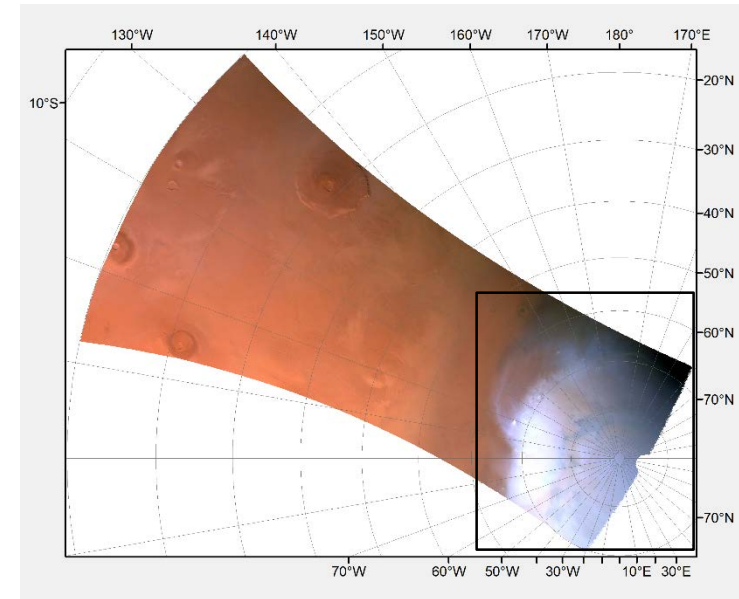


* Types of features in: Wang & Fisher (2009)

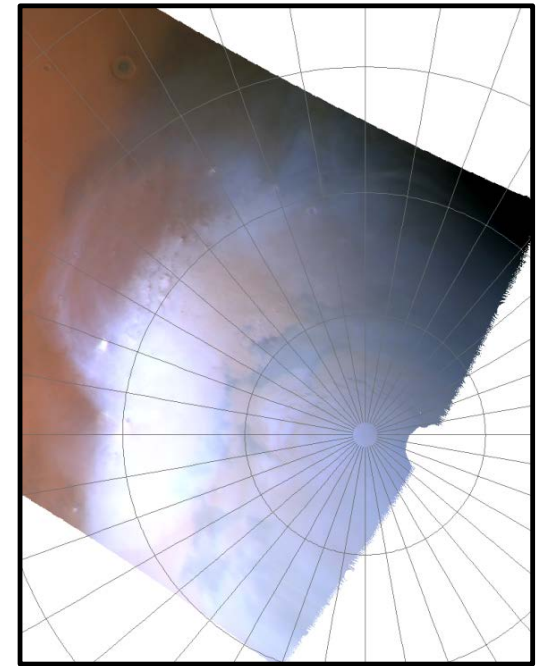
Shapeless dust areas over North Pole (HRSC)



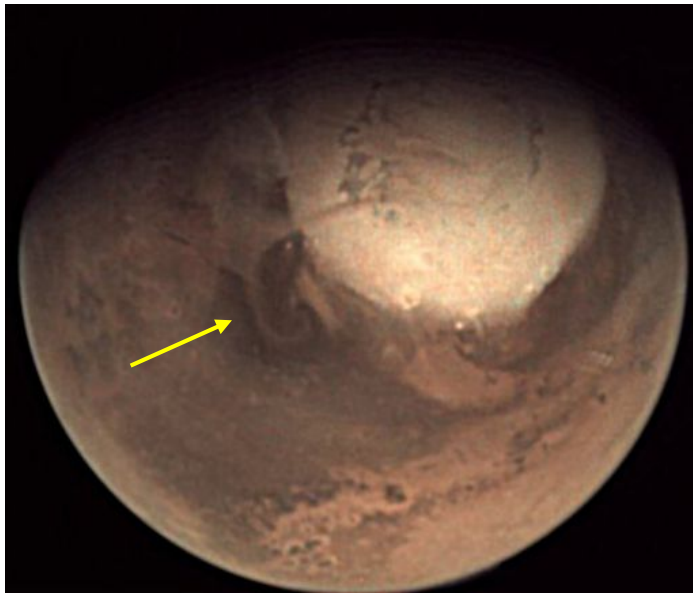
HRSC
6 June
Ls=35.4°



Dust patches
No texture
Dust total area:
 $7.92 \times 10^5 \text{ km}^2$

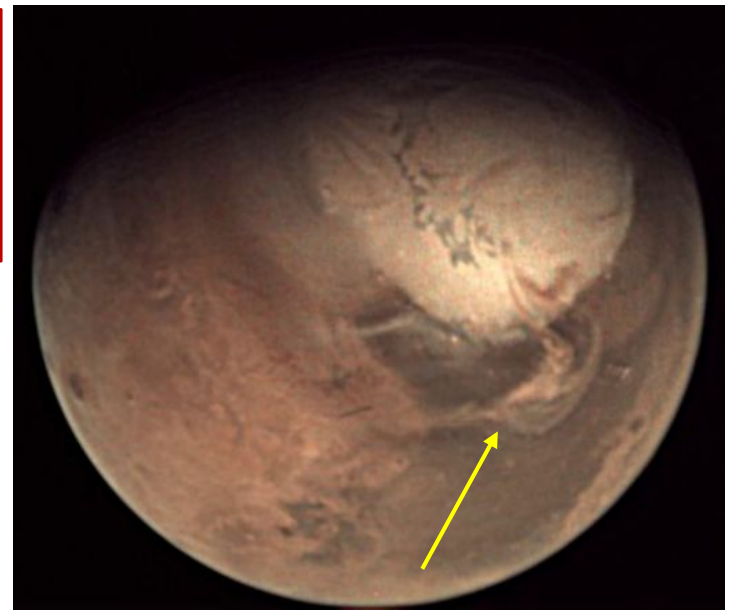


Long Filaments and Spirals (VMC)

