

The ARCHES Space-Analogue Demonstration Mission: Towards Heterogeneous Teams of Autonomous Robots for Collaborative Scientific Sampling in Planetary Exploration

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Abstract—Teams of mobile robots will play a crucial role in future missions to explore the surfaces of extraterrestrial bodies. Setting up infrastructure and taking scientific samples are expensive tasks when operating in distant, challenging, and unknown environments. In contrast to current single-robot space missions, future heterogeneous robotic teams will increase efficiency via enhanced autonomy and parallelization, improve robustness via functional redundancy, as well as benefit from complementary capabilities of the individual robots. In this article, we present our heterogeneous robotic team, consisting of flying and driving robots that we plan to deploy on scientific sampling demonstration missions at a Moon-analogue site on Mt. Etna, Sicily, Italy in 2021 as part of the ARCHES project. We describe the robots' individual capabilities and their roles in two mission scenarios. We then present components and experiments on important tasks therein: automated task planning, high-level mission control, spectral rock analysis, radio-based localization, collaborative multi-robot 6D SLAM in Moon-analogue and Mars-like scenarios, and demonstrations of autonomous sample return.

Index Terms—Space Robotics and Automation; Multi-Robot Systems; Autonomous Agents

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Fig. 1: Our Lightweight Rover Units (LRUs) [5] and drone ARDEA [7] test their skills for autonomous navigation, exploration, and sampling at a Moon-analogue site on Mt. Etna.

I. INTRODUCTION

TEAMS of mobile robots that operate on the surfaces of extraterrestrial bodies will play a crucial role in future space exploration endeavors [1], [2]. Compared to current single-robot missions, they will improve efficiency via parallelization and robustness via functional redundancy. In addition, heterogeneous teams benefit from complementary capabilities of their individual agents [3], [4]. However, several key tasks in lunar and planetary exploration missions, such as infrastructure installation and scientific sampling, remain challenging: robots operating in Global Navigation Satellite System (GNSS)-denied environments have to rely on space-suitable on-board sensors to navigate, explore, and sample unknown areas in rough terrain. Long round-trip times, ranging from seconds (Moon) to many minutes (Mars), render direct teleoperation impractical to impossible. Thus, local robot autonomy is required as the basis for any coordinated joint action within the team. For this, the robots need to localize themselves online, map their surroundings, and share information about the environment and their position therein.

In the ARCHES project by the Helmholtz Association, we develop heterogeneous, autonomous, and interconnected robotic systems and technologies to tackle these challenges [8]. We will show their capabilities during demonstration missions in the summer of 2021 at a Moon-analogue site on the volcano Mt. Etna on Sicily, Italy.¹ While in ROBEX [9], a single robot deployed a seismic sensor array, we focus in ARCHES on the cooperation in a heterogeneous team. Our robotic team includes a drone for fast scouting and rovers with scientific instruments and manipulation capabilities for infrastructure installation, close-up analyses, and sample return. In Fig. 1, we give an impression of our preliminary experiments at the test site conducted in 2017 and 2019. In this article, we present

- our heterogeneous robotic team and its capabilities
- two mission scenarios: the setup of a distributed radio telescope as well as geological exploration and sampling
- experiments testing key components for the missions

It is in parts based on our IROS 2019 workshop paper [10]. In addition to extended descriptions of central aspects, we present novel contributions going beyond the workshop:

- mission scenario for the installation of a low-frequency radio array (LOFAR) and a radio-based localization of the array elements via our novel Radio-CPT system
- high-level mission control with automated task planning, multi-robot monitoring and coordination, autonomous task execution, and shared robot autonomy
- rock and sand analysis via laser-induced breakdown spectroscopy (LIBS) to be conducted by our rover
- collaborative simultaneous localization and mapping (SLAM) with three agents at the Moon-analogue site

We showed more than 35 live runs of collaborative SLAM and autonomous sample-return missions in a Mars-like scenario at the International Astronautical Congress (IAC) 2018, the world's largest annual gathering of space professionals with more than 6500 participants. Accompanying this article, we provide a video thereof as well as of our novel three-agent 6D SLAM experiment at a Moon-analogue site on Mt. Etna.

II. RELATED WORK

In this section, we discuss related work on multi-robot teams for planetary exploration. Focusing on scientific measurements, [11], [12] employ teleoperated rovers to carry instruments, limiting the use of vision sensors to enhance the situational awareness of the operators. LUNARES [3] demonstrated the ability of a semi-autonomous heterogeneous robotic team to explore lunar craters, artificially replicated indoors with a challenging slope and analogous lighting characteristics. In [13], a team of two hybrid wheeled/legged robots is remotely operated via an augmented reality interface to collect geological samples in a Mars-analogue environment. Assisted teleoperation is investigated in [14], where a swarm of GPS-based drones is deployed on Lanzarote to gather aerial images with the goal to minimize the cognitive load of the operators. In contrast to these approaches and to [4], our

robotic team will primarily operate in full autonomy governed by a complex mission control framework. The authors of [15] present a multi-robot coordination framework to map planetary caves. However, they employ rather simple wheeled/tracked robots evaluated in generic artificial environments. In contrast to the LIDAR-based vision in [3], [4], [13], [15], we use stereo cameras as a space-qualifiable sensor setup. Furthermore, our goal is to deploy a robot team at an outdoor Moon-analogue site to explore, deploy sensors, and perform scientific analyses.