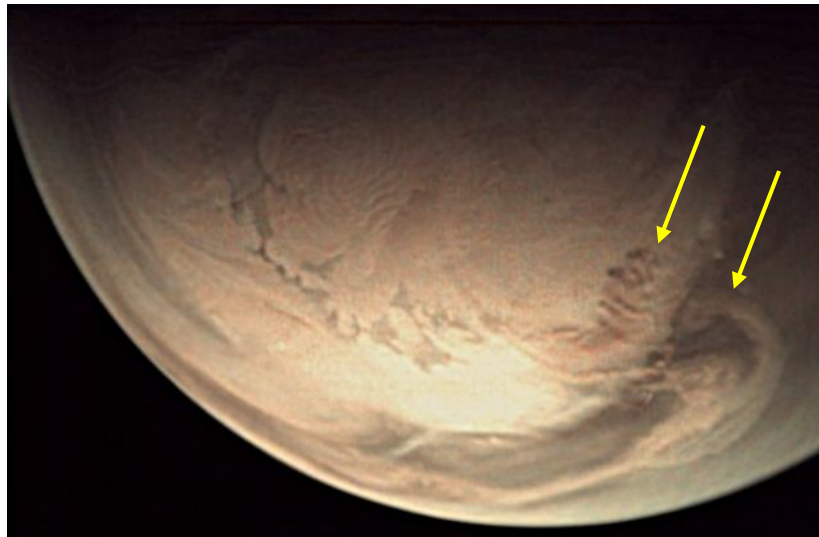


25 May

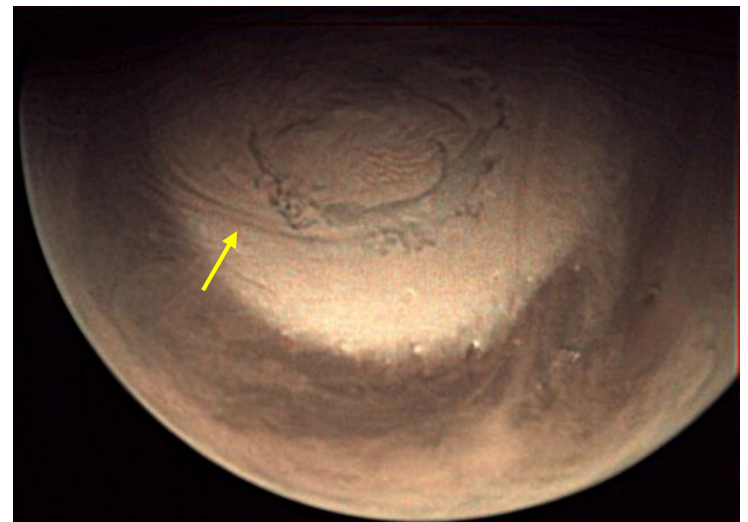
VMC
Ls=30-31°



27 May

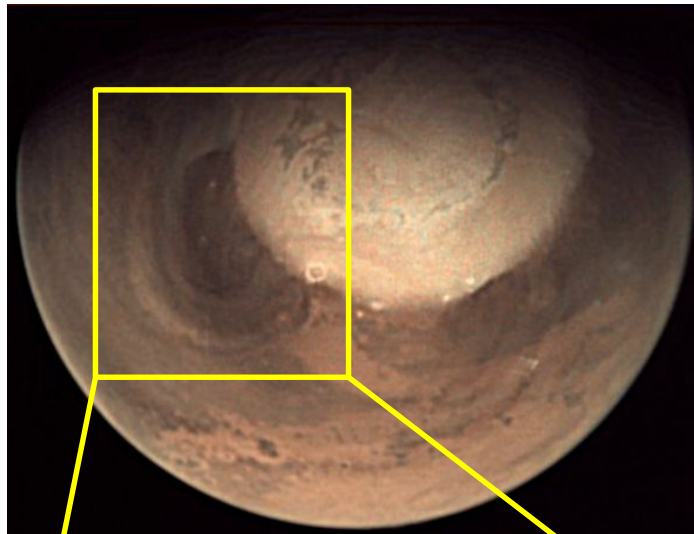


27 May

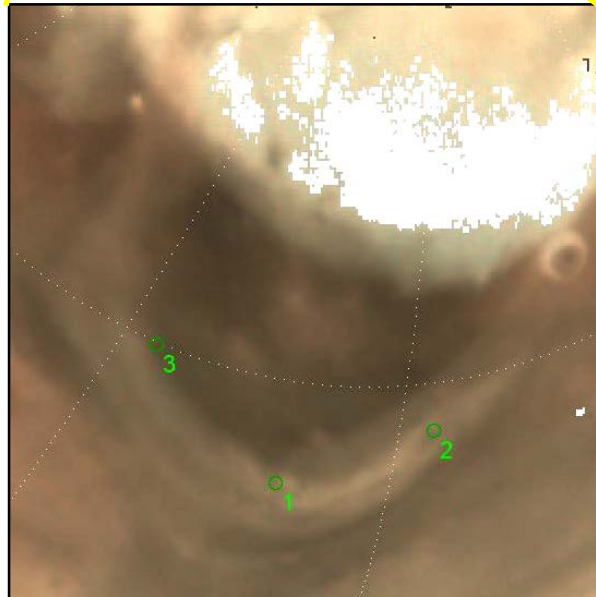


30 May

Arc shaped and “Flushing” dust storms (VMC)



3 June



Ls = 45° - 50°

Acidalia Planitia
(300°E - 330°E)

Arcs:

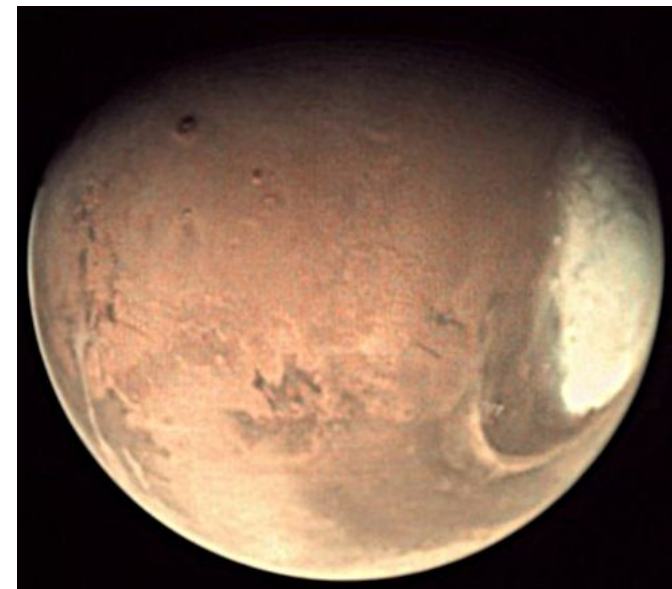
Length ~ 2350 km

Width ~ 250 km

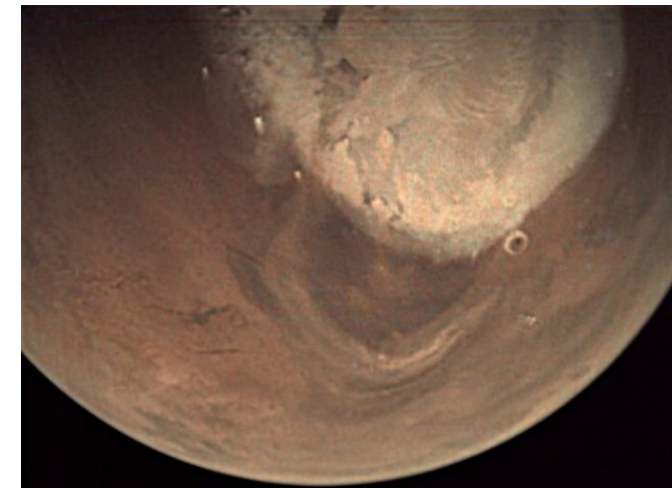
Expanding velocity:

Points (1-2-3)

$V = 29 \text{ ms}^{-1}$



5 July

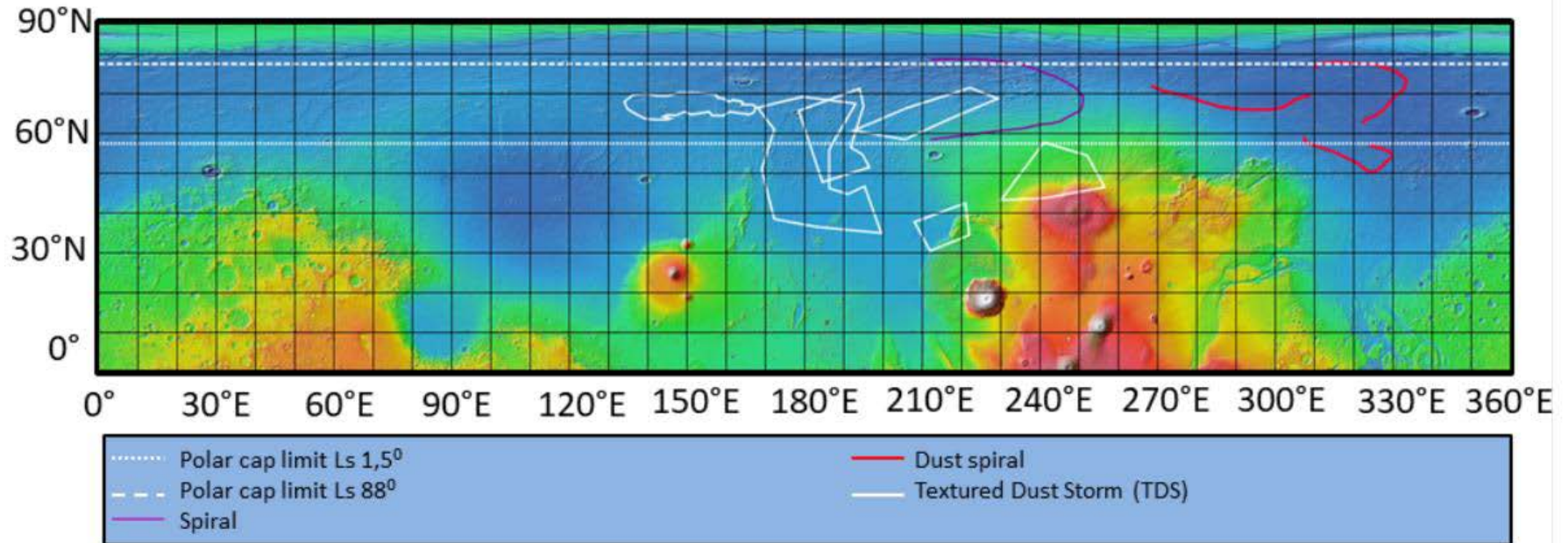


9 July

Textured Dust Storms (TDS)

May 22 – June 15 (2019)

$L_s = 31^\circ - 39^\circ$



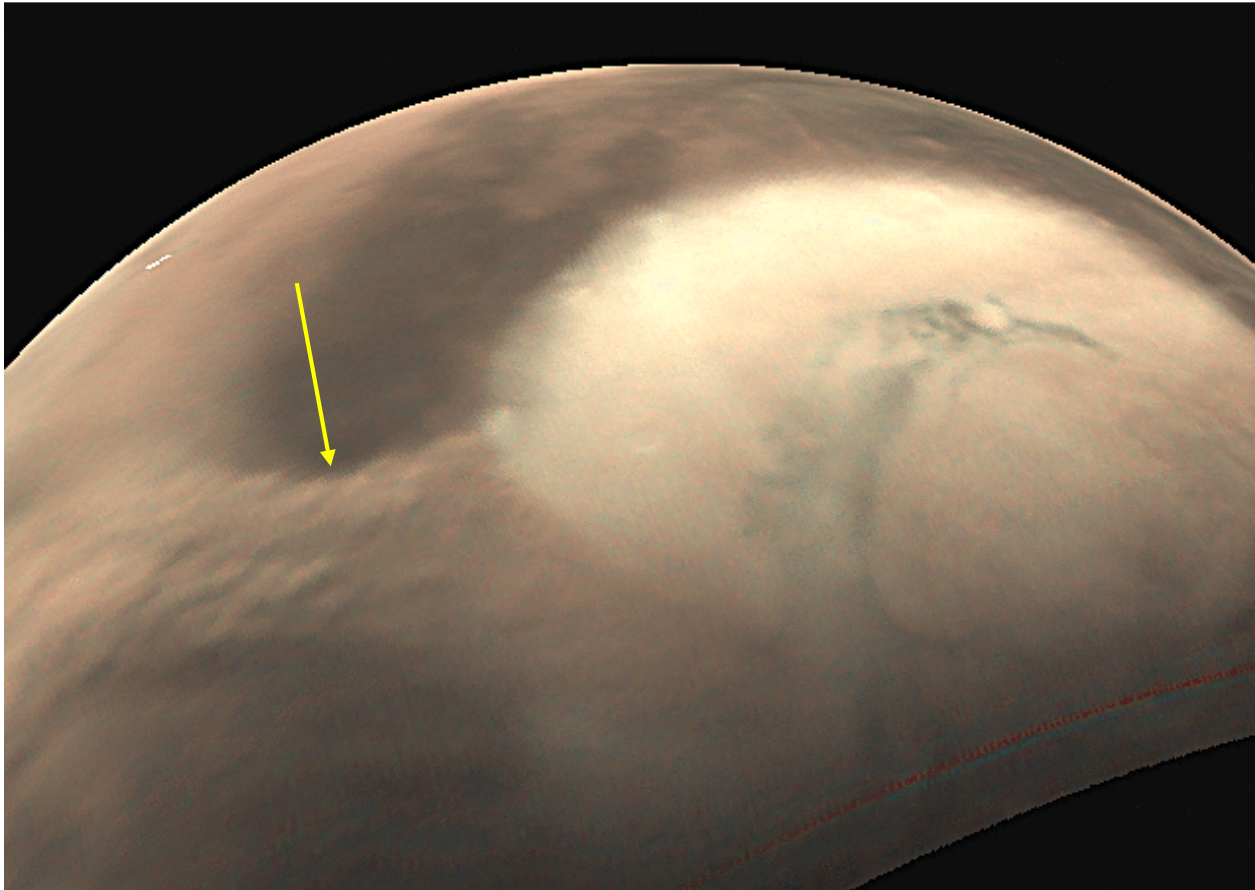
Local dust storms (*) showing texture at the top-surface (TDS)