III. HETEROGENEOUS ROBOTIC TEAM

Our heterogeneous robotic team consists of flying and driving robots as well as static and mobile infrastructure elements such as a lander and multi-purpose payload boxes [9], [17]. While drones on extraterrestrial bodies [2] require adaptations to their propulsion systems, e.g., realized by the Mars helicopter “Ingenuity” to be launched this summer [18], our drone allows us to implement and test transferable navigation and visual detection algorithms. We provide an overview of all systems in Table I. The robots benefit from their complementary capabilities: our drone ARDEA [6], [7] acts as a fast scout that can easily reach hard-to-access places such as craters or caves, whereas our first planetary exploration rover LRU1 [5] performs close-up inspections with its science camera and can transport ARDEA for energy efficiency reasons. Our second rover, LRU2 [5], [16] with its manipulator arm, takes soil and rock samples as well as transports, deploys, and manipulates payload boxes. The latter house scientific instruments and mobile infrastructure elements, e.g., to perform elemental analyses of rocks or to extend the robots’ communication range. The lander acts as a base station and defines an on-site coordinate frame as a global landmark for all robots.

IV. SCIENTIFIC EXPLORATION SCENARIOS

A. Low-Frequency Radio Array (LOFAR)

Setting up radio-astronomical instruments such as LOFARs on the far-side of the Moon allows probing the universe to unprecedented depth, as the Moon has no dense ionosphere and its far-side is one of the most radio-quiet areas in the inner solar system [19]. In this ARCHES scenario, we demonstrate key technologies to install, operate, and maintain a distributed

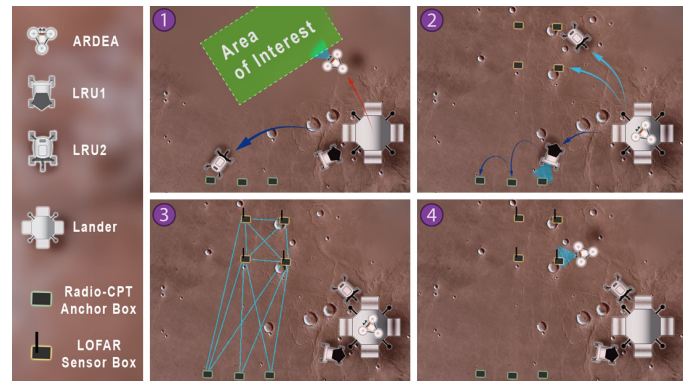


Fig. 2: ARCHES scenario with collaborative installation and maintenance of a low-frequency radio array (LOFAR).

¹The ARCHES demonstration mission was originally scheduled for 2020, but is now planned to be shifted to 2021 due to the COVID-19 pandemic.






System	System Features	Capabilities
ARDEA [6], [7] 	<ul style="list-style-type: none"> Flying system with high maneuverability Ultra wide-angle stereo (240° vertical, 80° horizontal FoV) Intel i7 dualcore at 3.0GHz (main computer), BeagleBone ARM (control), Spartan 6LX75 FPGA (dense stereo at 8 Hz) 	<ul style="list-style-type: none"> Fast autonomous scouting for exploration and mapping Identification of points of interest via semantic analyses Act as mobile radio communication relay Outdoor flight time of up to 10 min (104 Wh battery)
LRU1 [5], [9] 	<ul style="list-style-type: none"> Individually controlled & powered wheels on actuated bogies Pan/tilt camera head with stereo navigation cameras Scientific cameras: infrared camera, narrow-angle camera and stereo cameras with spectral filter wheels Take-off and landing platform for ARDEA Intel i7 quadcore at 2.7GHz (main computer), Intel Atom (control), Spartan 6LX75 FPGA (dense stereo at 11 Hz) 	<ul style="list-style-type: none"> Autonomous exploration and mapping in rough terrain Carrier for ARDEA Scientific analyses of rocks etc. with spectral and thermal cameras Autonomous operation time of up to 1.5 h in outdoor space-analogue environment (2 × 208 Wh batteries)
LRU2 [5], [16] 	<ul style="list-style-type: none"> Individually controlled & powered wheels on actuated bogies Pan/tilt camera head with stereo navigation cameras Manipulator arm with automatic docking interface Carriers for payload boxes (see below) and exchangeable tools (robotic hand to grasp stones & shovel to sample soil) Intel i7 quadcore at 2.7GHz (main computer), Intel Atom (control), Spartan 6LX75 FPGA (dense stereo at 15 Hz) 	<ul style="list-style-type: none"> Autonomous exploration and mapping in rough terrain Collection of geological samples (sand or small rocks) Deployment of scientific instruments (e. g., radio antennas or mobile spectrometers) Deployment of infrastructure (e. g., radio repeaters) Autonomous operation time of up to 1.5 h in outdoor space-analogue environment (2 × 208 Wh batteries)
Lander [9] 	<ul style="list-style-type: none"> Solar panels and batteries Computational power Communication link to mission control on Earth Storage for payload boxes with sample containers, scientific instruments, and infrastructure elements 	<ul style="list-style-type: none"> Perform computationally expensive tasks Save and aggregate data for downlink to Earth Gather solar power and recharge robots Store samples for future analysis Act as artificial landmark for navigation
Payload Boxes [9], [17] 	<p>Variants of payload box content:</p> <ul style="list-style-type: none"> Scientific instruments: <ul style="list-style-type: none"> Low frequency radio array (LOFAR) antenna elements Laser-induced breakdown spectroscopy (LIBS) Radio-based positioning system (Radio-CPT) & radio relay Soil or rock sample container <p>Each box has a common docking adapter to be grasped by LRU2.</p>	<p>Depending on content and placement of box:</p> <ul style="list-style-type: none"> Take scientific measurements Analyze elemental rock or soil composition Carry multiple soil or rock samples Localize payload boxes via Radio-CPT Extend communication range Provide power and communication to instruments Act as artificial landmark for navigation