Longitude range: $140^\circ\text{E} - 240^\circ\text{E}$

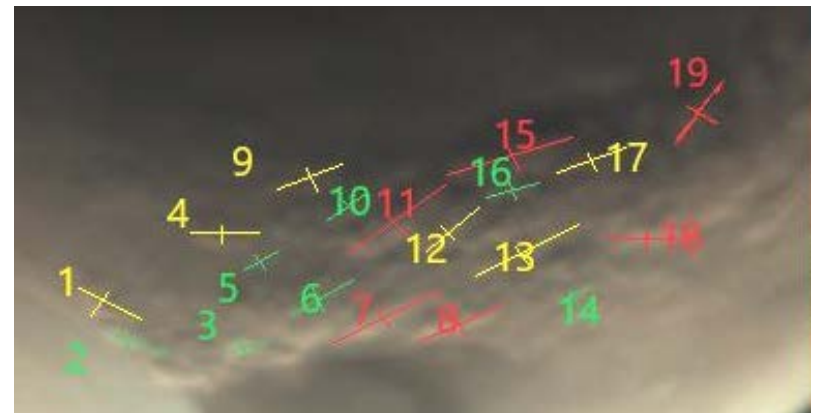
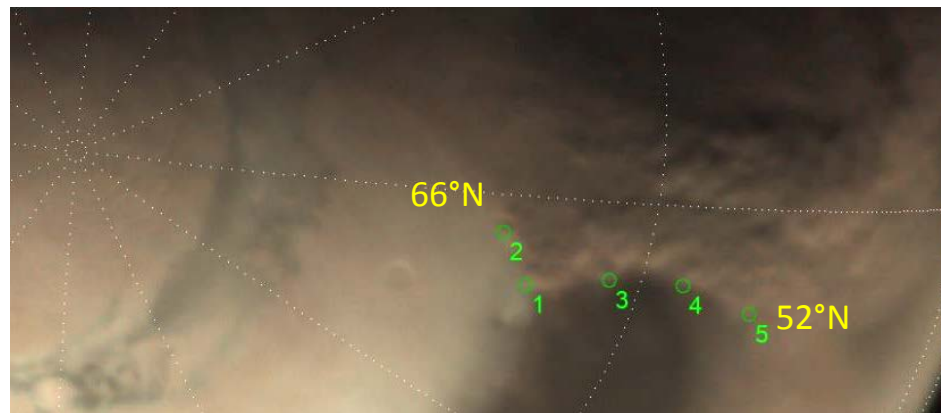
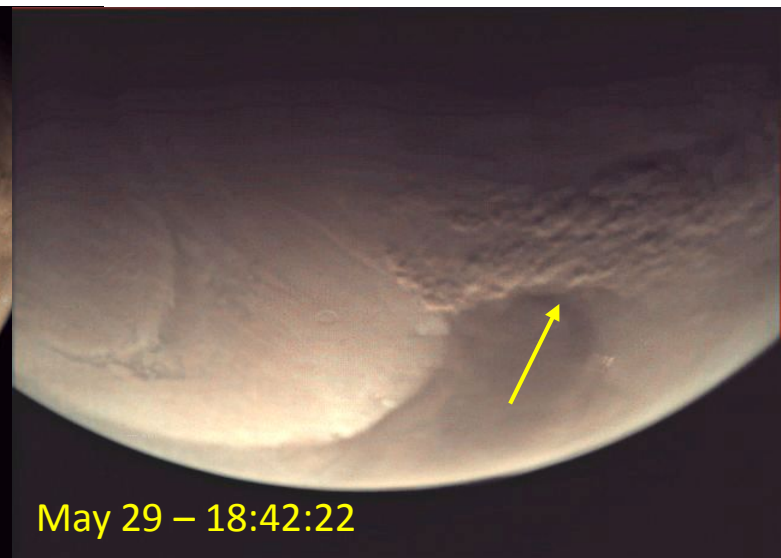
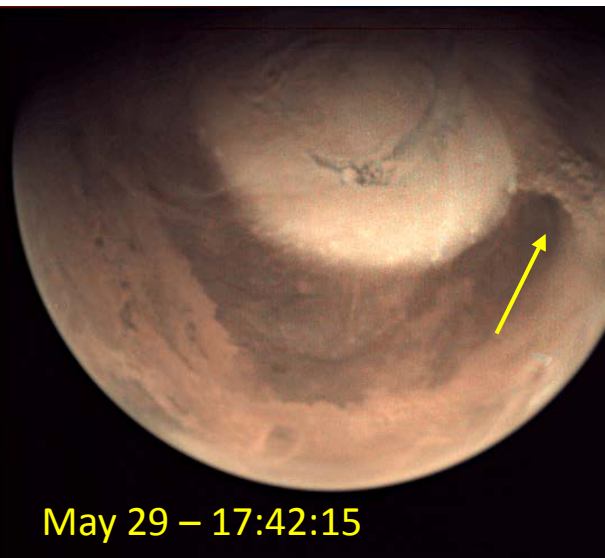
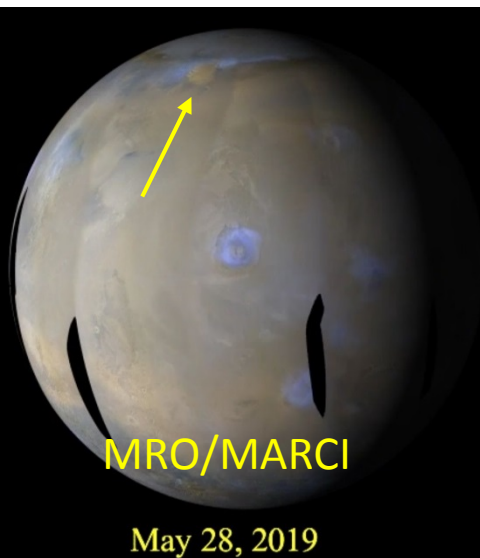
Acidalia, Arcadia and Amazonis Planitias

Local dust storm (*): area $< 1.6 \times 10^6 \text{ km}^2$

Local storm in Arcadia Planitia
Area = $1.36 \times 10^6 \text{ km}^2$
Ls = 32°



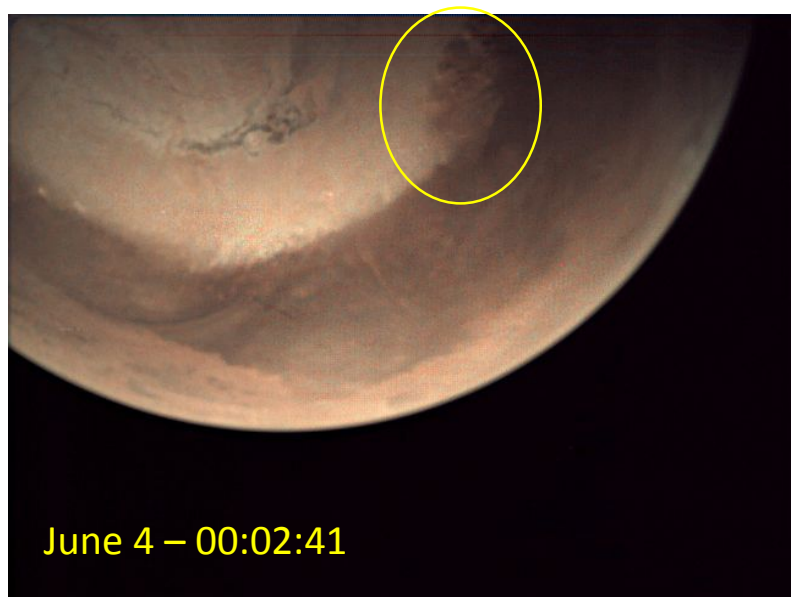
TDS1: 28 – 29 May 2019 (VMC)