TABLE I: Members of our heterogeneous robotic team with their complementary implemented and planned capabilities.

LOFAR for a carrier frequency between 20 MHz and 30 MHz in an area of about 100 m × 100 m. A LOFAR antenna element in the ARCHES analogue mission consists of a sensor payload box comprising a novel radio communication, positioning, and time-synchronization (Radio-CPT) system, a power supply and house-keeping unit, and a software-defined radio with a self-unfolding short dipole to receive radio signals from space.

In the LOFAR mission sketched out in Fig. 2, a team of scientists and operators select favorable array geometries for radio-astronomy observations. This selection process is assisted through a performance prediction software, taking into account resulting array directivity, wireless communication range, and Radio-CPT based self-localization accuracy. ARDEA then explores the area where sensor boxes shall be placed (①). In the meantime, LRU2 places Radio-CPT anchor boxes next to the lander, which are used for a precise localization of the antenna elements later on. A minimum of three anchors is needed to align the local LOFAR and global lander coordinate frames. Afterwards, LRU2 picks four LOFAR sensor boxes from the lander and places them in the afore defined array geometry, while LRU1 uses its cameras to measure in the Radio-CPT anchors locations w. r. t. the lander (②). At least three LOFAR boxes are required to determine the low frequency signal's direction of arrival (azimuth and elevation), while a larger number of boxes will allow an increased array aperture. For this, our system can be scaled up in a straightforward fashion. After placement, our Radio-CPT system automatically locates all boxes precisely: an

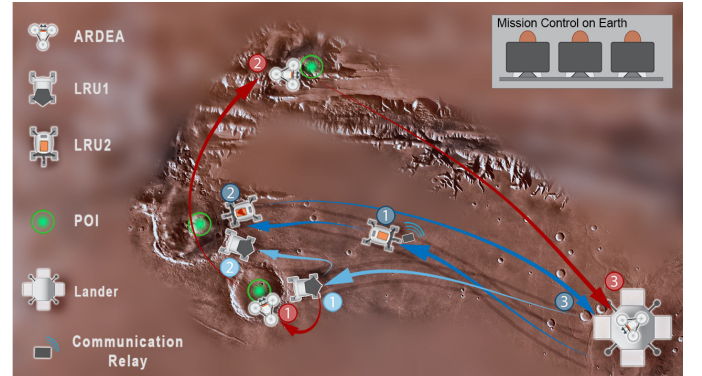


Fig. 3: ARCHES scenario with collaborative geological exploration and sampling (numbers indicate sequence of tasks).

important prerequisite to operate a distributed array. Short dipole antennas in each sensor box are released, the LOFAR is in-field calibrated, and the LOFAR signal reception starts (③). Pre-processed received signals are sent back to mission control for further processing and visualization by scientists, e. g., to identify S-type or L-type Jovian radio bursts [20]. In addition, scientists can select different modes of operation in post-processing, i. e., monitoring solar radio-bursts during daytime, estimating the angle of arrival of a strong source, or digitally steering the LOFAR beam in a specific direction. After disruptive events such as a dust storm on Mars, ARDEA can be sent to visually inspect the antenna array for damages (④).

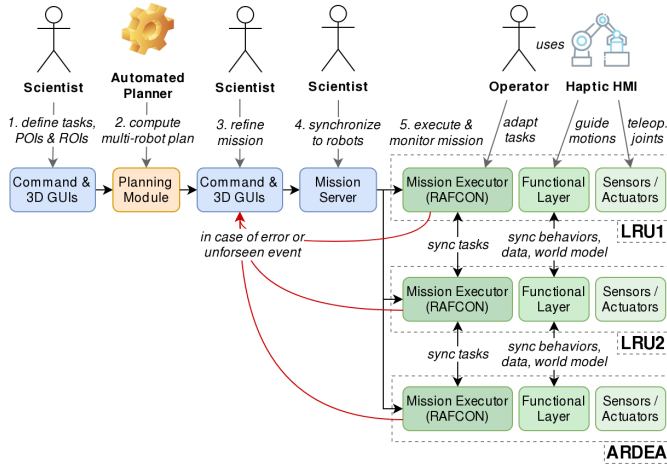


Fig. 4: Overview of our high-level mission control flow.

B. Geological Exploration and Sampling

The goal of the upcoming ARCHES geological sampling demonstration mission on Mt. Etna in 2021 is the creation of an accurate map of a target region, the spectral analysis of rocks, as well as the collection of sand and rock samples. For this, we will leverage the complementary capabilities of our heterogeneous team of (semi-)autonomous robots. In Fig. 3, we give an overview of the mission scenario. It involves the deployment of our robot systems and infrastructure elements listed in Table I. A landing spot had been selected by a team of scientists and mission operators based on prior knowledge from satellite imagery. It is located near three geologically interesting locations marked as points of interest (POIs) in Fig. 3. As a first step after landing, our rover LRU1 (1) is sent out to explore the closest POI in a nearby crater. It drives towards the target area while autonomously avoiding obstacles on its way. ARDEA, our drone, is carried by the rover on its transport and take-off platform. Upon reaching the target area, the 3D map created by LRU1 shows that the crater slope is too steep for the rover to enter. Therefore, mission control decides to launch the drone ARDEA (2) to further explore and map the crater from inside. At the same time, LRU2 (3) is sent out to place a payload box with a communication relay in between the lander and the two other, farther-away POIs to ensure radio coverage. In a second step, LRU1 (4) moves towards the next nearby, easier accessible, POI and uses its spectral cameras to identify stones for further inspection. Thus, mission control sends LRU2 (5) to deploy its LIBS instrument for laser spectroscopy and to collect the most promising rock sample, which it then returns to the lander (6). In the meantime, ARDEA (7) inspects and maps the final POI with its ultra-wide angle cameras to estimate the scientific gain of close-up inspection with the other robots. It then returns to the lander (8) for recharging.

V. HIGH-LEVEL MISSION CONTROL

For high-level mission control, we developed a framework for the efficient operation of multiple heterogeneous robots without requiring advanced robotics skills. The overall flow, see Fig. 4, consists of (1) task definition, (2) automated initial

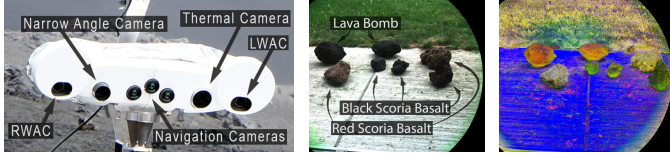


Fig. 5: Screenshot of the 3D GUI and status icons (per robot) of the ROSMC high-level mission control framework.

mission planning, (3) mission refinement, (4) synchronization to the robots, and (5) mission execution and monitoring.

To generate an initial mission plan, we use an automated planning algorithm that can handle cooperative tasks and precedence constraints relevant in both ARCHES missions. First, a greedy initialization similar to [21] provides a feasible initial solution. It is then refined via an optimization heuristic based on the idea of a relocation neighborhood known in vehicle routing [22], which we adapted to multi-robot task-allocation. For details and evaluations of the coordination algorithm, we refer to [23]. The tasks to be scheduled for the individual robots belong to predefined types. For ARCHES, the main ones include “take image” (ARDEA, LRU1, LRU2), “take science image” (LRU1), “take sample” (LRU2), and “take sample scientific mode” (LRU2 takes sample based on spectral images by LRU1). In addition, there are auxiliary task types for LRU2: “install radio range extender”, “pick payload box”, and “place payload box”. Each task is related to a POI for its execution. While main tasks are specified by the scientist, auxiliary tasks are added automatically whenever needed. Each auxiliary task adds a precedence constraint as, e.g., a sample container needs to be fetched prior to collecting a sample. The planner calculates an individual schedule for each robot, optimizing the overall mission time, driven distances, and energy consumption.

To enable intuitive task definition, mission refinement, and robust, distributed mission execution and monitoring, we introduce our novel ROSMC (ROS Mission Control) framework. It provides two graphical user interfaces (GUIs): a 3D GUI visualizing the global map, the agents’ poses, status, POIs, and ROIs (see Fig. 5) as well as a command GUI to define task sequences and parameterization. Mission control can adapt the pre-planned schedule of each robot: the 3D GUI allows to add tasks via a tree menu and to modify their POIs or ROIs, the command GUI to rearrange their order. The entire mission consists of a synchronized and a non-synchronized part, both being stored and tracked on a central mission server with low computational requirements. The former part, which is read-only in the GUIs, is sent to each robot so that they can execute the mission on board. The mission progress is stored locally and sent to the server when communication becomes available. In case of an error during runtime, the synchronized part can be deleted from the robot, adapted, and then re-synchronized. Powerful execution modalities based on hierarchical state machines modeled in our open source framework RAFCON [24] allow to pause and resume the