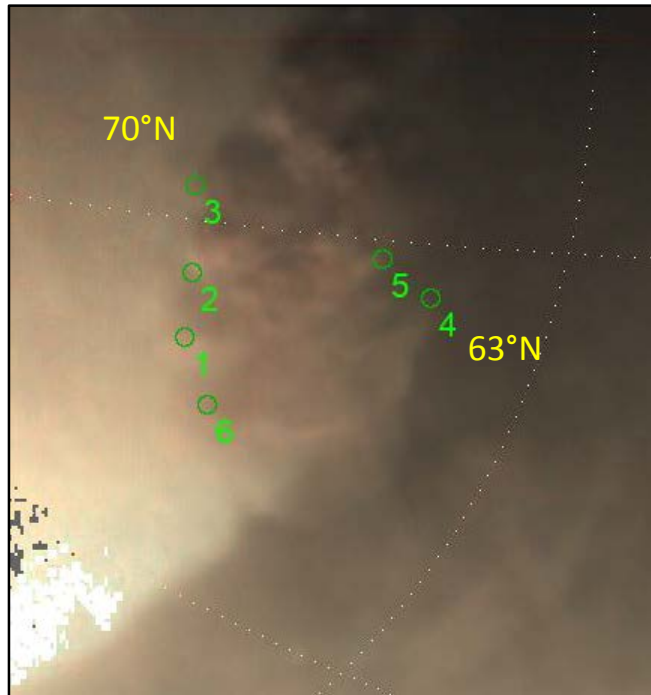
Texture type: pebbled-like
Cluster of cells ~ 20 - 40
Cells oriented by wind direction:
 $\langle V \rangle = 28 \text{ ms}^{-1}$ (20-40 ms^{-1})

Single cells:
Length = $140 \pm 45 \text{ km}$
Width = $61 \pm 15 \text{ km}$
length/width ~ 2.3
Separation = $180 \pm 60 \text{ km}$

TDS2: 3 – 10 June 2019 (VMC)



Local storm in Arcadia Planitia
Area = $1.6\text{-}2.1 \times 10^5 \text{ km}^2$
Ls = 34°