(a) ScienceCam sensor head with its seven cameras (b) RGB image and science product of volcanic rocks

Fig. 6: ScienceCam (a) and science product (b) visualized as $RGB = (Geo4, \frac{Geo1 - Blue}{Geo1 + Blue}, \frac{Geo6 - Geo1}{Geo6 + Geo1})$. Whereas the lava bomb and black scoria basalt look similar in the RGB image, they can be clearly distinguished in the science product.

behavior of each robot, to modify it, and to exchange the robots' entire command sequences during their mission. Task synchronization is achieved peer-to-peer without the server: robots communicate task ids with each other to coordinate inter-robot task dependencies. We are in preparation of a separate publication with details on ROSMC and intend to release it as open source in the future.

During mission execution, control can be switched to manual mode in case teleoperation is feasible and required. We employ a robotic arm as a generic haptic human-machine interface (HMI) to teleoperate different types of robots in a heterogeneous team. Its six degrees of freedom (DoF) allow for a direct mapping of the HMI states to the positions and orientations of the teleoperated robots, enabling an intuitive operation both in large-scale and precision tasks. Each DoF is actuated, which permits a continuous shift of the option-based levels of automation [25], covering the whole range from manual control via shared autonomy to full autonomy. Thus, operators can be supported with a cooperating automation to reduce their workload during teleoperation. As the haptic feedback is computed on the local HMI controller, the overall system is robust to communication delays.

These features render our mission control software more powerful than many approaches of related work, e.g., IMPACT [26], which does not allow a parameterization of reactive behavior, and M2PEM [27], which neither supports heterogeneous teams nor allows to switch to manual control.

VI. SPECTRAL ANALYSIS OF ROCK SAMPLES

The geological exploration of planetary surfaces aims at understanding the history of our solar system and supports the search for signs of life, resources, and habitable areas.

A. ScienceCam

LRU1 is equipped with a scientific camera, shown in Fig. 6 and introduced in [9] as ScienceCam. It has a concept similar to the PanCam of the ExoMars rover [28]. In addition to three navigation cameras, it houses wide-angle stereo (WAC, Manta G-145B NIR), a thermal (Xenics Serval-640), and a narrow-angle camera (Manta G-505C). The two WACs have spectral filter wheels, each containing three color filters (100 nm bandwidth) and six narrow band (10 nm bandwidth) geology filters ($Geo1 - Geo12$) across ranges of 440 – 660 nm and 720 – 1000 nm. By combining several bands, science products

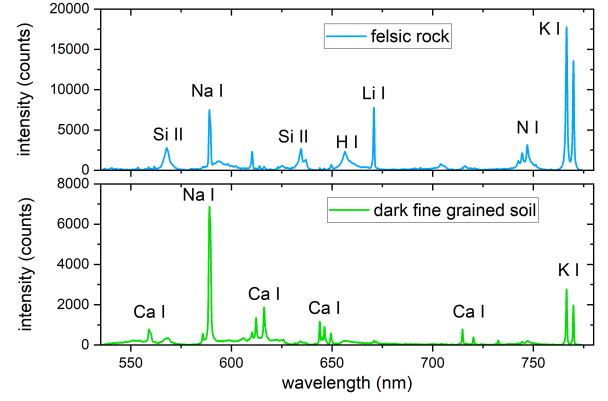


Fig. 7: LIBS spectra of a felsic rock and a dark fine-grained soil obtained with the ARCHES LIBS module for the LRU. The continuous emission was removed by a polynomial fit.

can be calculated to distinguish different types of rocks. In Fig. 6(b), we give an example based on a solely geometrically calibrated camera with the *Blue* (460 nm), *Geo1* (440 nm), *Geo4* (560 nm), and *Geo6* (660 nm) filters. For examination, we visualized the resulting channels as RGB: rocks that cannot be discriminated in the visible spectrum clearly stand out.

B. Laser-Induced Breakdown Spectroscopy (LIBS)

To perform elemental analysis of rocks and soil in its environment, LRU2 can equip itself with a laser-induced breakdown spectroscopy (LIBS) instrument that is housed in a payload box. LIBS can be used in close-up or remote configurations and gains more and more relevance for planetary in-situ exploration. Our LIBS module is connected to the rover's arm and built from commercial components (laser, spectrometer, micro-controller). It weighs about 1 kg and scans a small surface in close distance from the modular box with sampling rates of up to 30 Hz. A Nd:YAG laser with 8 mJ and 6 ns pulses at 1064 nm is used to create the micro plasma. The spectrometer covers a range of wavelengths of 550–770 nm, in which major rock-forming elements such as silicon, calcium, sodium, and potassium can be detected, see Fig. 7. Also minor and trace elements such as hydrogen and lithium can be observed, as well as the emission of nitrogen, which results from a breakdown of the ambient atmosphere.

VII. DISTRIBUTED COLLABORATIVE MULTI-ROBOT 6D SLAM IN MOON-ANALOGUE ENVIRONMENT

Multi-robot 6D Simultaneous Localization and Mapping (SLAM) allows robots to localize themselves w. r. t. each other and build a shared dense 3D model of their environment for autonomous exploration, mission planning, and the coordination of joint actions. We employ our online and on-board collaborative multi-robot SLAM framework [29], [30] to map rough-terrain environments in a distributed fashion: high-frequency and high-bandwidth sensor data is locally aggregated in real-time on each robot and then exchanged via so-called submaps between the systems. In contrast to other recent distributed SLAM approaches such as [31], [32], ours does not rely on a central server to delegate optimization and map management

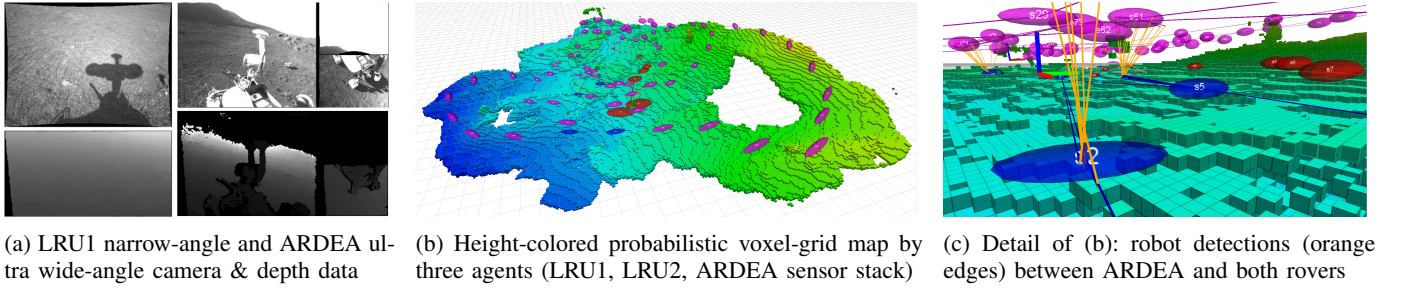


Fig. 8: Distributed multi-robot 6D SLAM in challenging conditions (harsh shadows, overexposures due to bright sunlight); data acquired in three-agent experiment at Moon-analogue site on Mt. Etna (10 cm map resolution, 1 m^2 background grid cells); ellipsoids show submap origins and positional uncertainty for LRU1 (blue), LRU2 (red), ARDEA sensor stack (pink).

<i>Multi-robot SLAM graph</i>	LRU1	LRU2	ARDEA
Number of robot poses	48	28	42
Number of submaps	7	10	78
Number of robot detections	0	17	42

TABLE II: Statistics for collaborative mapping exp. of Fig. 8.

tasks. Instead, every agent maintains a full SLAM graph for incremental online global optimization, exchanging constraints (submap matches and robot detections) and compacted submap data (3D voxel-grids and pointclouds) with no need for duplicate transmissions. This allows distributing the computational workload and keeping communication bandwidth requirements low. We refer to [29], [30] for further details on our SLAM framework, including more than 30 real-world experiments with five different robot platforms and evaluations regarding accuracy, runtime, and resource requirements.

Going beyond our previous two-agent SLAM experiments [29], [30], [10], we present first three-agent SLAM results on data acquired by a heterogeneous team of two LRU rovers and the sensor stack of our drone ARDEA at the Moon-analogue site [9]. In Fig. 8(a), we highlight challenges faced by the robots observing the world with their stereo cameras. In Fig. 8, we show the resulting dense multi-robot 3D map on a slope of Mt. Etna. Both rovers autonomously explored a given area by maximizing their expected information-gain [33]. In the meantime, we carried the sensor stack of ARDEA between them, simulating a flight that explores further areas, see the outer parts of the map in Fig. 8(b). Visual markers on both rovers allowed close-range detection by each other and by ARDEA in order to create inter-robot loop closures, shown as edges in the SLAM graph in Fig. 8(c). The robots thus acted as moving landmarks for each other, enabling a joint pose and map optimization. The resulting multi-robot graph has a total of only 213 nodes and 270 factors, see Table II, and thus allows for fast optimization.

VIII. RADIO-BASED LOCALIZATION (RADIO-CPT)

The LOFAR mission (Sec. IV-A) depends on precisely locating the LOFAR boxes, which will be achieved by our radio communication, positioning, and time-synchronization (Radio-CPT) system. The boxes are connected by bi-directional radio frequency (RF) links with a carrier frequency of $f_c = 5.8 \text{ GHz}$, a bandwidth of $B_c = 25 \text{ MHz}$, and a self-organized time division multiple access (SO-TDMA) protocol [34]. The localiza-