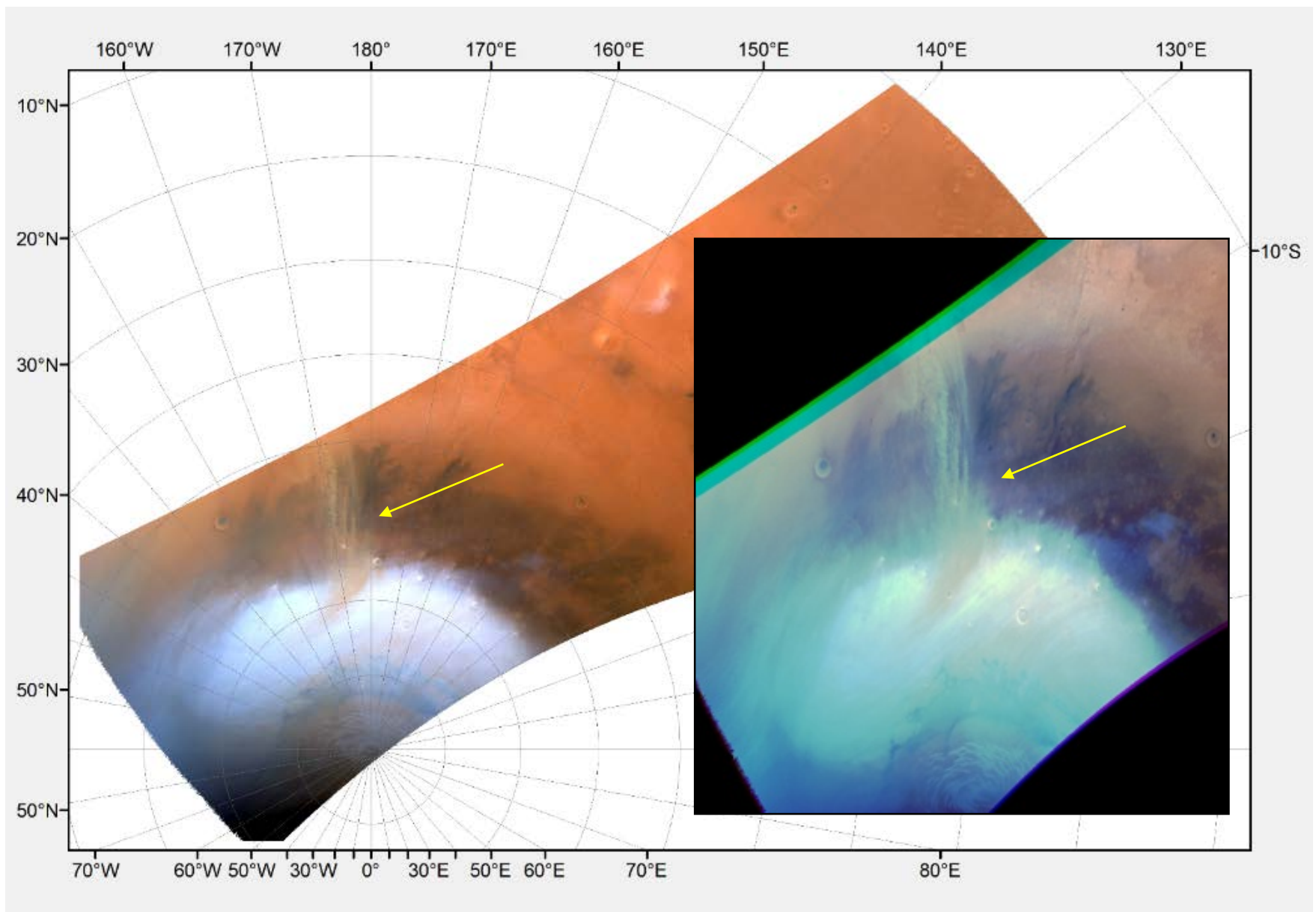


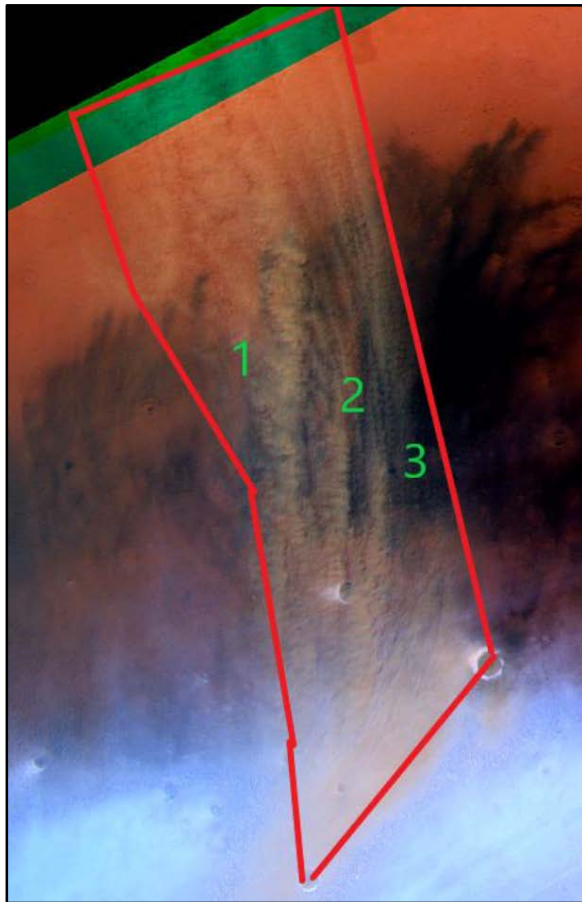
Texture type: pebbled-puffy
Cells oriented by wind direction:
 $\langle V \rangle = 38 \text{ ms}^{-1}$ (25-45 ms^{-1})



Single cells:
Length = $70 \pm 18 \text{ km}$
Width = $42 \pm 8 \text{ km}$
length/with ~ 1.6
Separation = $119 \pm 32 \text{ km}$

TDS3: 22 May 2019 (HRSC)





Local storm

Frontal shape

Three fronts:

Length = 950 km

Width = 85 km

Area = $4.3 \times 10^5 \text{ km}^2$

Ls = 28°



Single cells:

Length = 45 km

Width = 19 km

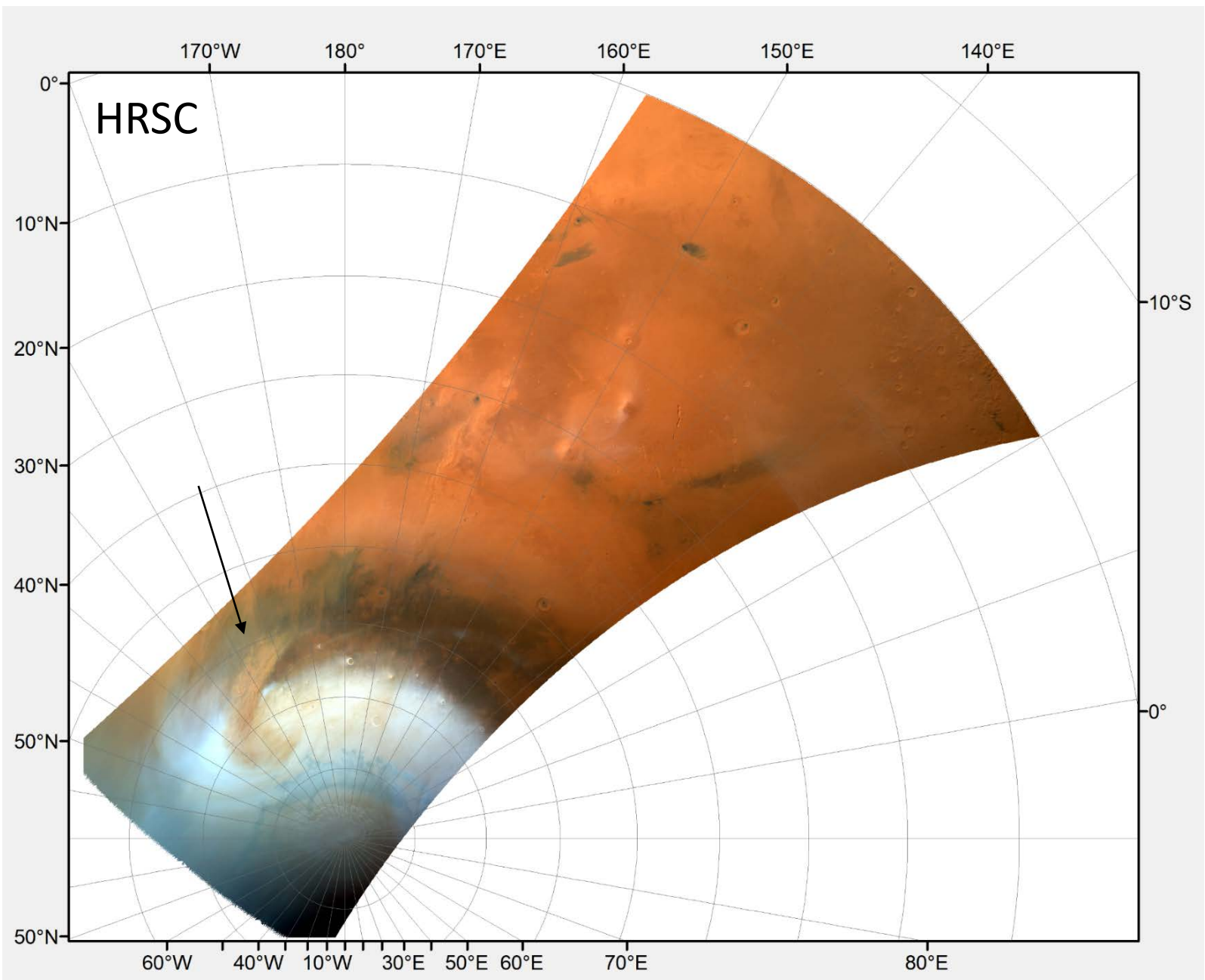
length/width $\sim 1.5 - 3.0$

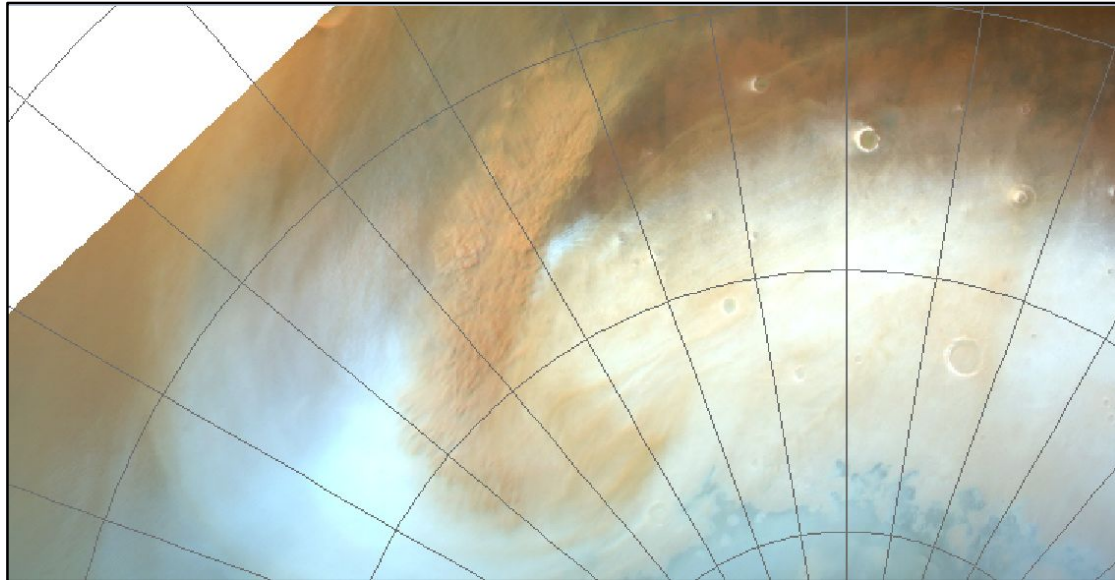
Separation = 35 - 40 km

Cells clustered in the fronts

Cells oriented transversal to front direction

TDS4: 26 May 2019 (HRSC)



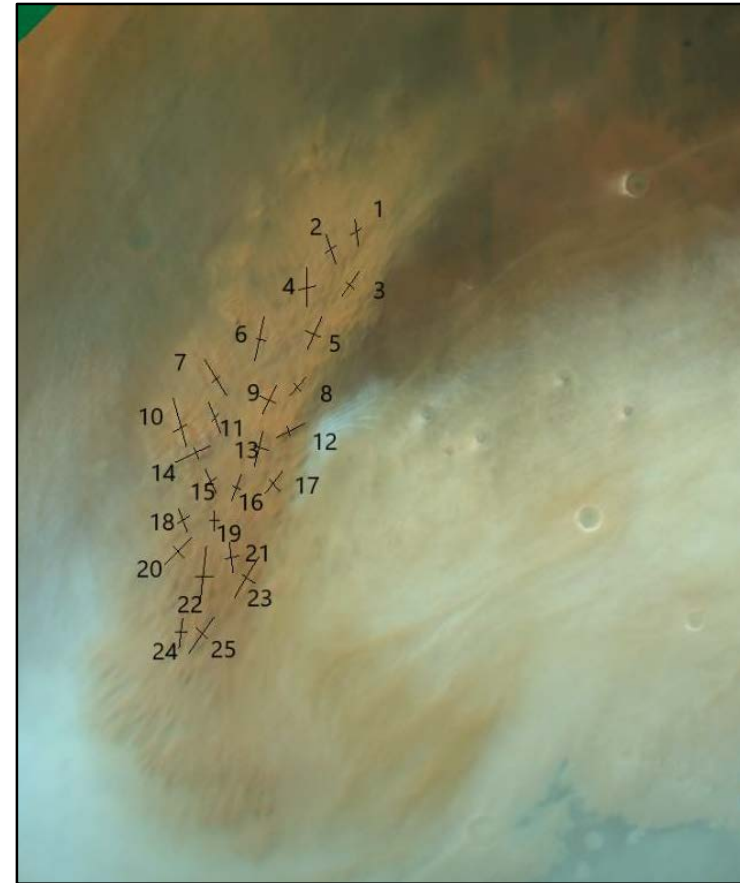


Local storm
Spiral shape
Area = $2.6 \times 10^5 \text{ km}^2$
Ls = 30°

Mixture of dust and water ice clouds

Cells oriented and concentrated mainly along the spiral arm

Spiral resulting from dynamical instability
(baroclinic nature?)



Single cells:
Length = $52 \pm 13 \text{ km}$
Width = $21 \pm 5 \text{ km}$
length/width ~ 2.5
Separation = $56 \pm 15 \text{ km}$

Conclusions

Four textured **local dust storms** North Pole edge in springtime ($L_s = 30^\circ - 40^\circ$)
Formed by clusters of 20-50 cells structured in oriented patterns
Single cell: length = 50 – 140 km, width = 20 – 60 km
Cells tops show anisotropy (length/width $\sim 1.5 - 3$)
Cells show pebbled and puffy-like textures (Kulowski et al., 2017)

Dust updrafts mechanisms:

- Mean horizontal winds ($V \sim 30 \text{ ms}^{-1}$) and high surface wind stresses
- Vertical wind shear due to dynamical instabilities (baroclinic)
- Dry convection driven dust heating (Spiga+, 2013; Heavens+, 2019)

Future work:

- Increase the number of cases studied
- Establish a relationship between wind speed and cell sizes
- Explore dust updraft models