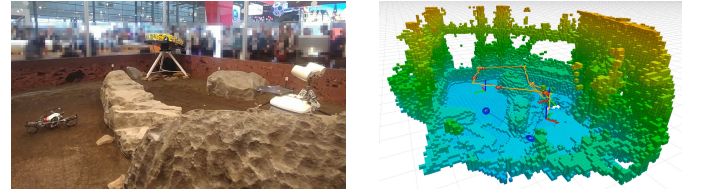


Fig. 9: Multi-robot SLAM with LRU1 and ARDEA at IAC 2018; all computation online and on board both robots.

tion accuracy depends on the clock accuracy, carrier frequency, transmit power, bandwidth, and the environment, which can promote multipath propagation [35]. In the ARCHES scenario, multipath effects are, however, not significant. We developed a distributed particle filter algorithm for radio localization [36], obtaining near-optimal performance with a complexity significantly lower than other state-of-the-art algorithms like [37]. In order to deal with the instability of low-cost oscillators, we implemented a distributed Kalman filter to track their characteristics using short TDMA frames (below 100 ms) [34].

We studied the joint estimation of LOFAR box location and LOFAR signal direction of arrival (DoA) in [38]. Our LOFAR setup targets signal reception at 20 MHz carrier frequency, i. e., a wavelength of 15 m. Accurate beam steering in azimuth and elevation requires an array localization accuracy of at least one order of magnitude higher than the carrier wavelength, which is 1.5 m for our system. We showed in simulation and multiple outdoor experiments [36], [34] that sub-meter accuracy can be achieved in an environment similar to the space-analogue mission, thereby verifying the feasibility of our approach. We plan to further reduce the impact of the LOFAR box location uncertainty on LOFAR DoA estimation via online phase calibration [39].

IX. INTEGRATION TESTS IN MARS-LIKE ENVIRONMENT

As a major milestone, we integrated several components for three different multi-system live demonstrations at IAC 2018 in a Mars-like sandbox of 50 m^2 , shown in Fig. 9. It features a lander mockup and artificial rocks as navigational and visual obstacles. All software required for autonomous robot behavior ran online and on board the three robots.

First, we demonstrated our distributed collaborative 6D SLAM with a heterogeneous team of our rover LRU1 and

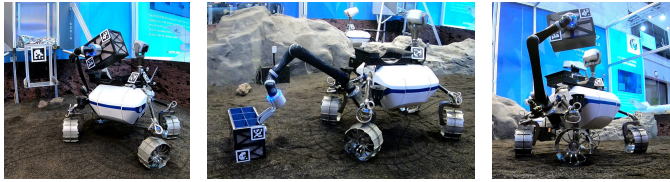


Fig. 10: Sequence of our IAC 2018 autonomous sample-return demonstration: (1) pick up container at lander, (2) extract sample, (3) reclaim container and return it to lander.

drone ARDEA in more than 35 successful live runs. On-board and online, both robots locally created partial 3D maps, exchanged, optimized, and merged them to a collaboratively built joint map of the environment. To connect the maps via loop closure constraints, ARDEA detected visual markers on LRU1 during its flight. We show the final 3D map in Fig. 9 and refer to [30], [7] for further details on this demonstration.

Second, in further 35 live runs, we demonstrated the robust integration of our on-board navigation, task execution, and manipulation required for a fully autonomous sample-return task, as shown in Fig. 10: LRU2 retrieved a sample container, transported it to a target location, placed it on the ground, grasped a shovel, scooped a soil sample into the payload box, and returned it to the lander. For details on the method and earlier experiments on Mt. Etna, we refer to [16].

Third, we integrated our Radio-CPT system in a robotic mission scenario and demonstrated real-time radio-localization of LRU2 carrying a Radio-CPT payload box in more than 20 live runs (> 30 min each). The high-rate real-time physical layer and MAC-layer processing ran in a distributed fashion in seven sensor boxes, and an off-board particle filter provided real-time 2D position estimates.

X. CONCLUSION AND FUTURE WORK

We summarize our methods to tackle the technical core challenges of the two ARCHES missions in Table III and give an evaluation regarding their maturity and open challenges. All of our robots are equipped with space-suitable sensor setups for visual-inertial navigation in GNSS-denied environments. A fully decentralized and in large parts distributed computation allows an efficient utilization of limited resources and ensures robustness to failures of individual agents. All modules run on-board and online, i. e., at a rate sufficient for reactive behaviors, and exchange locally pre-processed data for low-bandwidth communication. This allows full on-board autonomy, which is crucial in the light of unreliable and delayed communication links. We designed our system to scale for heterogeneous teams of $n < 10$ agents, thus swarms of hundreds of robots envisioned for the distant future will introduce further challenges, in particular w. r. t. data exchange. Our modular approach allows an easy adaption of core components to different robots and settings. Adding new types of sensors or robots, for example, does not require any changes to the other agents in a team as our multi-robot 6D SLAM connects them on the level of poses and aggregated (sub)map data. Furthermore, our framework RAFCON [24] (task execution) is available as open source and ROSMC (mission control) planned to be in the future.

For the ARCHES demonstration on Mt. Etna in 2021, we will bring together and extend the systems and capabilities shown at IAC 2018. As we experienced in the ROBEX campaigns [9], the additional challenges of full system integration, environment conditions (terrain, light, wind, local gravity, air density, etc. different to the lab), and logistics are not to be underestimated. Furthermore, we started to adapt individual modules for the Martian Moons eXploration mission (MMX) to navigate a rover on Phobos in 2026 [42].

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Technical Core Challenges	Components & Modules	Methods	Validation Status	Experiments & Demonstrations	Results & Open Challenges
High-Level Mission Control	ROSMC (Sec. V)	GUIs (in RViz + RQt) & data server	Simulation	User study: 9 people, three-robot simulations	Concepts for sub-groups and $n > 3$ robots; will require UI adaptations
Automated Task Planning	Multi-robot scheduling (Sec. V)	Greedy init. & local improvement heuristic [23]	Simulation	Evaluated on 600 instances (up to 3 robots, 15 tasks with 9 coop. tasks) [23]	Reliable generation of feasible optimized plans; future work: integration with mission control
Autonomous Task Execution	RAFCON (Sec. V, [24])	Hierarchical statemachines with concurrent execution	Space-like & analogue environments	Sim. & lab, SpaceBot-Camp 2015 [5], ROBEX 2017 [16], IAC 2018	Statemachines with 8 hierarchy levels and > 700 states per robot; future work: multi-robot world model
Shared Autonomy in Teleoperation	Haptic 6DoF HMI (Sec. V)	Haptic shared control [25]	Simulation & laboratory	Teleop. of LRU camera head, simulated drone, and 3DoF robotic arm	Results: 6DoF HMI with generic haptics & supporting automations; pending: user study on haptic HMI
Fast Scouting of Points of Interest	ARDEA wide-angle vision system (Sec. IX)	On-board robust dense stereo vision [6], [7]	Space-like & analogue environments	Laboratory [6], IAC 2018 [7], Autonomous waypoint flight on Mt. Etna 2019	Successful takeoff from rover; landing on LRU challenging due to its camera mast (work in progress)
Dense 3D Multi-Robot Mapping	6D Multi-Robot SLAM Framework (Sec. VII & IX)	Dense submaps, real-time filter + inc. graph optimization [29], [30]	Space-like & analogue environments	Sim. & lab., SpaceBot-Camp 2015, ROBEX 2017, IAC 2018 ($n > 35$ exp.) [5], [29], [30], [7]	Evaluated with two (on-board) & three (on datasets) agents; Open challenges: long-range mapping & multi-session relocalization [40]
Waypoint Navigation & Obstacle Avoidance	Obstacle mapping & 2D path planning (Sec. VII & IX)	Stereo-adaptive terrain classification [30] & adapted ROS nav. stack	Space-like & analogue environments	Sim. & lab, SpaceBotCamp 2015 (LRU), ROBEX 2016 & 2017 (LRUs), IAC 2018 [5], [30], [7]	Validated on LRU rovers; ARDEA: waypoint navigation & terrain following [41]; future work: full 3D path planning for ARDEA
Autonomous Exploration	Frontier exploration (Sec. VII)	Information gain-based expl. & active loop closing [33]	Space-like & analogue environments	Sim. (40 exp.), lab, outdoor & ROBEX 2017 ($n > 30$ real exp.) [33], [5], [30]	Tested single- and multi-robot [30] (rovers); future work: full 3D expl. for drone & multi-criteria decisions
Payload Transport & Manipulation	Compliant control, motion planning, world model (Sec. IX)	Impedance control, sampling-based planning, synchronized graph database [16]	Space-like & analogue environments	Simulation & laboratory, SpaceBot-Camp 2015 [5], ROBEX 2017 [16], IAC 2018 ($n > 35$ exp.)	Autonomous deployment & pickup of payload; challenges: high uncertainties due to object detection and elasticity in the rover; future work: compliant whole-body control
LOFAR Measurement	LOFAR Payload Boxes (Sec. IV-A)	Distributed RF-sampling & beam steering	Laboratory setup	Prelim. experiments: single-box RF reception on rooftop	Current & future work: integration in payload boxes, proof of concept experiments with multiple boxes
LOFAR Box Localization	Radio-CPT (Sec. VIII & IX)	Distributed particle filter [36]	Space-like & analogue environments	Grass field 2015 [36], gravel pit 2018 [34], IAC 2018, Mt. Etna 2019	Achieved sub-meter accuracy in most experiments; challenges: multipath propagation & instable clocks
Panoramic & Multi-Spectral Imaging	ScienceCam (Sec. VI-A)	Anomaly detection via combination of spectral images	Lab, space-analogue environment	Prelim. tests with volcanic rocks [10]; panorama at space-analogue site	Radiometric & spectral calibration difficult \rightarrow differentiation but no classification of minerals possible
Elemental Rock & Soil Analysis	LIBS (Sec. VI-B)	Optical emission spectroscopy	Lab tests of instrument	Eight reference spectra of geological samples	Validated on rocks & sand; integration with LRU payload box pending

TABLE III: Summary of our approaches to the core challenges of the ARCHES missions and evaluation of their maturity.

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